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# TREBALL DE FI DE GRAU

**TFG TITLE:** Determination of drag coefficients in a free molecular flow

**DEGREE:** Grau en Enginyeria d'Aeronavegació

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**DATE:** July 17, 2018



**Títol:** Determinació del coeficient d'arrossegament en un flux molecular lliure

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**Data:** 17 de juliol de 2018

## Resum

Aquest projecte té com a objectiu estudiar el coeficient d'arrossegament ( $C_D$ ) dels satèl·lits en un flux molecular lliure. L'estudi està basat principalment en els models de Schaaf & Chambre i de Schamberg. Són models antics (de 1958), però encara són les grans referències utilitzades en l'actualitat. S'analitzaran els dos models per geometries planes i esfèriques, tenint en compte reflexió especular i difusa.

Per presentar l'escenari del nostre estudi es fa una descripció de la termosfera i de l'interès en l'estudi d'aquesta regió de l'atmosfera. Una vegada està definit l'escenari, es descriu el tipus de flux amb el que es tracta en aquest estudi. S'expliquen els paràmetres que divideixen els gasos en diferents règims, així com les característiques bàsiques del flux molecular lliure i la interacció entre aquest i la superfície del satèl·lit.

El primer model d'estudi és el de Schaaf i Chambre, amb el qual s'han estudiat els dos tipus de reflexió (especular i difús) per a cada geometria. Els nostres resultats per la placa plana-paral·lela mostren que el paràmetre més influent és l'angle d'atac. En el cas de l'esfera, la simetria fa que aquest paràmetre sigui irrellevant i, a més, la temperatura de la superfície del satèl·lit tampoc no provoca variacions significatives.

Amb el model de Schamberg també es poden veure grans variacions de  $C_D$  en funció de l'angle d'atac en el cas pla-paral·lel. El cas de l'esfera dona també  $C_D$  constant pel cas especular. També hi ha una clara dependència amb la temperatura de la superfície als casos difús i entremig.

Els valors del  $C_D$  pel cas de reflexió difusa difereixen molt entre models. De fet, el model de Schamberg retorna uns valors que no concorden amb els valors empírics de satèl·lits reals en certes condicions. Per aquest motiu s'ha decidit finalment estudiar un tercer model, el model de Sentman per reflexió difusa. Aquest model dona resultats molt semblants al model de Schaaf i Chambre. Addicionalment, els resultats de la bibliografia apunten a les greus limitacions del model de Schamberg per estudiar el cas difús i angles d'atac petits. Hem pogut comprovar la necessitat pràctica d'utilitzar combinacions de models, en funció de les condicions de la reflexió.

Hem pogut concloure que un primer estudi de  $C_D$  (i, eventualment, de la densitat a la termosfera), hauria de fer ús de satèl·lits amb simetria esfèrica, per tal de minimitzar el nombre de variables d'entrada i acotar millor els paràmetres influents ineludibles. La problemàtica tècnica del llançament de satèl·lits esfèrics es podria resoldre situant aquests satèl·lits esfèrics a dintre d'estructures de geometria CubeSat durant el llançament, i permetent que s'obriessin quan estiguessin en òrbita, o fent servir dispensadors d'uns quants satèl·lits esfèrics integrats en CubeSats.



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## Overview

This project aims to study the drag coefficient ( $C_D$ ) for satellites in a free molecular flow. The study is mainly based on Schaaf & Chambre and Schamberg models. These are old models (from 1958), but still are the most widely used nowadays. The two models are analyzed for flat and spherical geometries, considering both specular and diffuse reflection. In order to set the scenario of our study, we describe the thermosphere and justify the interest of studying this region of the terrestrial atmosphere. Then, we describe the specific type of flux considered in this work. The parameters which divide the behaviour of the gases into different regimes, as well as the basic characteristics of the free molecular flow and the interaction between it and the surface of the satellite, are explained.

The first model studied is the one by Schaaf and Chambre, considering the two types of reflection for each geometry. We observe that, in the case of flat plate geometry, the angle of attack is the most important parameter for the determination of  $C_D$ . For the spherical case, we see that  $C_D$  adopts almost constant values, since the angle of attack is irrelevant due to the geometry, and the temperature does not cause great variations.

With respect to Schamberg model we obtain large variations of  $C_D$  due to variations in the angle of attack, as in the previous model. The case of the sphere also gives a constant  $C_D$  in the specular case. There is also a clear dependency on the wall temperature for the diffuse and intermediate cases.

The values obtained for  $C_D$  under the diffuse reflection show large discrepancies between models. In fact, Schamberg's model gives values that are at odds with empirical values under certain conditions. For this reason, we also analyze a third model: Sentman's model for diffuse reflection. This model gives results similar to those of Schaaf & Chambre. Additionally, the bibliography points to serious shortcomings of Schamberg's model in the diffuse reflection case and low angles of attack. We have checked the practical necessity of combining these models for different reflection conditions.

We have concluded that a first study of  $C_D$  (and, eventually, of the thermospheric density) should make use of spherical satellites in order to minimize the number of relevant variables, and to better identify the most influential parameters. The problem of launching spherical satellites could be solved by holding them inside a suitable satellite dispenser, perhaps with a CubeSat-like configuration, and release them once in orbit. These satellites dispensers could release either one or several spherical satellites.



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# **ACKNOWLEDGEMENTS**

I would like to express my sincere gratitude to Jordi Gutiérrez and Pilar Gil, for giving me the opportunity to develop this project with them and supporting me during all these months in those moments when the project seemed to break into pieces due to the difficulty in finding information and its accessibility.

Moreover, I would like to thank my family, specially my mother Pilar, my father Silvestre and my sister Marta, for their support and advice during the progress of this project.

Finally, I would like to thank my friends for being by my side all this time and never stop trying to help me in everything I have needed.





# INTRODUCTION

Is space overcrowded? In the last decades, the space competition has reached its limit, as it is reasonably easy and "Low-Cost" to send an object to the space. For that reason, space's population has increased a lot recently and therefore, space is overcrowded. That's why it is important to know and determine correctly the orbital trajectories of satellites in order to avoid collisions.

A correct assessment of spacecraft trajectories near the Earth requires a proper knowledge of the interactions between the Earth's atmosphere and the spacecrafts themselves. In spite of the great effort done in the last century, these interactions and, ultimately, their effects on the dynamics of spacecrafts is not well understood.

## State-of-the-art

The atmosphere of the Earth extends to about a 1000 km in height, becoming less and less dense with increasing altitude, but not showing a physical clear-cut limit.

Legally, airspace is limited to a height of 100 km (the von Kármán line), but its effects on satellites are relevant up to altitudes in excess of 400–500 km. Its most important influence is the decrease in semi-major axis (orbital spiraling-in), leading to the re-entry of satellites in the atmosphere.

The atmospheric density varies in a very wide range, between values of the order of 1 kg/m<sup>3</sup> near the Earth's surface, and down to values of the order of 10<sup>-15</sup> kg/m<sup>3</sup> at the uppermost layers of the atmosphere. As a consequence, a complete understanding of physics of the atmosphere involves a proper knowledge of fluid dynamics in very different regimes, ranging between the continuum (near the Earth's surface), and the rarefied fluid dynamics (at higher layers). Besides, it is important to realize, that atmospheric density values are not constant with time. In particular, density above a height  $\sim 100$  km strongly depends on the specific solar activity and geomagnetic conditions at a given time.

There are several mathematical models, such as NRMLISE-00 developed by the US Naval Research Laboratory (8), or JB2006, by Jacchia and Bowman (9) which forecast the air's density at altitudes between 100 km and 1000 km, and which are used, for example, to predict the timing of satellite re-entries. However all their outputs are –to some point– uncertain, even when considering the varying effects of solar activity (through the so-called  $F_{10.7}$  proxy), or the geomagnetic state (by means of the  $A_p$  parameter).

This poor knowledge of atmospheric density values is specially true in regions of low Earth orbit (hereafter referred to as LEO)<sup>1</sup> which are not densely populated by satellites. One of such regions, the lower thermosphere, is located between 100 and 300km, and is particularly important, because it dominates the last stages of satellite re-entries.

Most satellites found in the thermosphere are actually in their re-entry stage and lack a proper attitude control. This fact sets strong limitations on our knowledge of re-entry dynamics and, in particular, on the dynamics of drag. This problem is not solved by the existence of satellites orbiting at extremely low altitudes, like some of the Earth's gravitational field explorers (e.g. CHAMP, or GRACE), as well as some optical/IR intelligence

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<sup>1</sup>LEO orbits typically correspond to heights between 200 m and 2000 m above the Earth's surface.

satellites. The reason is that these spacecrafts do not provide information on the drag they experience. As a consequence, one of the most important aspects of spacecraft dynamics, that is, drag dynamics, remains poorly constrained.

## Scope

In this work we intend to improve the current knowledge of drag dynamics. This knowledge is expected to be used in the analysis of a specific space mission. The payload of this mission would be a swarm of very small, spherical satellites in the femtosatellite category (masses between 10 and 100 g).

These femtosatellites will carry a high-sensitivity, microelectromechanical system (MEMS) accelerometer, a MEMS Global Navigation Satellite System (MEMS GNSS) receiver, a transmitter, and a primary battery (as well as some other electronic devices required for the functioning of the satellite). The accelerometers will be able to measure accelerations as low as  $1 \mu\text{g}$  caused by the drag with the residual atmosphere. This drag can be modelled by means of expressions whose validity is extended to the case found in space from the continuous case (see [1.1](#)).

## Methodology and structure of the project

This project will be based on a thorough research of articles, books and thesis about the study of the thermosphere and drag modelling, and, once understood the basic foundations, we will develop a Matlab code for each drag model to obtain several plots regarding different cases of study. Each model will have two Matlab codes; a main menu, in which the case of study will be chosen, and a function in which all the calculations will be done. All Matlab codes developed for this project can be found in [Appendix C](#).

This work is structured as follows: in [chapter 1](#), an introduction to the theoretical foundations for the study of the thermosphere is given in order to set the basis of the following drag characterization. [chapter 2](#) develops the Schaaf and Chambre model for drag coefficients, where several cases are presented either for flat plate or sphere surfaces. [chapter 3](#) develops the Schamberg model for drag coefficients, which adds more parameters to vary depending on the characteristics of the environment and several cases are presented as in the previous chapter. [chapter 4](#) develops Sentman model for diffuse reflection, which gives more accurate values either for flat-plate or sphere when diffuse reflection occurs. Finally, [chapter 5](#) draws our conclusions and possible future developments.

# CHAPTER 1. THEORETICAL FOUNDATIONS FOR THE STUDY OF THE THERMOSPHERE

Aerodynamic drag is one of the most significant orbital perturbations in Low Earth Orbit (LEO) as it reduces the orbital energy by interchanging momentum between the upper atmosphere and the spacecraft, resulting in a change of the eccentricity towards a more circular orbit and a reduction of the semi-major axis until the eventual re-entry of the spacecraft. That is the reason why modelling this force and its variability is a key issue to determine and predict spacecraft trajectories as well as measure atmospheric density via satellite drag. A common approach to calculate the drag acceleration experienced by a body is:

$$\vec{a} = \frac{1}{2} \rho v_{\text{rel}}^2 \frac{S C_D}{m} \hat{v}_{\text{rel}} \quad (1.1)$$

Where  $\rho$  represents the atmospheric neutral density,  $v_{\text{rel}}$  is the relative velocity of the vehicle with respect to the atmosphere,  $C_D$  is the drag coefficient,  $S$  is the reference surface area and  $m$  the mass of the body.

Despite the huge theoretical and experimental efforts done in the last decades (REFERENCES), three of the above parameters:  $\rho$ ,  $v_{\text{rel}}$  and  $C_D$  remain highly uncertain.

- **Neutral density ( $\rho$ )** is the main source of error, as the atmosphere is a moving mass of air and its density varies with time, altitude and location. The solar extreme ultra-violet radiation is absorbed by the atmospheric constituents and hence causes the corresponding density variations. Geomagnetic fluctuations caused by solar storms, which occur more frequently during periods of peak solar activity, also increase thermospheric density.
- **Relative velocity  $v_{\text{rel}}$ :** when calculating the relative velocity of the spacecraft with respect to the atmosphere it is a common to assume that the atmosphere rotates integrally with the Earth. However, the existence of winds may cause a significant departure from the former hypothesis. Winds can increase or decrease the drag coefficient depending on their blowing direction. This fact is an important source of uncertainty as wind velocities can be as high as several hundred m/s, and the reliability of their predictions is very limited.
- **Drag coefficient  $C_D$ :** the traditional perspective assumes a constant  $C_D$ , because precise modeling of this parameter is lacking. However, it is well known that the drag coefficient is not constant and, in fact, it can vary in a wide range of values, mostly depending on the body shape and the atmospheric temperature, as well as on chemical composition of the surfaces and the remaining atmosphere.

In spite of all the difficulties which involve the computation of satellite drag, its importance has maintained and even increased the interest in spacecraft drag modelling in recent times. An improvement on the knowledge of this topic will significantly increase the accuracy of orbit determination and space objects tracking.

## 1.1. The Thermosphere

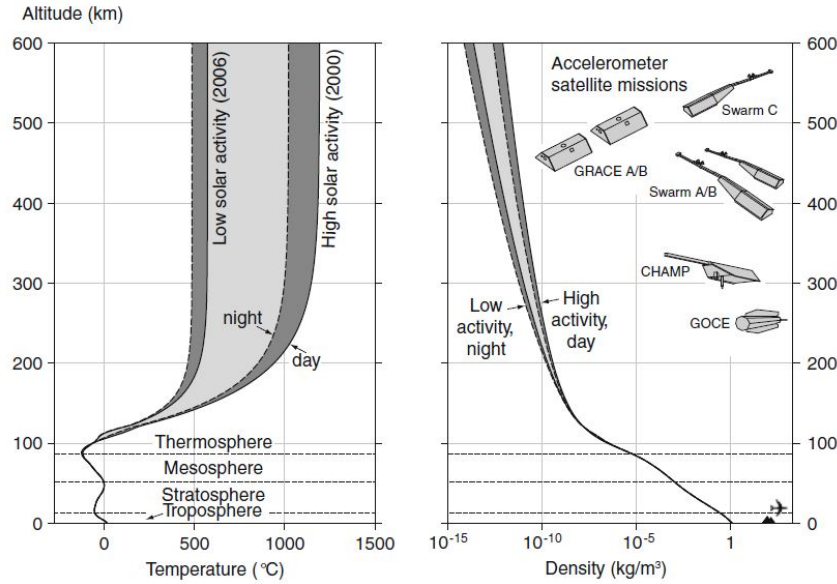


Figure 1.1: Altitude profiles of atmospheric temperature (*left*) and density (*right*) according to NRLMSISE-00 model.

The standard classification of the different parts of the Earth's atmosphere is shown in the left panel of Figure 1.1. The neutral atmospheric region of interest for satellite aerodynamics is the thermosphere, which is an atmospheric layer above an altitude of 85 km and about 500–600 km. This layer absorbs the extreme ultraviolet (EUV) energy from the Sun so the temperature profile of the thermosphere increases sharply with altitude at the lower part, while remaining almost constant with height at the upper part. This limit temperature is called exospheric temperature, and it may change with time, with values ranging, approximately, between 500°C and 2000°C.

Neutral density in the thermosphere is one of the most important variables to model for applications in solar-terrestrial physics and drag computations for satellite orbit determination. Both temperature and neutral density vary under the influence of complex interactions between the Earth system and solar processes, specifically on the amount of energy received by the thermosphere, mostly from solar flux and geomagnetic activity.

### 1.1.1. Why study the thermosphere?

Thermosphere density models are applied in scientific investigations and in many types of satellite orbit computations. The accuracy of density models influences scientific results but it also affects requirements for space mission operations, tracking systems and propellant consumption. Some of the main reasons to study the thermosphere are summarized as follows.

#### 1.1.1.1. Applications in Space Mission Analysis and Operations

The atmospheric drag force causes all LEO objects to spiral downward, so they may eventually re-enter the densest atmospheric layers. Since density and orbital velocity decrease with altitude, the drag force is stronger at the perigee of the orbit, and at this point the transformation of kinetic energy into heat is more efficient. The decrease in mechanical energy naturally leads to a circularization of the orbit. Figure 1.2 shows the change in shape of the PAM-D object with catalogue number 22659.

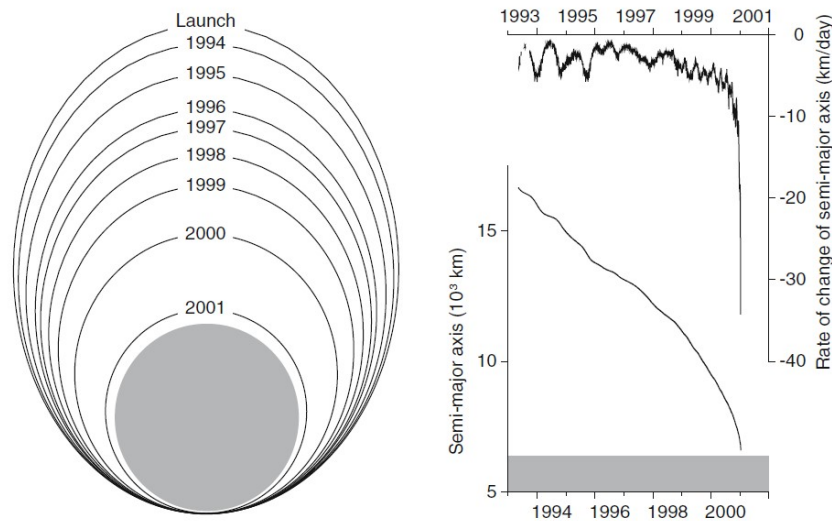


Figure 1.2: *Left*: The changing shape of the orbit from its launch until re-entry. *Right*: Time evolution of the semi-major axis and its rate of change. Credit (1)

#### 1.1.1.2. Lifetime Analysis

As shown in Figure 1.2, there is a rate of decay of the semi-major axis of LEO orbiting objects. This variation is mainly caused by changes in atmospheric density. An accurate modelling of these variations is important for predicting the lifetime of satellite missions.

#### 1.1.1.3. Re-Entry Operations

Most of the objects that re-enter into the denser layers of the atmosphere burn up and are disintegrated. Occasionally, though, some parts of these objects may reach the ground. Figure 1.3 shows what happened during the final week of the orbital lifetime of object 22659. Re-entry and collision events are rare and they normally occur over oceans or deserts, where there is not a high risk of damage. In some cases re-entries involve spacecrafts with large masses or carrying hazardous payloads, such as nuclear reactors. In such cases re-entry trajectories are as carefully controlled as possible, although density fluctuations may affect the accuracy of the control. It is also interesting to ponder that an uncertainty of just 15 minutes in the re-entry prediction, which is beyond our current capabilities, is equivalent to a displacement of about one sixth of the orbit, making the forecast of the re-entry region highly uncertain.

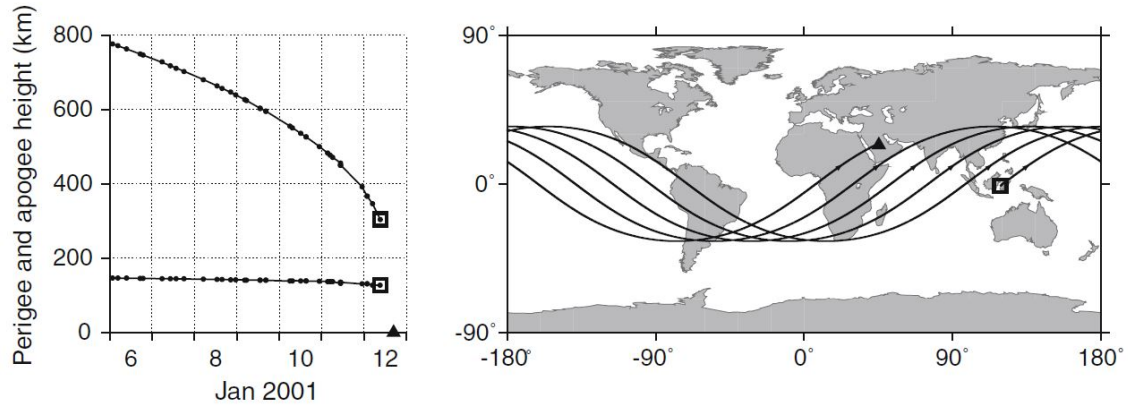


Figure 1.3: *Left*: Graph of apogee and perigee heights during the final lifetime prior re-entry. *Right*: Map of the predicted orbital ground track. Last TLE is indicated by an open square and re-entry point is indicated by a triangle. Credit (1)

#### 1.1.1.4. Manoeuvre Planning for Orbit Maintenance

Some satellites are required to follow a specific ground track pattern or remain within a certain altitude. The orbit decay of these types of satellite is compensated using thrusters. At high latitudes, the drift is driven by luni-solar perturbations, while at equator is driven by aerodynamic drag.

### 1.1.2. Structure and variability

The atmosphere is usually defined as a fragile, finite and thin layer of gases that surrounds our planet, but since the numerical density of particles forming the atmosphere decreases exponentially with altitude there is not a clear boundary between the atmosphere and outer space. Although density at orbital altitudes is at least a billion times smaller than at sea level, there's still drag force due to high orbital velocities. Figure 1.1 shows the variations of both temperature and density due to solar activity as defined in (1). As explained before, there is a temperature maximum (exospheric temperature) that is reached at the thermosphere. Notice that variations in temperature lead to density variations.

#### 1.1.2.1. Solar Activity Variation

One of the main factors that produce density and temperature variations is the solar activity. The thermosphere only absorbs extreme ultraviolet (EUV) and X-ray wavelengths, which can be extremely variable.

The amplitude of the density variation due to solar activity increases with altitude along the thermosphere, and it can vary up to a factor of 20. This density increase during peak solar activity periods is the dominant factor affecting spacecraft aerodynamic performances.

The index that measures the solar radiation is the  $F_{10.7}$ . It is a measure of the solar flux emitted at a wavelength of 10.7 cm measured in solar flux units (sfu), equivalent to  $10^4$  jansky ( $10^{-22}$  W/m·Hz).

### 1.1.2.2. Geomagnetic Activity Variation

Active regions are related to solar flares, which eject large amounts of charged particles into space. These particles are mostly deflected by the Earth's magnetic field, and only a fraction can enter the atmosphere through the polar cusps of the magnetosphere, causing geomagnetic storms usually accompanied by auroral displays. During these storms, vast amounts of energy are deposited in the polar regions of the thermosphere causing density variations up to one order of magnitude.

The index that measures the geomagnetic activity is called  $A_p$  Index, which is obtained from magnetic field variations measured at different locations for a given day. It is given in nanoTesla (nT).

### 1.1.3. Empirical density models

At orbital altitudes, the atmospheric composition varies with altitude, as shown in Fig 1.7. This composition fluctuates as the atmosphere expands and contracts under the influence of the solar cycle or geomagnetic activity.

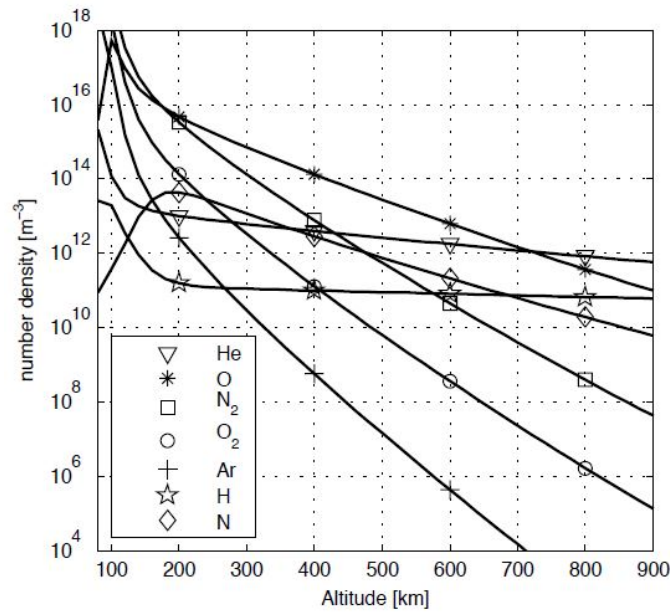


Figure 1.4: Variation of the Earth's atmospheric composition with altitude as defined by NRLM-SISE00 model.

There exists a wide variety of models which can be used to predict thermospheric characteristics which provide both neutral density and temperature, as functions of the spacecraft position, solar radiation and geomagnetic activity. Two of the most recent models are the NRLMSISE-00 and the JB2006.

The **NRLMSISE-00** model, developed by the US Naval Research Laboratory, provides, among other parameters, temperature, total mass density, and gas composition (in number density of each element or molecule). The required inputs are altitude, latitude, longitude and the two indexes  $F_{10.7}$  and  $A_p$ . It covers all the range from sea level to the exosphere.



The **JB2006** model provides neutral density and temperature from 120 km to the exosphere. The inputs are again the  $F_{10.7}$  and  $Ap$  indexes, but this model incorporates two new indexes for solar radiation, called S10 and Mg10, in order to obtain improved accuracy in density variation calculations.

The ECSS standard on Space environment estates that NRLMSISE-00 model should be used for calculating neutral temperature, gas composition and density below 120 km, while the JB2006 model should be used for computing the total density above 120 km.

## 1.2. Gas dynamics

The range of variation of atmospheric density values is huge. Density varies between values close to  $1 \text{ kg/m}^3$  near the Earth's surface, and down to  $10^{-15} \text{ kg/m}^3$  at the outermost layers of the thermosphere. These density variations have important consequences with respect to the behaviour of matter, which can be safely interpreted as a continuous flow at lower altitudes, but requires to be understood as a rarefied medium at altitudes  $\gtrsim 100 \text{ km}$ . We now describe the main parameters which determine the atmospheric fluid regimes.

### 1.2.1. Mean free path

The ideal gas model assumes that gas molecules only have collisions with the walls of the container, which produce the gas pressure, and not between each other. However, there are phenomena that are occasioned due to the collisions between molecules, such as diffusion. That's why it is important to know the distance travelled between collisions.

The mean free path of a gas is an important fundamental concept of the kinetic theory of gases. It is defined as the average distance traveled by a gas molecule between two consecutive collisions (see Figure 1.5). Gas mean free path may be estimated from kinetic theory too (10).

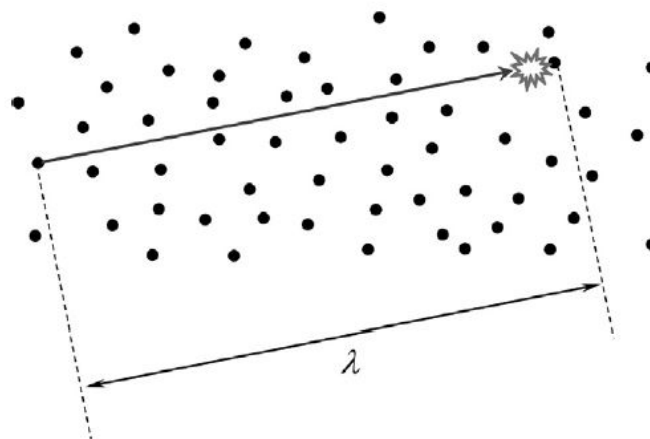


Figure 1.5: Mean free path ( $\lambda$ ). Credit (2)



### 1.2.2. Knudsen number

Whether a gas flow can be interpreted as a continuum or a rarefied medium depends on the relation between the molecular free path ( $\lambda$ ) and the characteristic dimensions of the flow field ( $L$ ). If the free path is much shorter than the dimensions of the flow field, an arbitrary molecule has a high chance of colliding with another, that is, the medium is dense and can be understood as a continuum. On the other hand, if  $\lambda$  is comparable to  $L$ , the flow is rarefied. The Knudsen number

$$Kn = \frac{\lambda}{L} \quad (1.2)$$

is a measure of how rarefied a fluid is. Non-negligible  $Kn$  values correspond to rarefied fluids. Besides,  $Kn$  is related to the Reynolds number  $Re$  and the Mach number  $M$  through the expression:

$$Kn = 1.26\sqrt{\gamma} \frac{M}{Re} \quad (1.3)$$

where  $\gamma$  is the isentropic exponent,  $Re$  is the Reynolds number, that is, a quotient between inertial and viscous forces in the fluid. A low  $Re$  is associated to a laminar fluid regime, and a high  $Re$  is associated to a turbulent regime. Finally  $M$  is the Mach number, that is, the quotient between the local fluid speed and the sound speed in the fluid.

### 1.2.3. Division of flow regimes

According to the range of the Knudsen number  $Kn$ , the flow regimes are:

$Kn < 0.01$	Continuum flow regime
$0.01 < Kn < 0.1$	Slip flow regime
$0.1 < Kn < 10$	Transitional regime
$Kn > 10$	Free molecular regime

Table 1.1: The regimes of gas dynamics in terms of  $Kn$ . Credit (7)

In the *slip flow regime*, some phenomena appear in the gas flow different from ordinary flows, such as velocity slip or temperature jump. Ordinary gas dynamics equations are still valid but it is necessary to introduce some modifications into the boundary conditions. The layer of gas immediately adjacent to the fluid surface cannot be assumed at rest, but moving with a non-negligible tangential velocity.

In the *transitional regime*, between the slip flow and the free molecular flow, the collisions with the surface and between molecules have the same importance, so the analysis of this type of flow becomes difficult and molecular gas dynamics methods must be applied.

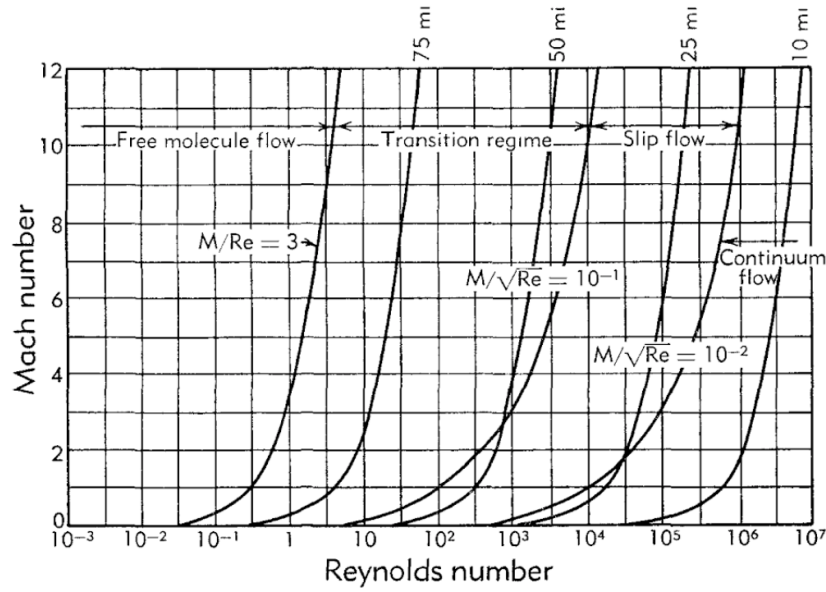


Figure 1.6: Different dynamic flow regimes in the Re-M diagram. Credit (3).

At large Knudsen numbers, where the gas is very rarefied, collisions of the gas molecules with the surface of the body prevail. When  $\lambda$  is much larger than  $L$ , the molecules reflected from the surface of the body collide with other molecules only after a large distance. Momentum and energy of the incoming flow can be easily calculated as the velocity distribution function is not influenced by the presence of the body. This regime is called the *free molecular flow*, and is the regime of interest in the study of spacecraft drag (see Figure 1.7).

The behaviour of flows, either in the continuum regime, in the rarefied regime, or in the intermediate cases (the slip flow, or the transition regime), is determined by the characteristic Mach number and Reynolds number of the flow (11; 3). Figure 1.6 shows the boundaries for the different regimes of fluid dynamics.

### 1.3. Free Molecular Flow

The relevant regime when studying gas dynamics at the characteristic heights of the thermosphere corresponds to the free molecular flow, that is, to very high Knudsen number  $Kn$ . The relation between  $Kn$  and density implies a variation of  $Kn$  as function of the solar activity. In Figure 1.7 we can see that high solar activity and thus increasing density leads  $Kn$  values almost one order of magnitude lower for typical heights of the lower thermosphere.

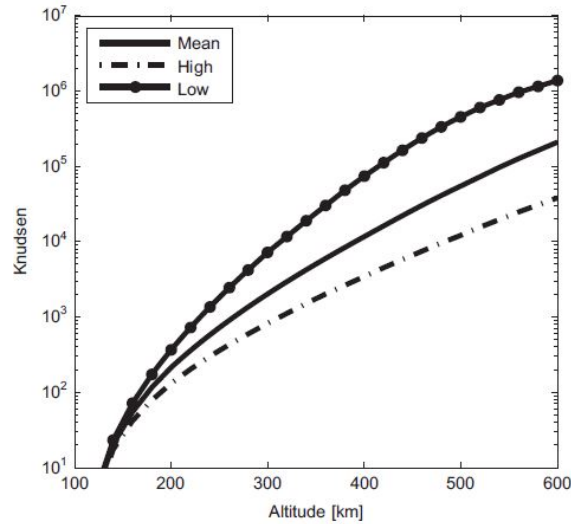


Figure 1.7: Variation of the Knudsen number versus height in the Earth's atmosphere. Credit (2).

Note that the safe approximation to the free molecular flow has an important consequence in terms of the interaction flow-spacecraft and, ultimately, on the calculation of drag coefficients which we will consider in the following chapters. The low number density of fluid particles allows the molecules which hit the body and are re-emitted to travel very far before they interact with other molecules. Thus the fluxes of incident and re-emitted particles can be treated separately, and only their interactions with the spacecraft are to be considered.

Besides, typical spacecraft velocities at the thermosphere are high enough that the relative velocity of the surrounding gas with respect to the spacecraft,  $v_{\text{rel}}$ , correspond to the hyperthermal case. In this case the random motion of gas particles can be neglected, because the associated velocities are much lower than  $v_{\text{rel}}$ .

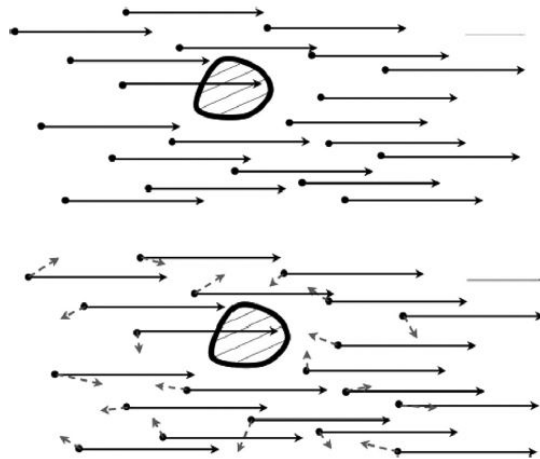


Figure 1.8: Sketch depicting a hyperthermal (upper panel) and a normal velocity fluid (lower panel). Credit (2).

## 1.4. Gas-Surface Interactions

As previously mentioned, under conditions of free molecular flow, collisions between molecules are extremely rare, even between incident and reflected particles. Then, reflected particles do not have a relevant effect on incident flow.

Assuming only gas-surface interaction, the force on the surface is related to the incident (subindex 'i') and reflected momentum (subindex 'r') as follows:

$$\frac{d\vec{f}}{dA} = (p_i + p_r)\vec{n} + (\tau_i - \tau_r)\vec{t} \quad (1.4)$$

Normal momentum flux is represented by  $p$  and tangential momentum flux by  $\tau$  (see Figure 1.9). Normal and tangential incident fluxes  $(p_i, \tau_i)$  depend on the incident velocity and the mass flux.

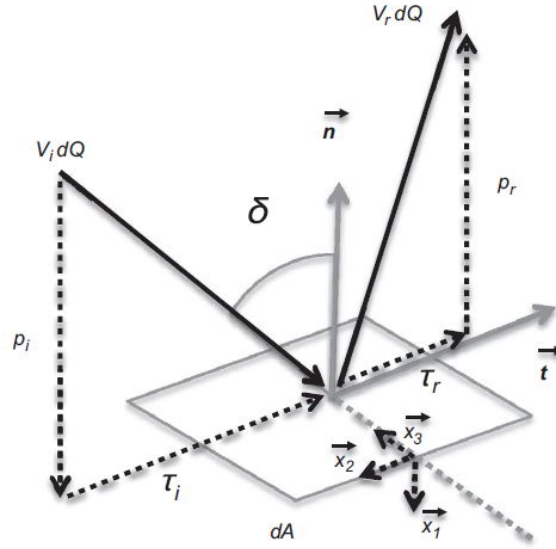


Figure 1.9: Incident and reflected fluxes on a convex body. Credit (2).

The incident fluxes are described by the following equations:

$$p_i = \frac{\rho V_i^2}{2s^2} \Gamma_1(s \cos \delta) \quad (1.5)$$

$$\tau_i = \frac{\rho V_i^2}{2s} \sin \delta \Gamma_2(s \cos \delta) \quad (1.6)$$

where

$$\Gamma_1(x) = \frac{1}{\sqrt{\pi}} \left\{ x \exp(-x^2) + \frac{\sqrt{\pi}}{2} (1 + 2x^2) [1 + \operatorname{erf}(x)] \right\} \quad (1.7)$$

$$\Gamma_2(x) = \frac{1}{\sqrt{\pi}} \left\{ x \exp(-x^2) + \sqrt{\pi} x [1 + \operatorname{erf}(x)] \right\} \quad (1.8)$$

$$\operatorname{erf}(x) = \frac{2}{\sqrt{\pi}} \int_0^x e^{-t^2} dt \quad (1.9)$$

Regarding Eqs. 1.5 and 1.6,  $\rho$  corresponds to the neutral density of the gas,  $V_i$  is the velocity of the incident molecules computed as the sum of the gas velocity and the orbital velocity, and  $s$  is the velocity ratio, which is calculated as the ratio between the orbital velocity and the gas velocity.

Figure 1.10 shows the normal and tangential momentum versus angle of attack for some specific data extracted from an NRL-MSISE00 online calculator. Having described the incident fluxes, the next key point is the determination of the reflected fluxes, which is a difficult and complex problem.

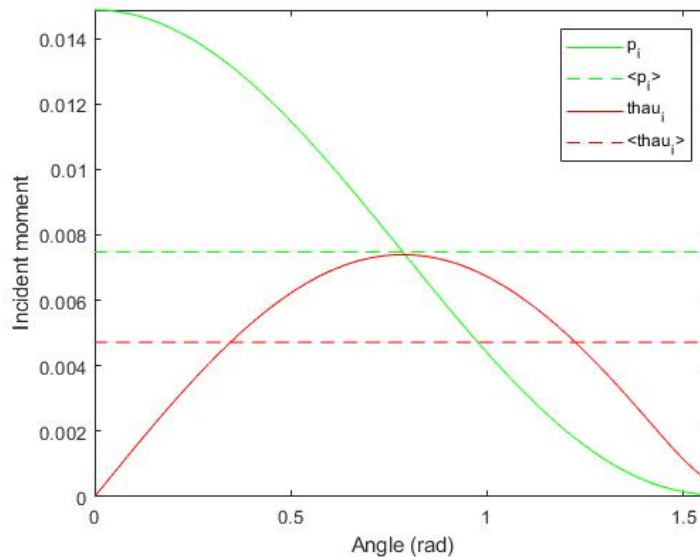


Figure 1.10: Incident normal and tangential momentum versus angle of attack.



# CHAPTER 2. SCHAAF AND CHAMBRE MODEL FOR DRAG COEFFICIENTS

## 2.1. Model description

Schaaf and Chambre (1958) model (3) for the calculation of the drag coefficient contemplates the calculation of incident and reflected (or re-emitted) fluxes of momentum, and of the energies of atmospheric particles incident on the spacecraft and reflected by it. It is based on the following assumptions:

- i) This model is restricted to the case of free molecular flow, convenient for the study of the dynamics of the thermosphere. Let us briefly recall that a free molecular flow is characterised by long mean free paths, and thus by the low probability of particle collisions. This allows to make the safe assumption that the flows of incident and reflected particles from the spacecraft will not interfere, and thus can be treated separately.
- ii) Only selected average parameters are necessary in order to characterise the interaction flow-spacecraft. No complete determination of the velocity distributions of reflected (or re-emitted) particles as a function of incident velocities is required.

### 2.1.1. Energy considerations

The energy accommodation coefficient  $\alpha$  was defined by Smoluchowski (1898) and Knudsen (1911), in terms of the incident and re-emitted energy flux ( $dE_i$  and  $dE_r$  respectively), and the energy flux which would be carried away if all incident molecules were re-emitted following a Maxwellian distribution corresponding to a surface temperature  $T_w$ :

$$\alpha = \frac{dE_i - dE_r}{dE_i - dE_w} \quad (2.1)$$

$\alpha$  close to zero corresponds to the case of negligible energy transfer between incident molecules and surface.  $\alpha$  values close to 1 correspond to cases in which  $dE_i \sim dE_w$ , and thus to a case in which reemitted particles have 'accommodated' their energies to the Maxwellian case.

### 2.1.2. Flux of momentum considerations

A proper account of the effects of linear momentum exchange requires to consider separately the normal ( $p$ ) and tangential ( $\tau$ ) components. Following the same nomenclature for subindexes as in the former subsection, (3) defined the following phenomenological coefficients <sup>1</sup>:

$$\sigma = \frac{\tau_i - \tau_r}{\tau_i - \tau_w} \quad \sigma' = \frac{p_i - p_r}{p_i - p_w} \quad (2.2)$$

---

<sup>1</sup>Note that all the coefficients described in this subsection are phenomenological averages. They may, and probably depend on other parameters, such as gas and surface temperatures, gas pressure, gas characteristics, and direction of the mean flow with respect to the spacecraft surface.

Perfect specular reflection ( $p_r = p_i$ ,  $\tau_r = \tau_i$ ) would correspond to  $\sigma = \sigma' = 0$ . If, additionally, accommodation were complete, then  $\alpha = 0$ . The opposite limit case, that is, complete diffuse reflection ( $p_r = p_w$ ,  $\tau_r = \tau_w$ ), would correspond to  $\sigma = \sigma' = \alpha = 1$ .

At the time in which the work by Schaaf and Chambre was presented, only experimental values of  $\alpha$  and  $\sigma$  at low velocities could be obtained. Such values were close to 1 (see Table A.1 from Appendix A), and thus supported the assumption of almost completely diffuse re-emission.

At this point we have the expressions for  $p_i$  (Equation 1.5) and  $\tau_i$  (Equation 1.6) obtained in Chapter 1. From Equation 2.2,  $p_r$ ,  $\tau_r$  can be expressed as functions of  $\sigma$ ,  $\sigma'$  and  $p_w$  (note that  $\tau_w=0$  for a Maxwellian distribution):

$$p_r = (1 - \sigma')p_i + \sigma'p_w \quad (2.3)$$

$$\tau_r = (1 - \sigma)\tau_i \quad (2.4)$$

Thus the net normal momentum flux ( $p$ ) and tangential momentum flux ( $\tau$ ) are:

$$p = p_i + p_r = (2 - \sigma')p_i + \sigma'p_w \quad (2.5)$$

$$\tau = \tau_i - \tau_r = \sigma\tau_i \quad (2.6)$$

We know the relation between force per unit area:

$$\frac{d\vec{F}}{dA} = (p_i + p_r)\vec{n} + (\tau_i - \tau_r)\vec{\tau} \quad (2.7)$$

and finally, from the expression for the drag coefficient as function of the drag acceleration (immediate from equation 1.1), we can obtain the geometry-dependant drag coefficient expression either for the flat plate or for the sphere, which are going to be analyzed in the following sections.

Along the next subsections we analyze the variation of  $C_D$  as a function of spacecraft wall temperature,  $T_w$ ,  $\sigma$  and angle of attack of the incident molecules,  $\theta$ , for two basic spacecraft geometries: spherical and flat plate.

## 2.2. Flat plate

For a flat plate at an angle of attack  $\theta$  and one side exposed to the flow has for specular reflection

$$C_{D,spec} = \frac{4 \sin \theta}{\sqrt{\pi} S^2} \left[ (S \sin \theta) e^{-(S \sin \theta)^2} + \sqrt{\pi} \left[ \frac{1}{2} + (S \sin \theta)^2 \right] \text{erf}(S \sin \theta) \right] \quad (2.8)$$

and for diffuse reflection

$$C_{D,diff} = \frac{2}{\sqrt{\pi} S} \left[ e^{-(S \sin \theta)^2} + \sqrt{\pi} S \sin \theta \left( 1 + \frac{1}{2 S^2} \right) \text{erf}(S \sin \theta) + \frac{\pi S}{S_w} \sin^2 \theta \right] \quad (2.9)$$

where:

$$S_w = \frac{U}{\sqrt{2RT_w}} \quad (2.10)$$

All models studied in this section have the following nomenclature for and easier identification:  $M - F - T_w T_w T_w - \sigma - \theta \theta$



### 2.2.1. Completely specular reflection

In this model, the study will consider completely specular reflection, which means that  $\sigma = 0$ . We will only consider the case in which the angle of attack is varied (from  $0^\circ$  to  $90^\circ$ ). Note that the equation corresponding to the specular case (Equation 2.8) does not include  $T_w$ , and thus we will disregard its effect at this point of the project.

Table 2.1 shows the values of the parameters for each case. The five cases are plotted in the same figure in order to facilitate analysis and comparison (see Figure 2.1).

Model	$T_w(^{\circ}C)$	$\sigma$	$\theta(^{\circ})$
M-F-ANY-0-var	Any	0	$[0,90]$

Table 2.1: Analyzed Schaaf and Chambre model for flat plate and specular reflection.  $\theta = [0, 90]^\circ$  and  $\sigma = 0$ .

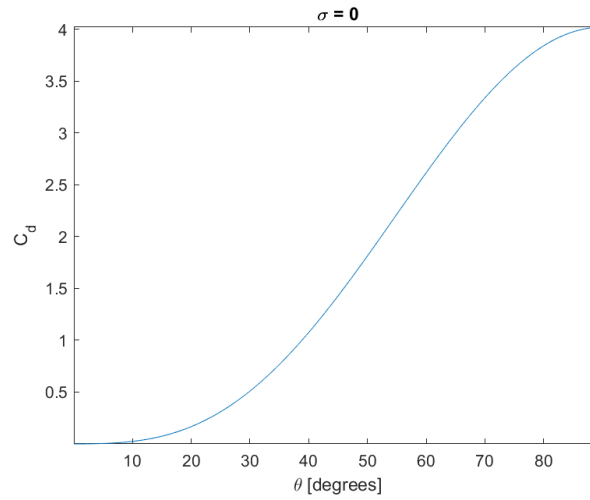


Figure 2.1: Analyzed Schaaf and Chambre models for flat plate and specular reflection.  $\theta = [0, 90]^\circ$  and  $\sigma = 0$ .

Figure 2.1 shows that  $C_D$  increases significantly with the angle of attack as its value goes from almost 0 to around 4.

### 2.2.2. Completely diffuse reflection

We now present results for the completely diffuse case, in which  $\sigma = 1$ . Two studies will be performed, one with constant  $T_w$ , and another with constant  $\theta$ .

#### 2.2.2.1. Constant $T_w$

In this case,  $T_w$  will be kept constant along each calculation, while the angle of attack will be varied from  $0^\circ$  to  $90^\circ$ . The drag coefficient will be computed for several temperatures. Table 2.2 shows the values of the parameters for each case.

Model	$T_w(^{\circ}C)$	$\sigma$	$\theta(^{\circ})$
M-F-000-0-var	0	1	[0,90]
M-F-025-0-var	25	1	[0,90]
M-F-050-0-var	50	1	[0,90]
M-F-075-0-var	75	1	[0,90]
M-F-100-0-var	100	1	[0,90]

Table 2.2: Analyzed Schaaf and Chambre models for flat plate and diffuse reflection.  $\theta = [0, 90]^{\circ}$ ,  $\sigma = 1$  and  $T_w = [0 \ 25 \ 50 \ 75 \ 90]^{\circ}C$ .

Figure 2.2 shows the effects of the angle of attack at different temperatures. In the diffuse cases, there is a slight variation with  $T_w$  that increases with  $\theta$ . Despite existing an increase in  $C_D$  as temperature increases, the variation is practically negligible.

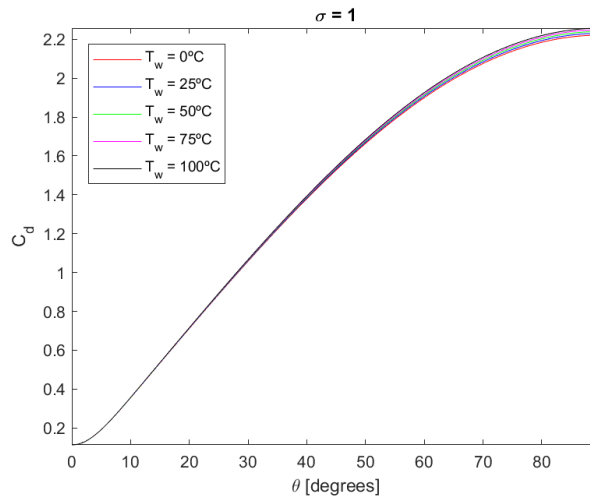


Figure 2.2: Analyzed Schaaf and Chambre models for flat plate and diffuse reflection.  $\theta = [0, 90]^{\circ}$ ,  $\sigma = 1$  and  $T_w = [0 \ 25 \ 50 \ 75 \ 90]^{\circ}C$ .

#### 2.2.2.2. Constant $\theta$

In this case,  $\theta$  will be kept constant along each calculation, while the temperature of the wall will be varied from  $0^{\circ}C$  to  $100^{\circ}C$ . The drag coefficient will be computed for several angles of attack. Table 2.3 shows the values of the parameters for each case.

Model	$T_w(^{\circ}C)$	$\sigma$	$\theta(^{\circ})$
M-F-var-0-00	[0,100]	1	0
M-F-var-0-15	[0,100]	1	15
M-F-var-0-30	[0,100]	1	30
M-F-var-0-45	[0,100]	1	45
M-F-var-0-60	[0,100]	1	60
M-F-var-0-75	[0,100]	1	75
M-F-var-0-90	[0,100]	1	90

Table 2.3: Analyzed Schaaf and Chambre models for flat plate and diffuse reflection.  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\sigma = 1$  and  $T_w = [0, 100]^{\circ}C$ .

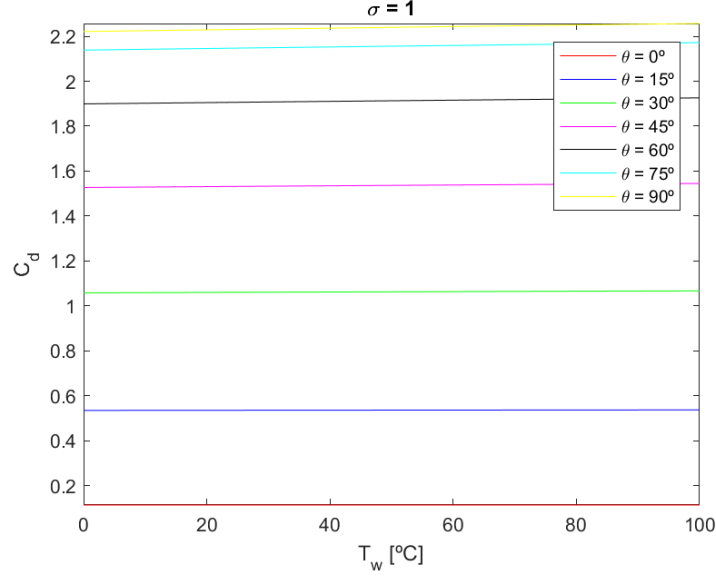


Figure 2.3: Analyzed Schaaf and Chambre models for flat plate and diffuse reflection.  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\sigma = 1$  and  $T_w = [0, 100]^\circ C$ .

Figure 2.3 is consistent with Figure 2.2, that is, there's not a significant variation in  $C_D$  with temperature. It is reasonable to assume that  $T_w$  is not relevant when computing the drag coefficient, whereas the angle of attack must definitely be taken into consideration.

### 2.2.3. Specular - Diffuse reflection

In this case, the analysis intends to show the difference in  $C_D$  as  $\sigma$  varies from specular reflection to diffuse reflection. This study will be performed fixing  $T_w = 50^\circ C$  and varying the angle of attack. We fix  $T_w$  because, as seen previously, it does not play a relevant role in  $C_D$  calculation. Table 2.4 shows the values of the parameters for each model.

Figure 2.4 shows that there is an important variation of  $C_D$  for different angles of attack. It can also be seen that the larger the angle of attack, the greater the variation of  $C_D$  from specular to diffuse reflection.

Model	$T_w(^{\circ}C)$	$\sigma$	$\theta(^{\circ})$
M-F-050-var-00	50	[0,1]	0
M-F-050-var-15	50	[0,1]	15
M-F-050-var-30	50	[0,1]	30
M-F-050-var-45	50	[0,1]	45
M-F-050-var-60	50	[0,1]	60
M-F-050-var-75	50	[0,1]	75
M-F-050-var-90	50	[0,1]	90

Table 2.4: Analyzed Schaaf and Chambre models for flat plate and transition from specular to diffuse reflection.  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\sigma = [0, 1]$  and  $T_w = 50^\circ C$ .

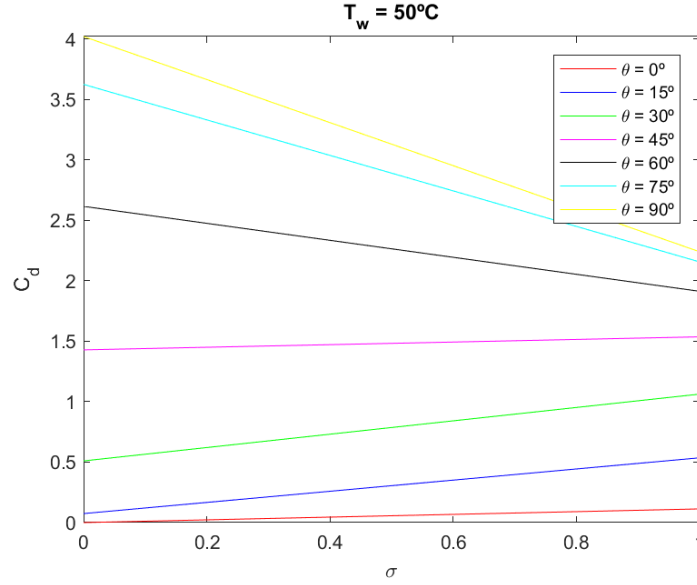


Figure 2.4: Analyzed Schaaf and Chambre models for flat plate and transition from specular to diffuse reflection.  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\sigma = [0, 1]$  and  $T_w = 50^\circ\text{C}$ .

## 2.3. Sphere

Spherical geometry implies a reduction of input parameters, as angles of attack will not be considered. The analysis structure of the different cases is the same as for the flat plate. A sphere with projected area  $A$  has the following drag coefficient for diffuse reflection:

$$C_{D,diff} = \frac{e^{-S^2/2}}{\sqrt{\pi}S^2}(1 + 2S^2) + \frac{4S^4 + 4S^2 - 1}{2S^4} \text{erf}(S) + \frac{2\sqrt{\pi}}{3S_w} \quad (2.11)$$

and for specular reflection:

$$C_{D,spec} = C_{D,diff} - \frac{2\sqrt{\pi}}{3S_w} \quad (2.12)$$

where:

$$S_w = \frac{U}{\sqrt{2RT_w}} \quad (2.13)$$

All models studied in this section have the following nomenclature for an easier identification:  $M - S - T_w T_w T_w - \sigma$

### 2.3.1. Completely specular reflection

In this case, the only difference with respect to the flat plate is that there's no need to take into account the angles of attack. Table 2.5 shows the values of  $T_w$  and  $\sigma$  for this model.

Model	$T_w(^{\circ}C)$	$\sigma$
M-S-var-0	[0,100]	0

Table 2.5: Analyzed Schaaf and Chambre models for sphere and specular reflection.  $\sigma = 0$  and  $T_w = [0, 100]^{\circ}C$ .

For this model, we only have one calculation, as the only parameter that varies is the temperature. In this case,  $C_D$  adopts a constant value equal to 2.02, see Figure B.1 from Appendix B, as  $C_D$  (Equation 2.12) does not depend on  $T_w$  when reflection is specular. This happens due to the fact that, in this case, molecules do no exchange temperature with the surface so re-emission is independent of  $T_w$ .

### 2.3.2. Completely diffuse reflection

Again, the only relevant parameter is  $T_w$  (from  $0^{\circ}C$  to  $100^{\circ}C$ ), and the difference is that  $\sigma = 1$ . Table 2.6 shows the values of the parameters for this model.

Model	$T_w(^{\circ}C)$	$\sigma$
M-S-var-1	[0,100]	1

Table 2.6: Schaaf and Chambre models for sphere. Diffuse reflection.  $T_w = [0, 100]^{\circ}C$ .

In this case, it can be observed from Figure 2.5 that when diffuse reflection occurs,  $T_w$  plays a role in  $C_D$  calculation. However, it can be seen that the variation of  $C_D$  due to temperature changes is smaller than 0.1, so  $T_w$  would not be a determining factor.

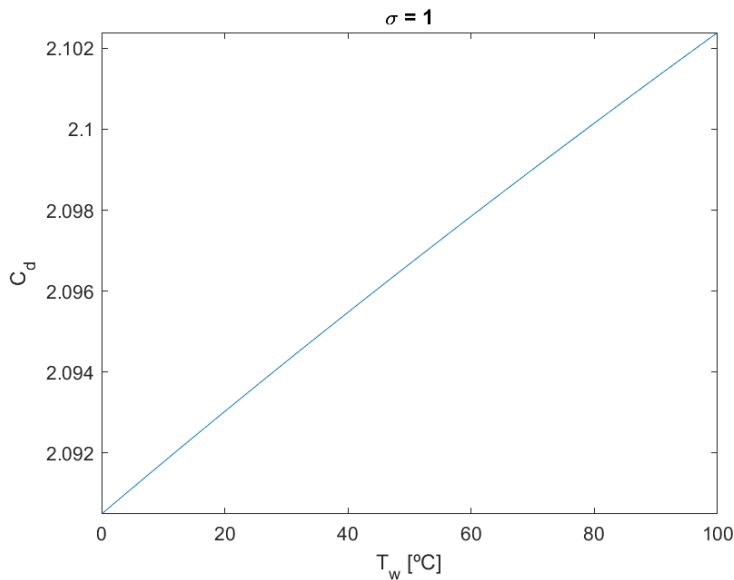


Figure 2.5: Schaaf and Chambre models for sphere. Diffuse reflection.  $T_w = [0, 100]^{\circ}C$ .

### 2.3.3. Transition Specular - Diffuse reflection

In this case the value of  $\sigma$  will vary from 0 to 1 and different values of  $T_w$  will be analyzed. Table 2.7 shows the values of the parameters for each model.

Model	$T_w(^{\circ}C)$	$\sigma$
M-S-000-var	0	[0,1]
M-S-025-var	25	[0,1]
M-S-050-var	50	[0,1]
M-S-075-var	75	[0,1]
M-S-100-var	100	[0,1]

Table 2.7: Schaaf and Chambre models for sphere. Transition specular-diffuse reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^{\circ}C$ .

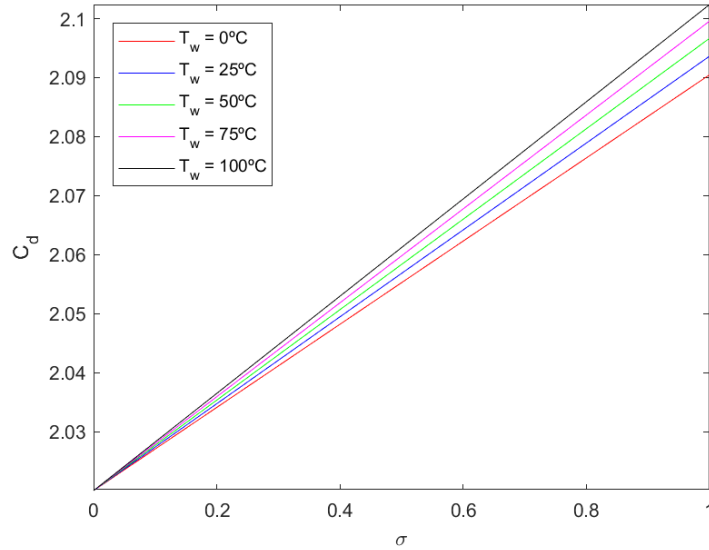


Figure 2.6: Schaaf and Chambre models for sphere. Transition specular-diffuse reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^{\circ}C$ .

Figure 2.6 ratifies what we have seen in the specular and diffuse cases. For specular reflection,  $T_w$  does not affect when computing the  $C_D$ , but as  $\sigma$  approaches to the diffuse reflection value, variation of  $T_w$  implies a variation of  $C_D$ . Again, the variation of  $C_D$  is too low to be relevant.

# CHAPTER 3. SCHAMBERG MODEL FOR DRAG COEFFICIENTS

## 3.1. Model description

An alternative model for surface interactions was proposed by Schamberg (5) in 1958 and is used to calculate force coefficient of flat plate and convex bodies in "hypersonic" free-molecule flow, in which the random thermal motion of molecules is negligible relative to the speed of the body. We are going to use formulae to estimate the effects of uncertainty in the surface interaction on the drag of satellites for flat plate and spherical geometries.

Mean free path of molecules in the gas is so great that the frequency of collisions between gas molecules is entirely negligible relative to the frequency of collisions between gas molecules and the surface of the body analyzed. However, these interaction effects are not understood in detail and there's no sufficient empirical information available.

The analytic formulation of an hypothesis for the gas-surface interaction in a rarefied gas goes back to Maxwell who postulated that a certain fraction,  $\sigma$ , of the incident molecules are temporarily absorbed by the surface and, after a certain time, re-emitted diffusely with a mean temperature equal to the temperature of the wall, while the rest of molecules are assumed to be reflected specularly with  $\theta_r = \theta_i$  and  $V_r = V_i$ .

Von Smoluchowski introduced the concept of "accommodation" to account for his experiments on heat transfer, which showed the re-emission velocity of gas molecules is determined by a temperature of re-emission,  $T_r$ , which takes an intermediate value between the incident temperature and the surface temperature, that is

$$T_r = T_i + \alpha(T_w - T_i) \quad (3.1)$$

where  $\alpha$  is a positive number called the "accommodation coefficient" and its value varies between 0 and 1.

The forces on and heat transfer to simple bodies in free-molecule flow have been calculated by several scientists using the empirical "constants"  $\sigma$  and  $\alpha$  as arbitrary parameters to describe the interaction between the gas and the surface, and these results are summarized in the Schaaf and Chambre model (3). As  $\sigma$  and  $\alpha$  denote the exchange of tangential momentum and energy, a third coefficient was introduced later,  $\sigma'$ , to parametrize the normal momentum exchange. The fact is that there's no empirical information to justify that these coefficients are "constants" which do not depend on the angle of incidence. The alternative model for describing gas-surface interaction proposed by Schamberg intends an improvement about the aspect by choosing different parameters which are expected to be more independent of the local angle of attack of the body.

Even using simple coefficients like  $\sigma$  and  $\alpha$ , the interaction model leads to lengthy expressions that are hard to solve, but fortunately, for calculating the drag of earth satellites, a simplification is possible. This can be done because at altitudes lower than 1000 miles, where the atmosphere is sufficiently dense to produce appreciable neutral-particle drag, the satellite speed is, at least, 6 times bigger than the thermal speed of the molecules, so the drag can be approximated by its asymptotic values for infinite speed ratio and can be

calculated with much simpler equations. For example, for a flat plate at a speed ratio of 6 and at  $45^\circ$  angle of attack, the drag coefficient differs from that for infinite speed ratio by only a 6%, and this accuracy, apparently, is more than adequate.

### 3.1.1. Analytic representation of surface interaction

We consider a uniform, parallel stream of molecules colliding with a flat surface at an angle of incidence  $\theta_i$  and incident velocity  $V_i$ . After hitting the surface, the molecules are expected to be re-emitted or reflected in a conical or wedge-shaped beam having a half-angular width denoted by  $\phi_0$ . The direction of re-emission is denoted as  $\theta_r$ .

There's a need to make some phenomenological assumptions concerning three questions(5):

- How does the angles of reflection  $\theta_r$  vary with the angles of incidence  $\theta_i$ , the incident speed  $V_i$ , type of gas and surface, etc?
- What is the distribution of the number and speed of re-emitted molecules within the reflected beam?
- How does the speed  $V_r$  of the reflected molecules vary with the incident speed  $V_i$ , and with the temperature and nature of the surface?

First of all we need to parametrize the relation between the angles of reflection and incidence somehow including the limiting cases of specular reflections, in which  $\theta_i = \theta_r$ , and diffuse reflection, in which  $\theta_r = 90^\circ$  for any value of  $\theta_i$ . It looks like  $\theta_r \geq \theta_i$ . The following form is chosen as being both simple and convenient for easier manipulation:

$$\cos \theta_r = (\cos \theta_i)^v, \quad v \geq 1 \quad (3.2)$$

Then for specular reflection

$$v = 1, \quad \cos \theta_r = \cos \theta_i \quad (3.3)$$

and for diffuse reflection

$$v \rightarrow \infty, \quad \theta_r = \frac{\pi}{2} \quad (3.4)$$

There is evidence that the distribution of molecules is symmetric about the axis of the beam and can be approximated by a cosine distribution:

$$n_r(\phi) = K \cos\left(\frac{\phi}{\phi_0} \cdot \frac{\pi}{2}\right) \quad (3.5)$$

where  $n_r$  is the number of re-emitted molecules per unit time whose direction of re-emission lies between  $\phi$  and  $\phi + d\phi$  and  $K$  is a constant of proportionality that depends on the total number of molecules re-emitted per unit time and on whether the re-emitted beam is wedge-shaped or conical.

Little is known about the distribution of speeds, so it will be assumed that re-emission speed  $V_r$  is constant independent of re-emission angles  $\theta_r$ .



The magnitude of the speed  $V_r$  can be related to the incident velocity  $V_i$  and the temperature of the surface,  $T_w$ , by using the coefficient  $\alpha$ ,

$$\frac{V_r}{V_i} = \sqrt{\frac{T_r}{T_i}} = \sqrt{1 + \alpha \left( \frac{T_w}{T_i} - 1 \right)} \quad (3.6)$$

### 3.1.2. Drag coefficient of flat plates and convex bodies

The drag coefficient is defined in the conventional manner as

$$C_D = \frac{D}{\frac{1}{2} \rho_i V_i^2 A} \quad (3.7)$$

It turns out that the hyperthermal drag coefficients for all geometries can be expressed in this simple form:

$$C_D = 2 \left[ 1 + \phi(\phi_0) \frac{V_r}{V_i} f(v, shape) \right] \quad (3.8)$$

The first term inside the bracket, unity, represents the drag that the incident molecules would produce if all of them stuck to the body ( $V_r = 0$ ), that takes the value 2 and it's independent of the body shape.

The second term represents the contribution to the drag of the momentum of the molecules which are re-emitted from the surface of the body. It can be positive or negative depending on the direction of re-emission.

The factor  $\phi(\phi_0)$  is defined as:

$$\phi(\phi_0) = \frac{1 - (2\phi_0/\pi)^2}{1 - 4(2\phi_0/\pi)} \cdot \frac{\frac{1}{2} \sin 2\phi_0 - (2\phi_0/\pi)}{\sin \phi_0 - (2\phi_0/\pi)} \quad (3.9)$$

The term  $\phi_0$  is shown in Figure 3.1. The values for the limiting cases are for specular reflection  $\phi(0) = 1$  and for diffuse reflection  $\phi(\frac{\pi}{2}) = \frac{2}{3}$ .

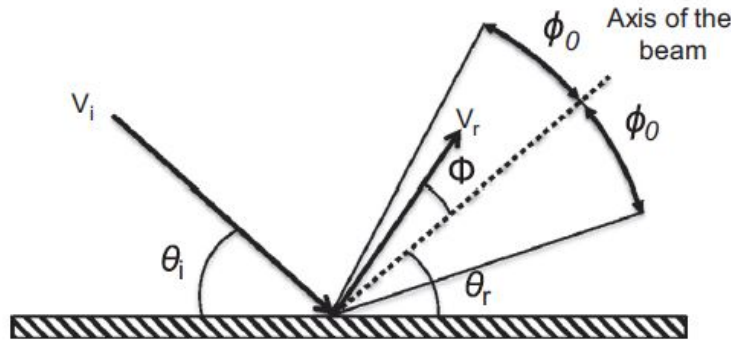


Figure 3.1: Schamberg's GSIM. Credit (2)

The factor  $f(v, shape)$  accounts for the combined effect of the reflection law of Eq.(3.2) and the shape of the body.

It is defined for flat plate at an angle of attack  $\theta_i$  as

$$f(v, \theta_i) = [\sin \theta_i \sqrt{1 - \cos^{2v} \theta_i} - (\cos \theta_i)^{v+1}] \quad (3.10)$$

and for sphere as

$$f(v) = 2 \left[ I_1(v) - \frac{1}{v+3} \right] \quad (3.11)$$

where

$$I_1(v) = \int_0^1 x \sqrt{(1 - x^{2v})(1 - x^2)} \, dx \quad (3.12)$$

the limiting values of this definite integral are

$$I_1(v = 1) = \frac{1}{4}; \quad I_1(v \rightarrow \infty) = \frac{1}{3} \quad (3.13)$$

## 3.2. Flat plate

All models studied in this section have the following nomenclature for an easier identification:  $M - F - T_w T_w T_w - \alpha \alpha - \theta \theta - \phi - \frac{1}{v}$

### 3.2.1. Completely specular reflection

In this model, the study will consider completely specular reflection (that is,  $\phi_0 = 0^\circ$ ), which means that  $\phi = 1$  and  $\frac{1}{v} = 1$ . Two different cases will be studied, one in which  $T_w$  and  $\alpha$  will be kept constant, and the other in which  $T_w$  and  $\theta$  will be kept constant.

#### 3.2.1.1. Constant $T_w$ and $\alpha$

In this case, the study will represent the drag coefficient for a flat plate and specular reflection. To do so, the different models have been structured keeping  $\alpha$  constant among each computation and for each accommodation factor, several  $T_w$  values have been used [0 50 100] °C. Table 3.1 shows the parameter values for each model. Some other intermediate values have also been computed.

Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-00-var-1-1	0	0	[0,90]	1	1
M-F-050-00-var-1-1	50	0	[0,90]	1	1
M-F-100-00-var-1-1	100	0	[0,90]	1	1
M-F-000-05-var-1-1	0	0.5	[0,90]	1	1
M-F-050-05-var-1-1	50	0.5	[0,90]	1	1
M-F-100-05-var-1-1	100	0.5	[0,90]	1	1
M-F-000-10-var-1-1	0	1	[0,90]	1	1
M-F-050-10-var-1-1	50	1	[0,90]	1	1
M-F-100-10-var-1-1	100	1	[0,90]	1	1

Table 3.1: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\theta = [0, 90]^{\circ}$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ .

After structuring the different models, it's time for calculation. When designing the plots, two different combinations have been computed. Figure 3.2 is representing  $C_D$  vs.  $\theta$  and each colored line corresponds to a constant value of  $\alpha$ , while each line style corresponds to a different value of  $T_w$ . To validate the information shown in the first plot, Figure 3.3 shows  $C_D$  vs.  $\theta$ , but in this case, each colored line represents a constant value of  $T_w$  while each line style corresponds to a different value of  $\alpha$ .

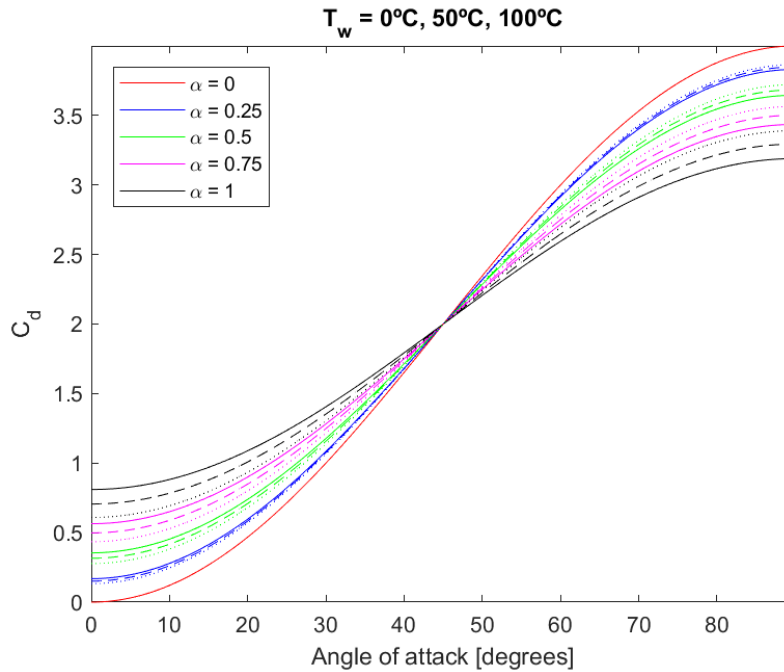


Figure 3.2: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\theta = [0, 90]^{\circ}$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ . Note that solid lines correspond to  $T_w = 0^{\circ}C$ , dashed lines correspond to  $T_w = 50^{\circ}C$  and dotted lines correspond to  $T_w = 100^{\circ}C$ .

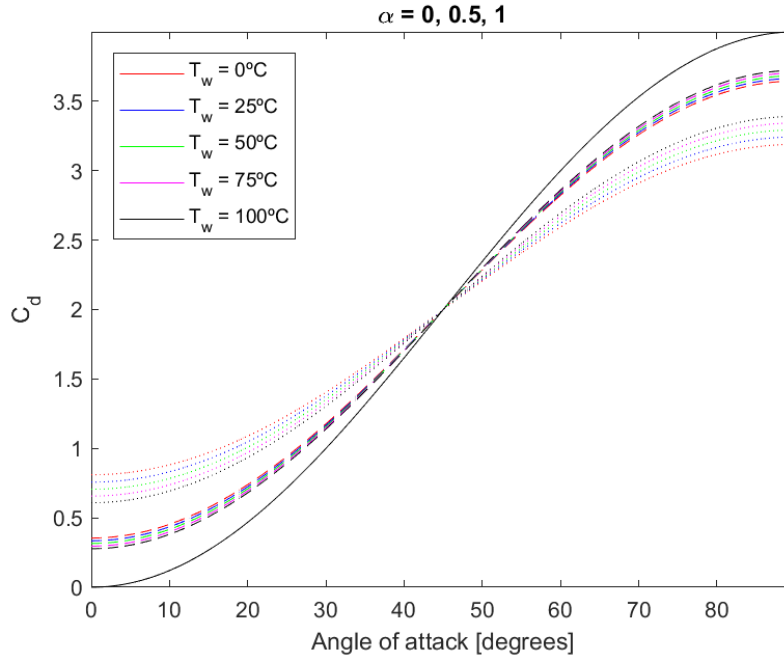


Figure 3.3: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^\circ\text{C}$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\theta = [0, 90]^\circ$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ . Note that solid lines correspond to  $\alpha = 0$ , dashed lines correspond to  $\alpha = 0.5$  and dotted lines correspond to  $\alpha = 1$ .

From both Figure 3.2 and Figure 3.3 it can be seen that the angle of attack plays a very important role when computing the drag coefficient as the variation of  $C_D$  is very big. Regarding the accommodation coefficient and the different values of  $T_w$ , it can be affirmed, as expected, that the greater the accommodation coefficient, the greater  $C_D$  variation with temperature. This makes perfect sense due to the fact that, for  $\alpha = 0$  there is no accommodation of the molecule with the surface so the temperature of the surface is not playing any role there, but when accommodation appears,  $T_w$  starts to take center stage due to the heat exchange between the molecule and the surface, so this involves an energy exchange then the value of  $C_D$  varies depending on the amount of energy transferred.

#### 3.2.1.2. Constant $T_w$ and $\theta$

In this case, the study will be almost the same as the previous case but the varying parameters. Now,  $T_w$  and  $\theta$  will be kept constant and  $\alpha$  will vary from 0 to 1. Table 3.2 shows the structure of the different models studied.

Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-var-00-1-1	0	[0,1]	0	1	1	M-F-075-var-45-1-1	75	[0,1]	45	1	1
M-F-025-var-00-1-1	25	[0,1]	0	1	1	M-F-100-var-45-1-1	100	[0,1]	45	1	1
M-F-050-var-00-1-1	50	[0,1]	0	1	1	M-F-000-var-60-1-1	0	[0,1]	60	1	1
M-F-075-var-00-1-1	75	[0,1]	0	1	1	M-F-025-var-60-1-1	25	[0,1]	60	1	1
M-F-100-var-00-1-1	100	[0,1]	0	1	1	M-F-050-var-60-1-1	50	[0,1]	60	1	1
M-F-000-var-15-1-1	0	[0,1]	15	1	1	M-F-075-var-60-1-1	75	[0,1]	60	1	1
M-F-025-var-15-1-1	25	[0,1]	15	1	1	M-F-100-var-60-1-1	100	[0,1]	60	1	1
M-F-050-var-15-1-1	50	[0,1]	15	1	1	M-F-000-var-75-1-1	0	[0,1]	75	1	1
M-F-075-var-15-1-1	75	[0,1]	15	1	1	M-F-025-var-75-1-1	25	[0,1]	75	1	1
M-F-100-var-15-1-1	100	[0,1]	15	1	1	M-F-050-var-75-1-1	50	[0,1]	75	1	1
M-F-000-var-30-1-1	0	[0,1]	30	1	1	M-F-075-var-75-1-1	75	[0,1]	75	1	1
M-F-025-var-30-1-1	25	[0,1]	30	1	1	M-F-100-var-75-1-1	100	[0,1]	75	1	1
M-F-050-var-30-1-1	50	[0,1]	30	1	1	M-F-000-var-90-1-1	0	[0,1]	90	1	1
M-F-075-var-30-1-1	75	[0,1]	30	1	1	M-F-025-var-90-1-1	25	[0,1]	90	1	1
M-F-100-var-30-1-1	100	[0,1]	30	1	1	M-F-050-var-90-1-1	50	[0,1]	90	1	1
M-F-000-var-45-1-1	0	[0,1]	45	1	1	M-F-075-var-90-1-1	75	[0,1]	90	1	1
M-F-025-var-45-1-1	25	[0,1]	45	1	1	M-F-100-var-90-1-1	100	[0,1]	90	1	1
M-F-050-var-45-1-1	50	[0,1]	45	1	1						

Table 3.2: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^{\circ}C$ ,  $\alpha = [0,1]$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ .

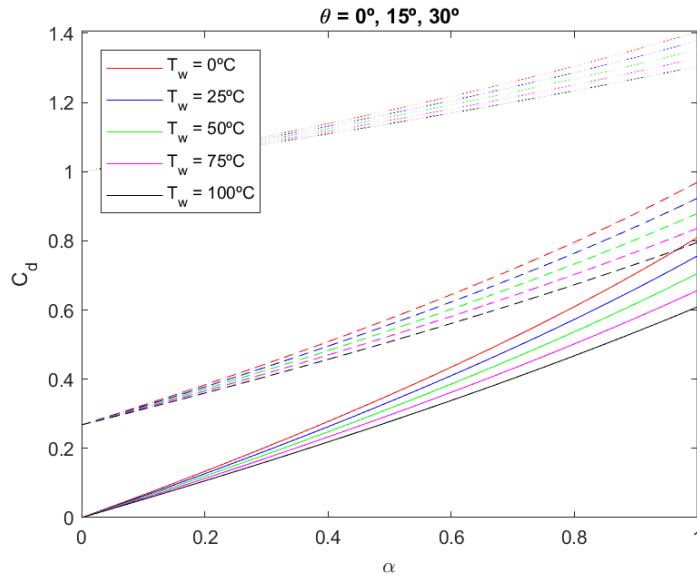


Figure 3.4: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^{\circ}C$ ,  $\alpha = [0,1]$ ,  $\theta = [0 \ 15 \ 30]^{\circ}$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ . Note that solid lines correspond to  $\theta = 0^{\circ}$ , dashed lines correspond to  $\theta = 15^{\circ}$  and dotted lines correspond to  $\theta = 30^{\circ}$ .

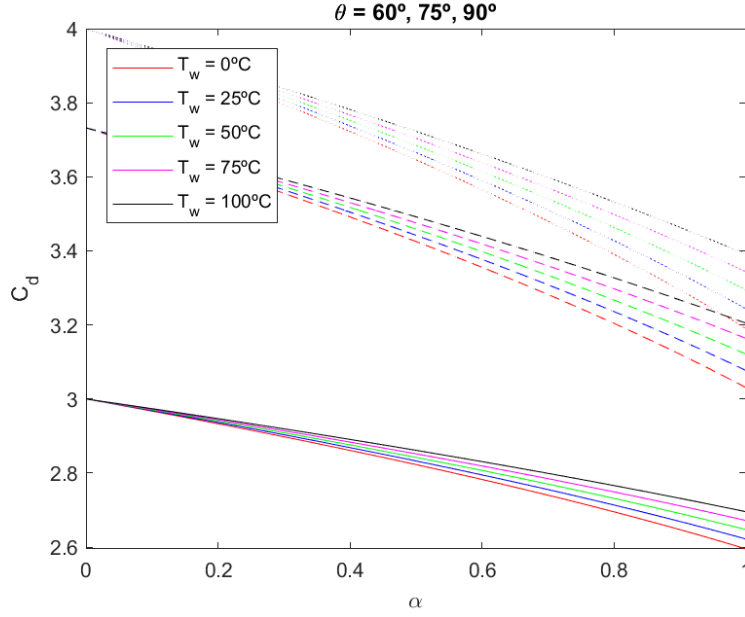


Figure 3.5: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0\ 25\ 50\ 75\ 100]^\circ\text{C}$ ,  $\alpha = [0,1]$ ,  $\theta = [60\ 75\ 90]^\circ$ ,  $\phi = 1$  and  $\frac{1}{V} = 1$ . Note that solid lines correspond to  $\theta = 60^\circ$ , dashed lines correspond to  $\theta = 75^\circ$  and dotted lines correspond to  $\theta = 90^\circ$ .

As it can be seen from the two plots, variation of  $C_D$  is highly related with the angle of attack. Figure 3.4 and Figure 3.5 show again that when there's no accommodation,  $T_w$  does not cause a variation in  $C_D$ , but for  $\alpha = 1$  it does. Note that the difference in  $C_D$  caused by the temperature of the surface is greater for the extreme cases than for the ones close to  $\theta = 45^\circ$ , in which  $T_w$  does not affect even if accommodation occurs. In the case in which  $\theta = 45^\circ$ , see Figure B.2 from Appendix B, it can be seen that there's no variation in  $C_D$ .

### 3.2.2. Completely diffuse reflection

In this model, the study will consider completely diffuse reflection, which means that  $\phi = \frac{2}{3}$  and  $\frac{1}{V} = 0$ . Two different cases will be studied, one in which  $T_w$  and  $\alpha$  will be kept constant, and the other in which  $T_w$  and  $\theta$  will be kept constant.

#### 3.2.2.1. Constant $T_w$ and $\alpha$

In this case, the study will represent the drag coefficient for a flat plate and diffuse reflection. To do so, the different models have been structured keeping  $\alpha$  constant among each computation and for each accommodation factor, several  $T_w$  values have been used  $[0\ 50\ 100]^\circ\text{C}$ . Table 3.3 shows the parameter values for each model. Some other intermediate values have also been computed.

Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-00-var-2/3-0	0	0	[0,90]	$\frac{2}{3}$	0
M-F-050-00-var-2/3-0	50	0	[0,90]	$\frac{2}{3}$	0
M-F-100-00-var-2/3-0	100	0	[0,90]	$\frac{2}{3}$	0
M-F-000-05-var-2/3-0	0	0.5	[0,90]	$\frac{2}{3}$	0
M-F-050-05-var-2/3-0	50	0.5	[0,90]	$\frac{2}{3}$	0
M-F-100-05-var-2/3-0	100	0.5	[0,90]	$\frac{2}{3}$	0
M-F-000-10-var-2/3-0	0	1	[0,90]	$\frac{2}{3}$	0
M-F-050-10-var-2/3-0	50	1	[0,90]	$\frac{2}{3}$	0
M-F-100-10-var-2/3-0	100	1	[0,90]	$\frac{2}{3}$	0

Table 3.3: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\theta = [0,90]^{\circ}$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ .

The following two plots show  $C_D$  vs.  $\theta$ . In Figure 3.6 colored lines correspond to different values of  $\alpha$  and each line style correspond to different values of  $T_w$ , while in Figure 3.7 colored lines correspond to different  $T_w$  values and each line style corresponds to different values of  $\alpha$ .

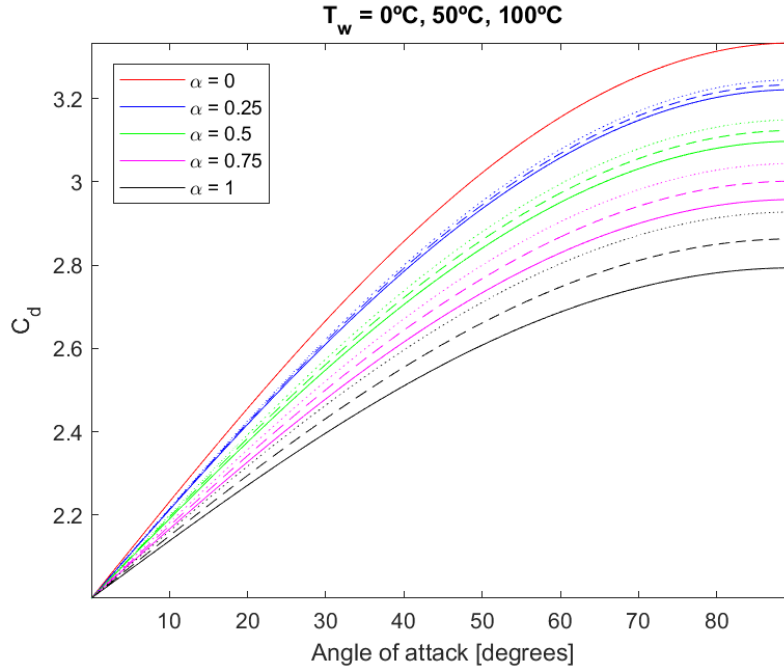


Figure 3.6: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.25 \ 0.5 \ 0.75 \ 1]$ ,  $\theta = [0,90]^{\circ}$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ . Note that solid lines correspond to  $T_w = 0^{\circ}C$ , dashed lines correspond to  $T_w = 50^{\circ}C$  and dotted lines correspond to  $T_w = 100^{\circ}C$ .

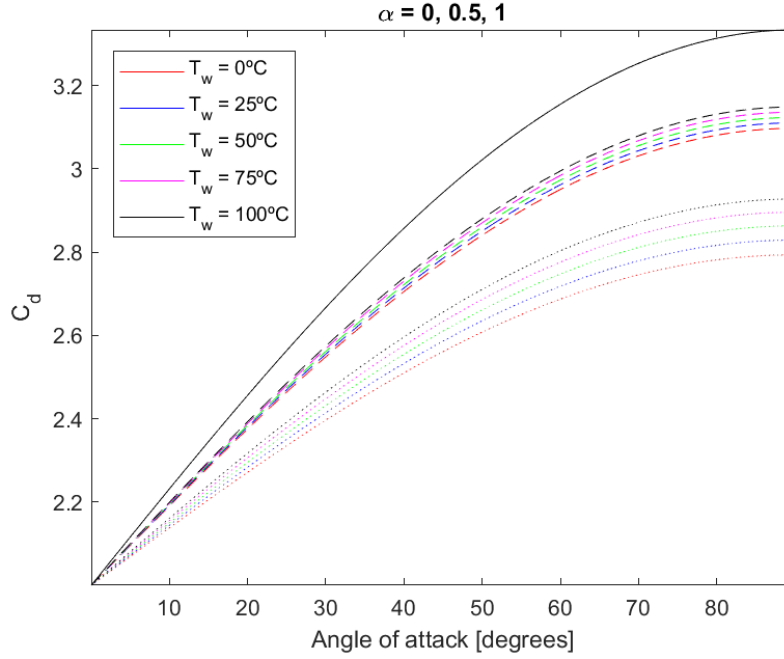


Figure 3.7: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^\circ\text{C}$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\theta = [0, 90]^\circ$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ . Note that solid lines correspond to  $\alpha = 0$ , dashed lines correspond to  $\alpha = 0.5$  and dotted lines correspond to  $\alpha = 1$ .

From both Figure 3.6 and Figure 3.7 it can be seen that the behaviour of  $C_D$  is similar to the specular reflection case, but for diffuse reflection there are no values of  $C_D$  lower than 2 and for  $\theta = 0^\circ$ , the value of  $T_w$  or  $\alpha$  does not cause a variation in  $C_D$ .

### 3.2.2.2. Constant $T_w$ and $\theta$

Table 3.4 shows the structure of the different models studied. As it can be seen from Figure 3.9 and Figure 3.10, variation of  $C_D$  decreases as  $\theta$  increases (note that  $\Delta C_D$  between  $15^\circ$  and  $30^\circ$  is much bigger than between  $75^\circ$  and  $90^\circ$ ). There's a special case when  $\theta = 0^\circ$ , that can be seen in Figure 3.8, that differs a lot with the other values of  $\theta$ . This could be due to the fact that maybe, this model is not taking into account the thermal speed of molecules and, as a result, the behaviour of  $C_D$  for  $\theta = 0^\circ$  should not be taken into account.



Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-var-00-2/3-0	0	[0,1]	0	$\frac{2}{3}$	0	M-F-075-var-45-2/3-0	75	[0,1]	45	$\frac{2}{3}$	0
M-F-025-var-00-2/3-0	25	[0,1]	0	$\frac{2}{3}$	0	M-F-100-var-45-2/3-0	100	[0,1]	45	$\frac{2}{3}$	0
M-F-050-var-00-2/3-0	50	[0,1]	0	$\frac{2}{3}$	0	M-F-000-var-60-2/3-0	0	[0,1]	60	$\frac{2}{3}$	0
M-F-075-var-00-2/3-0	75	[0,1]	0	$\frac{2}{3}$	0	M-F-025-var-60-2/3-0	25	[0,1]	60	$\frac{2}{3}$	0
M-F-100-var-00-2/3-0	100	[0,1]	0	$\frac{2}{3}$	0	M-F-050-var-60-2/3-0	50	[0,1]	60	$\frac{2}{3}$	0
M-F-000-var-15-2/3-0	0	[0,1]	15	$\frac{2}{3}$	0	M-F-075-var-60-2/3-0	75	[0,1]	60	$\frac{2}{3}$	0
M-F-025-var-15-2/3-0	25	[0,1]	15	$\frac{2}{3}$	0	M-F-100-var-60-2/3-0	100	[0,1]	60	$\frac{2}{3}$	0
M-F-050-var-15-2/3-0	50	[0,1]	15	$\frac{2}{3}$	0	M-F-000-var-75-2/3-0	0	[0,1]	75	$\frac{2}{3}$	0
M-F-075-var-15-2/3-0	75	[0,1]	15	$\frac{2}{3}$	0	M-F-025-var-75-2/3-0	25	[0,1]	75	$\frac{2}{3}$	0
M-F-100-var-15-2/3-0	100	[0,1]	15	$\frac{2}{3}$	0	M-F-050-var-75-2/3-0	50	[0,1]	75	$\frac{2}{3}$	0
M-F-000-var-30-2/3-0	0	[0,1]	30	$\frac{2}{3}$	0	M-F-075-var-75-2/3-0	75	[0,1]	75	$\frac{2}{3}$	0
M-F-025-var-30-2/3-0	25	[0,1]	30	$\frac{2}{3}$	0	M-F-100-var-75-2/3-0	100	[0,1]	75	$\frac{2}{3}$	0
M-F-050-var-30-2/3-0	50	[0,1]	30	$\frac{2}{3}$	0	M-F-000-var-90-2/3-0	0	[0,1]	90	$\frac{2}{3}$	0
M-F-075-var-30-2/3-0	75	[0,1]	30	$\frac{2}{3}$	0	M-F-025-var-90-2/3-0	25	[0,1]	90	$\frac{2}{3}$	0
M-F-100-var-30-2/3-0	100	[0,1]	30	$\frac{2}{3}$	0	M-F-050-var-90-2/3-0	50	[0,1]	90	$\frac{2}{3}$	0
M-F-000-var-45-2/3-0	0	[0,1]	45	$\frac{2}{3}$	0	M-F-075-var-90-2/3-0	75	[0,1]	90	$\frac{2}{3}$	0
M-F-025-var-45-2/3-0	25	[0,1]	45	$\frac{2}{3}$	0	M-F-100-var-90-2/3-0	100	[0,1]	90	$\frac{2}{3}$	0
M-F-050-var-45-2/3-0	50	[0,1]	45	$\frac{2}{3}$	0						

Table 3.4: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^{\circ}C$ ,  $\alpha = [0,1]$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ .

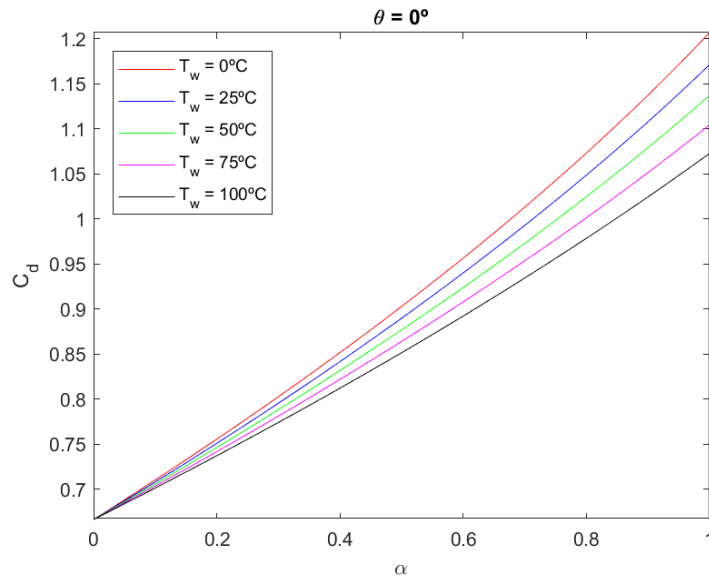


Figure 3.8: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0 \ 25 \ 50 \ 75 \ 100]^{\circ}C$ ,  $\alpha = [0,1]$ ,  $\theta = 0^{\circ}$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ .

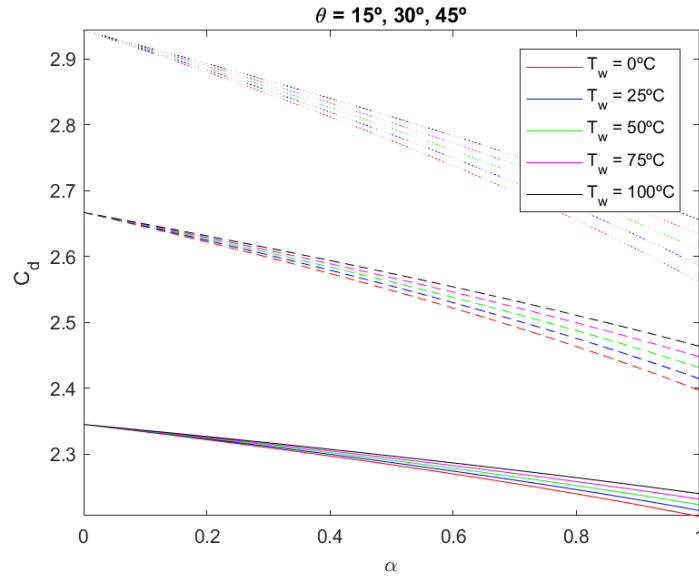


Figure 3.9: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0\ 25\ 50\ 75\ 100]^\circ\text{C}$ ,  $\alpha = [0,1]$ ,  $\theta = [15\ 30\ 45]^\circ$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ . Note that solid lines correspond to  $\theta = 15^\circ$ , dashed lines correspond to  $\theta = 30^\circ$  and dotted lines correspond to  $\theta = 45^\circ$ .

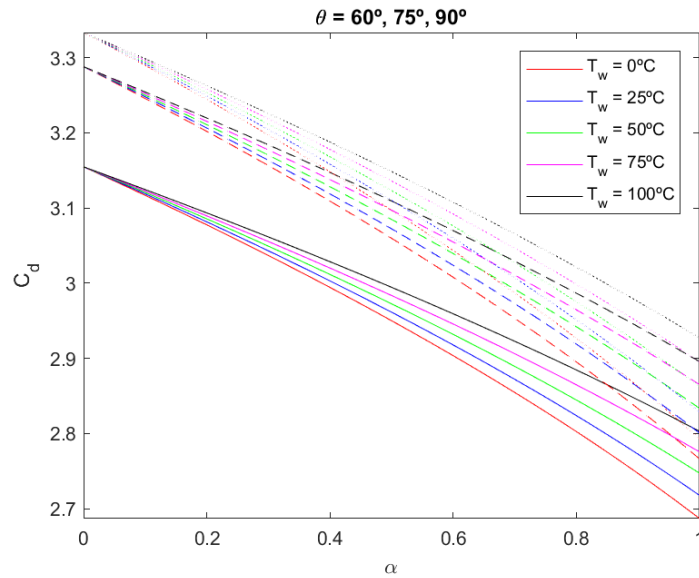


Figure 3.10: Analyzed Schamberg models for flat plate and diffuse reflection.  $T_w = [0\ 25\ 50\ 75\ 100]^\circ\text{C}$ ,  $\alpha = [0,1]$ ,  $\theta = [60\ 75\ 90]^\circ$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ . Note that solid lines correspond to  $\theta = 60^\circ$ , dashed lines correspond to  $\theta = 75^\circ$  and dotted lines correspond to  $\theta = 90^\circ$ .

### 3.2.3. Specular - Diffuse reflection

The study now will be focused on the transition from specular to diffuse reflection, so  $\phi$  will vary from 1 to  $\frac{2}{3}$  and  $\frac{1}{v}$  will vary from 1 to 0. For easier identification, it will be structured in three different sections depending on the value of  $\alpha$ . Each section will include

a parameters table and two plots, and they will be discussed all together at the end of Section 3.2.3.

### 3.2.3.1. $\alpha = 0$

Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-00-00-var-var	0	0	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-00-45-var-var	100	0	45	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-00-00-var-var	50	0	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-00-60-var-var	0	0	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-00-00-var-var	100	0	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-00-60-var-var	50	0	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-00-15-var-var	0	0	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-00-60-var-var	100	0	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-00-15-var-var	50	0	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-00-75-var-var	0	0	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-00-15-var-var	100	0	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-00-75-var-var	50	0	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-00-30-var-var	0	0	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-00-75-var-var	100	0	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-00-30-var-var	50	0	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-00-90-var-var	0	0	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-00-30-var-var	100	0	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-00-90-var-var	50	0	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-00-45-var-var	0	0	45	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-00-90-var-var	100	0	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-00-45-var-var	50	0	45	$[1, \frac{2}{3}]$	$[1, 0]$						

Table 3.5: Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = 0$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ .

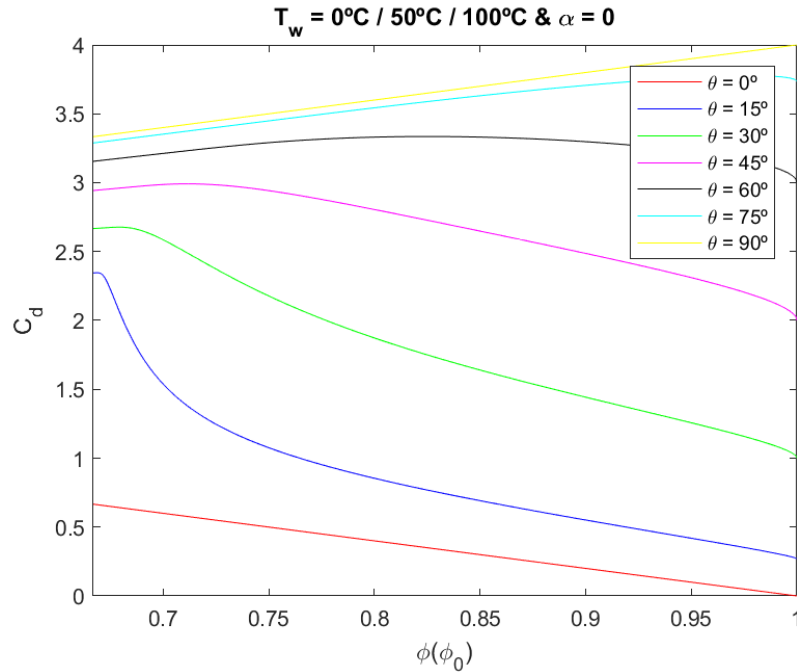


Figure 3.11:  $(C_D \text{ vs. } \phi(\phi_0))$  - Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = 0$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ .

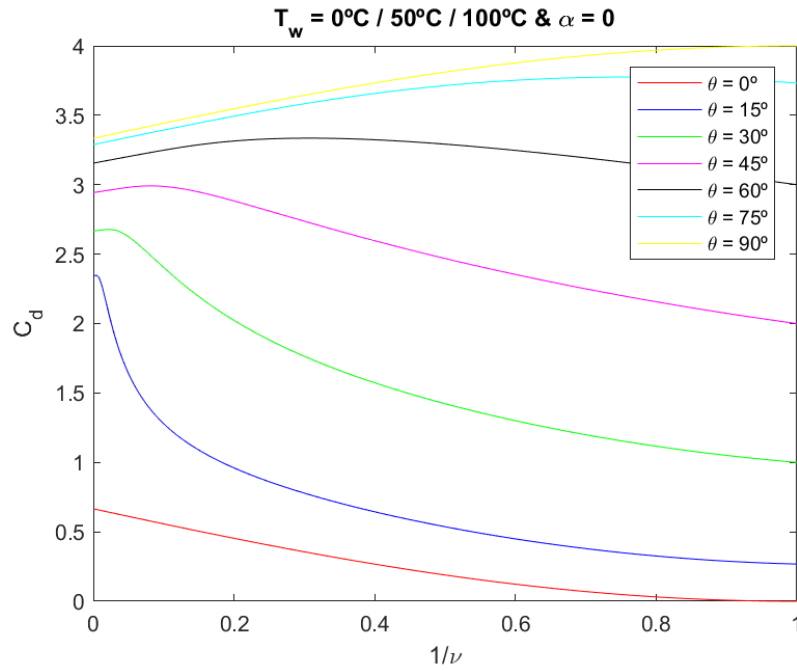


Figure 3.12: ( $C_D$  vs.  $\frac{1}{v}$ ) - Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$ ,  $\alpha = 0$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ .

### 3.2.3.2. $\alpha = 0.5$

Model	$T_w(^{\circ}\text{C})$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}\text{C})$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-05-00-var-var	0	0.5	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-05-45-var-var	100	0.5	45	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-05-00-var-var	50	0.5	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-05-60-var-var	0	0.5	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-05-00-var-var	100	0.5	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-05-60-var-var	50	0.5	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-05-15-var-var	0	0.5	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-05-60-var-var	100	0.5	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-05-15-var-var	50	0.5	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-05-75-var-var	0	0.5	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-05-15-var-var	100	0.5	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-05-75-var-var	50	0.5	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-05-30-var-var	0	0.5	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-05-75-var-var	100	0.5	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-05-30-var-var	50	0.5	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-05-90-var-var	0	0.5	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-05-30-var-var	100	0.5	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-05-90-var-var	50	0.5	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-05-45-var-var	0	0.5	45	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-05-90-var-var	100	0.5	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-05-45-var-var	50	0.5	45	$[1, \frac{2}{3}]$	$[1, 0]$						

Table 3.6: Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$ ,  $\alpha = 0.5$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ .

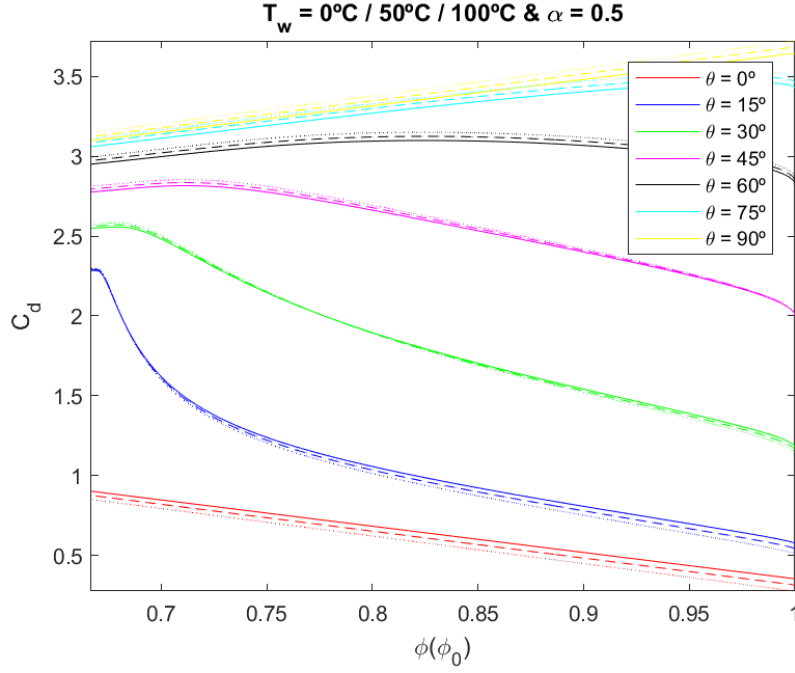


Figure 3.13: ( $C_D$  vs.  $\phi(\phi_0)$ ) - Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$ ,  $\alpha = 0.5$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{\nu} = [1, 0]$ . Note that solid lines correspond to  $T_w = 0^\circ\text{C}$ , dashed lines correspond to  $T_w = 50^\circ\text{C}$  and dotted lines correspond to  $T_w = 100^\circ\text{C}$ .

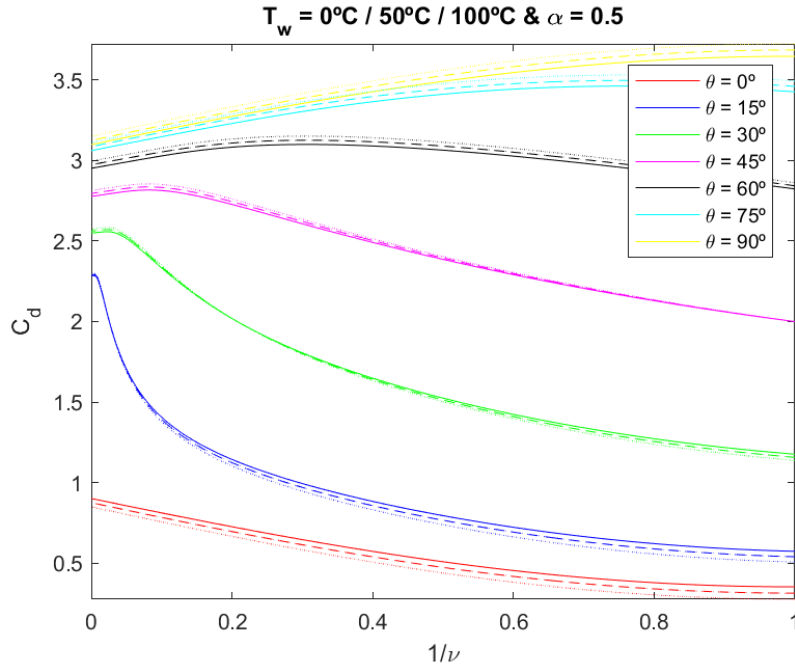


Figure 3.14: ( $C_D$  vs.  $\frac{1}{\nu}$ ) - Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$ ,  $\alpha = 0.5$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{\nu} = [1, 0]$ . Note that solid lines correspond to  $T_w = 0^\circ\text{C}$ , dashed lines correspond to  $T_w = 50^\circ\text{C}$  and dotted lines correspond to  $T_w = 100^\circ\text{C}$ .

3.2.3.3.  $\alpha = 1$ 

Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}C)$	$\alpha$	$\theta(^{\circ})$	$\phi$	$\frac{1}{v}$
M-F-000-10-00-var-var	0	1	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-10-45-var-var	100	1	45	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-10-00-var-var	50	1	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-10-60-var-var	0	1	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-10-00-var-var	100	1	0	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-10-60-var-var	50	1	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-10-15-var-var	0	1	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-10-60-var-var	100	1	60	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-10-15-var-var	50	1	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-10-75-var-var	0	1	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-10-15-var-var	100	1	15	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-10-75-var-var	50	1	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-10-30-var-var	0	1	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-10-75-var-var	100	1	75	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-10-30-var-var	50	1	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-000-10-90-var-var	0	1	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-100-10-30-var-var	100	1	30	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-050-10-90-var-var	50	1	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-000-10-45-var-var	0	1	45	$[1, \frac{2}{3}]$	$[1, 0]$	M-F-100-10-90-var-var	100	1	90	$[1, \frac{2}{3}]$	$[1, 0]$
M-F-050-10-45-var-var	50	1	45	$[1, \frac{2}{3}]$	$[1, 0]$						

Table 3.7: Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = 1$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ .

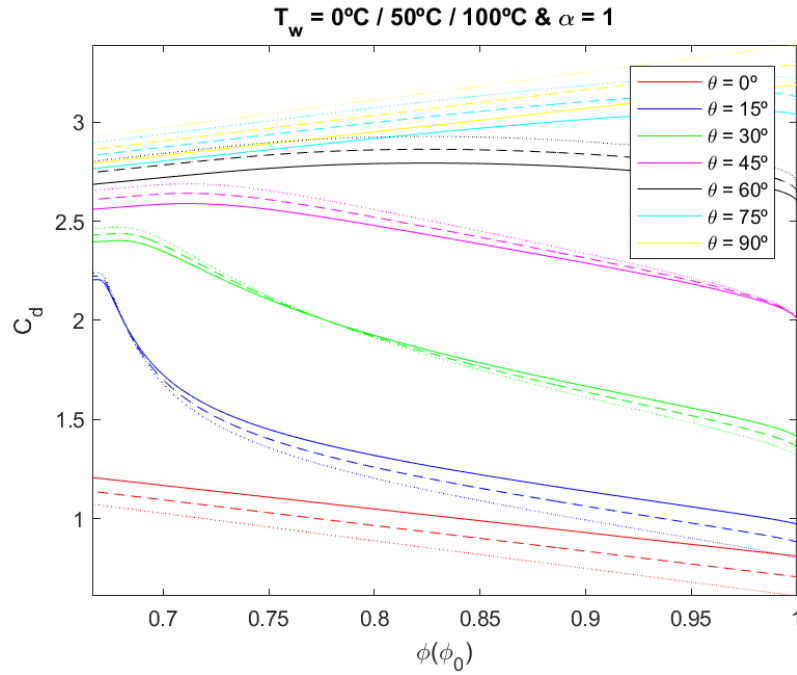


Figure 3.15: ( $C_D$  vs.  $\phi(\phi_0)$ ) - Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = 1$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^{\circ}$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ . Note that solid lines correspond to  $T_w = 0^{\circ}C$ , dashed lines correspond to  $T_w = 50^{\circ}C$  and dotted lines correspond to  $T_w = 100^{\circ}C$ .

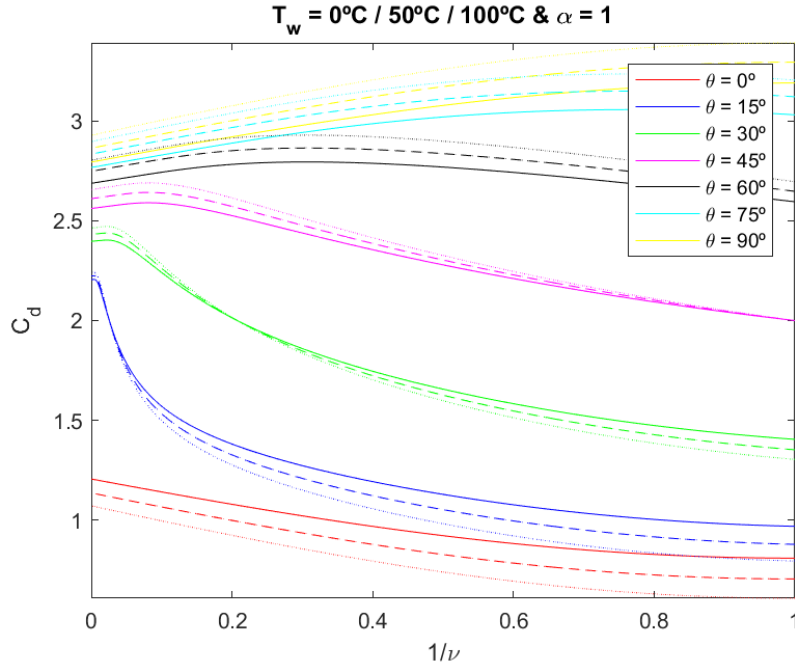


Figure 3.16: ( $C_D$  vs.  $\frac{1}{v}$ ) - Analyzed Schamberg models for flat plate and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$ ,  $\alpha = 1$ ,  $\theta = [0 \ 15 \ 30 \ 45 \ 60 \ 75 \ 90]^\circ$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ . Note that solid lines correspond to  $T_w = 0^\circ\text{C}$ , dashed lines correspond to  $T_w = 50^\circ\text{C}$  and dotted lines correspond to  $T_w = 100^\circ\text{C}$ .

#### 3.2.3.4. Comparison

In the previous figures it can be seen that the relation between  $\alpha$  and  $T_w$  is the same either for specular or diffuse reflection. As  $\alpha$  increases, different values of  $T_w$  cause more variation in  $C_D$ . Note that  $\Delta C_D$  caused by  $T_w$  is greater for extreme values of  $\theta$  than for intermediate ones, while  $\Delta C_D$  between different values of  $\theta$  becomes smaller for extreme values.

### 3.3. Sphere

All models studied in this section have the following nomenclature for an easier identification:  $M - S - T_w T_w T_w - \alpha \alpha - \phi - \frac{1}{v}$

#### 3.3.1. Completely specular reflection

This study is going to be focused on specular reflection, which implies  $\phi = 1$  and  $\frac{1}{v} = 1$ . It will be divided into two cases;  $T_w$  is going to be kept constant for the first case and  $\alpha$  for the second case.

### 3.3.1.1. Constant $T_w$ and $\alpha$

This case includes two different combinations. One in which  $T_w = [0 \ 50 \ 100]^\circ\text{C}$  and  $\alpha = [0, 1]$ , and the other one in which  $T_w = [0, 100]^\circ\text{C}$  and  $\alpha = [0 \ 0.5 \ 1]$ . Table 3.8 shows the values of both cases.

Model	$T_w(^{\circ}\text{C})$	$\alpha$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}\text{C})$	$\alpha$	$\phi$	$\frac{1}{v}$
M-S-000-var-1-1	0	[0,1]	1	1	M-S-var-00-1-1	[0,100]	0	1	1
M-S-050-var-1-1	50	[0,1]	1	1	M-S-var-05-1-1	[0,100]	0.5	1	1
M-S-100-var-1-1	100	[0,1]	1	1	M-S-var-10-1-1	[0,100]	1	1	1

Table 3.8: Analyzed Schamberg models for sphere and specular reflection.  $T_w = [0, 100]^\circ\text{C}$ ,  $\alpha = [0, 1]$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ .

Specular reflection cases using Schamberg's model yield constant  $C_D$  values exactly equal to 2 (see Figure B.3 and Figure B.4 from Appendix B). No dependence either on wall temperature or accommodation coefficient is detected, which is consistent with equation for  $C_D$  in spherical case.

## 3.3.2. Completely diffuse reflection

In this section, the procedure to develop is computing  $C_D$  for diffuse reflection on a sphere, which involves  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ . The same structure of the previous section will be followed here.

### 3.3.2.1. Constant $T_w$ and $\alpha$

This case will be performed using the same procedure as for specular reflection. Table 3.9 shows the values of each parameter

Model	$T_w(^{\circ}\text{C})$	$\alpha$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}\text{C})$	$\alpha$	$\phi$	$\frac{1}{v}$
M-S-000-var-2/3-0	0	[0,1]	$\frac{2}{3}$	0	M-S-var-00-2/3-0	[0,100]	0	$\frac{2}{3}$	0
M-S-050-var-2/3-0	50	[0,1]	$\frac{2}{3}$	0	M-S-var-05-2/3-0	[0,100]	0.5	$\frac{2}{3}$	0
M-S-100-var-2/3-0	100	[0,1]	$\frac{2}{3}$	0	M-S-var-10-2/3-0	[0,100]	1	$\frac{2}{3}$	0

Table 3.9: Analyzed Schamberg models for sphere and diffuse reflection.  $T_w = [0, 100]^\circ\text{C}$ ,  $\alpha = [0, 1]$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ .



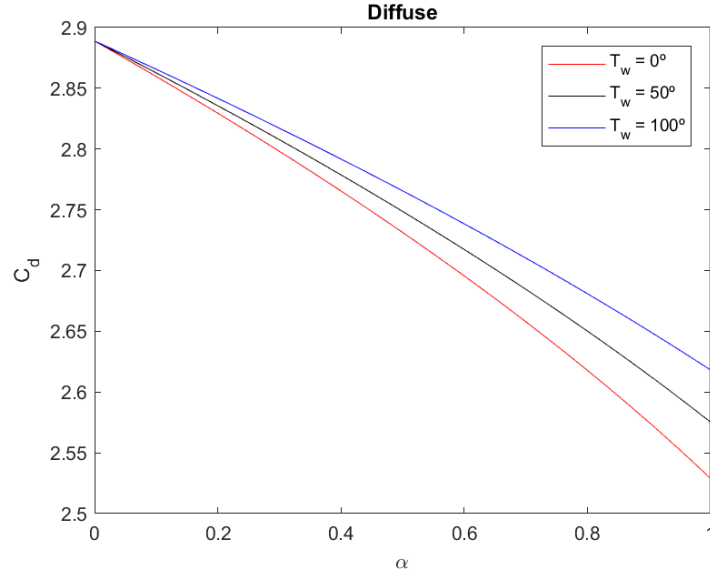


Figure 3.17: Analyzed Schamberg models for sphere and diffuse reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$ ,  $\alpha = [0, 1]$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ .

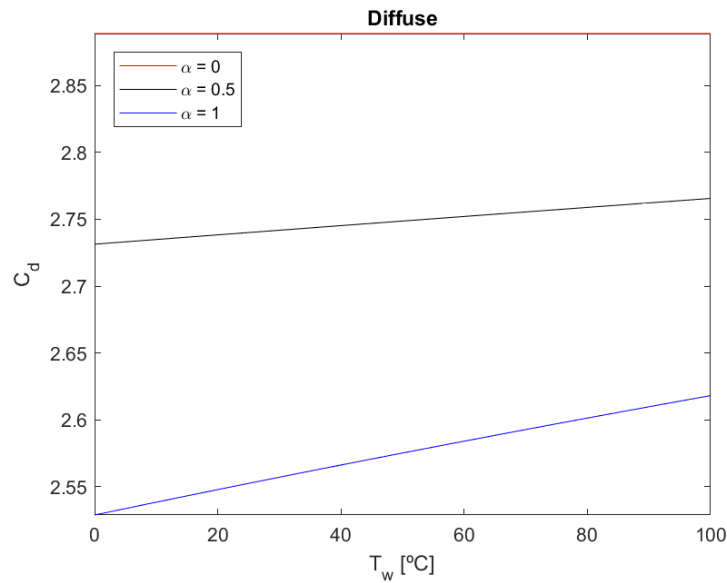


Figure 3.18: Analyzed Schamberg models for sphere and diffuse reflection.  $T_w = [0, 100]^\circ\text{C}$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\phi = \frac{2}{3}$  and  $\frac{1}{v} = 0$ .

Regarding both Figure 3.17 and Figure 3.18, it can be seen that as  $\alpha$  increases, the variation in  $C_D$  due to  $T_w$  increases too.

### 3.3.3. Specular - Diffuse reflection

Finally, the transition from specular to diffuse reflection is going to be studied. Now  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ . Table 3.9 shows the values of the parameters for each model.

Model	$T_w(^{\circ}C)$	$\alpha$	$\phi$	$\frac{1}{v}$	Model	$T_w(^{\circ}C)$	$\alpha$	$\phi$	$\frac{1}{v}$
M-S-000-00-var-var	0	0	$[\frac{2}{3}, 1]$	$[0, 1]$	M-S-100-05-var-var	100	0.5	$[\frac{2}{3}, 1]$	$[0, 1]$
M-S-050-00-var-var	50	0	$[\frac{2}{3}, 1]$	$[0, 1]$	M-S-000-10-var-var	0	1	$[\frac{2}{3}, 1]$	$[0, 1]$
M-S-100-00-var-var	100	0	$[\frac{2}{3}, 1]$	$[0, 1]$	M-S-050-10-var-var	50	1	$[\frac{2}{3}, 1]$	$[0, 1]$
M-S-000-05-var-var	0	0.5	$[\frac{2}{3}, 1]$	$[0, 1]$	M-S-100-10-var-var	100	1	$[\frac{2}{3}, 1]$	$[0, 1]$
M-S-050-05-var-var	50	0.5	$[\frac{2}{3}, 1]$	$[0, 1]$					

Table 3.10: Analyzed Schamberg models for sphere and transition from specular to diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ .

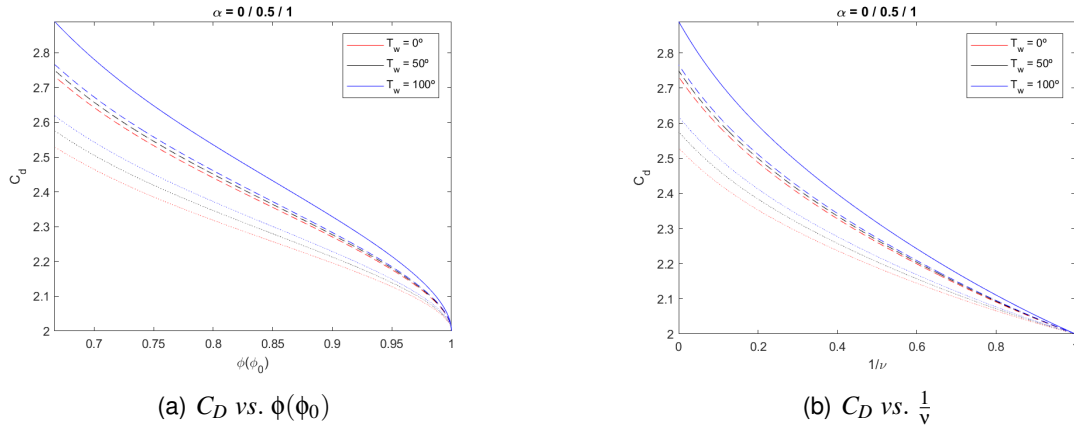


Figure 3.19: Schamberg models for sphere. Specular-Diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ . Note that solid lines correspond to  $\alpha = 0$ , dashed lines to  $\alpha = 0.5$  and dotted lines to  $\alpha = 1$ .

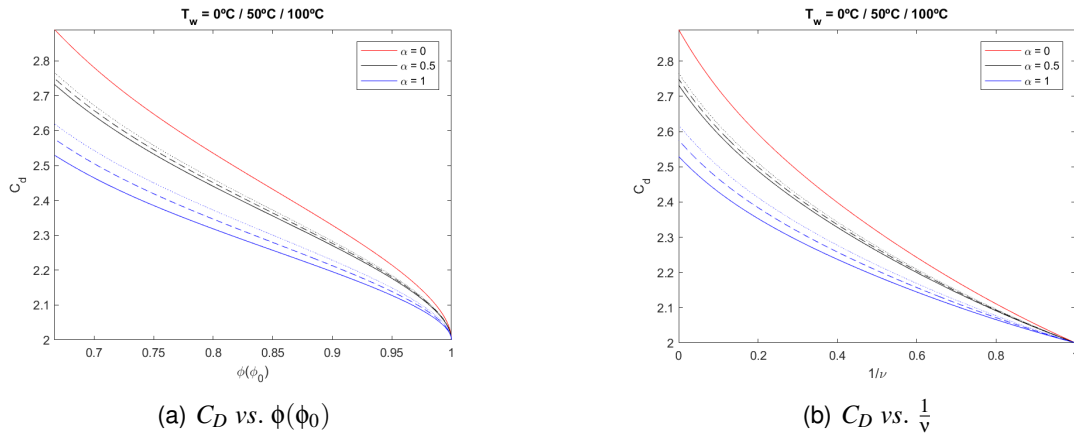


Figure 3.20: Schamberg models for sphere. Specular-Diffuse reflection.  $T_w = [0 \ 50 \ 100]^{\circ}C$ ,  $\alpha = [0 \ 0.5 \ 1]$ ,  $\phi = [1, \frac{2}{3}]$  and  $\frac{1}{v} = [1, 0]$ . Note that solid lines correspond to  $T_w = 0^{\circ}C$ , dashed lines to  $T_w = 50^{\circ}C$  and dotted lines to  $T_w = 100^{\circ}C$ .

As it can be seen in Figure 3.19 and Figure 3.20, the  $\Delta C_D$  caused by  $T_w$  and  $\alpha$  variations is greater as the reflection becomes more diffuse. It can also be observed that  $\Delta C_D$  between different values of  $T_w$  is larger for high values of  $\alpha$ .

# CHAPTER 4. SENTMAN MODEL FOR DRAG COEFFICIENTS

In this chapter, we are going to expose Sentman model for drag coefficients that could fit when computing drag coefficient for diffuse reflection.

## 4.1. Model Description

The accommodation coefficient in this model is treated as an empirical parameter effective over the entire satellite surface. Momentum is highly sensitive to accommodation, as postulated in (12), and is strongly related to atomic oxygen absorption of the satellite's surface. So the variation of  $C_D$  with altitude is driven by changes in atomic oxygen absorption which indicates that density changes also produce variations in  $C_D$ . The accommodation coefficient is defined by (3) as:

$$\alpha = \frac{T_{k,in} - T_{k,out}}{T_{k,in} - T_w} \quad (4.1)$$

where  $T_{k,in}$  is the kinetic temperature of incoming molecules,  $T_{k,out}$  the distribution temperature of the reflected molecules and  $T_w$  the surface temperature and can be written as:

$$T_{k,in} = \frac{mV_r^2}{3k_b} \quad (4.2)$$

$$T_{k,out} = \frac{m}{3k_b} V_r^2 (1 - \alpha) + \alpha T_w \quad (4.3)$$

There are typically two types of reflection ascribed to the gas-surface interactions in low-earth orbit, as can be seen in Figure 4.1, that's why this combination of models intends to be useful for computing  $C_D$  for both types of reflection.

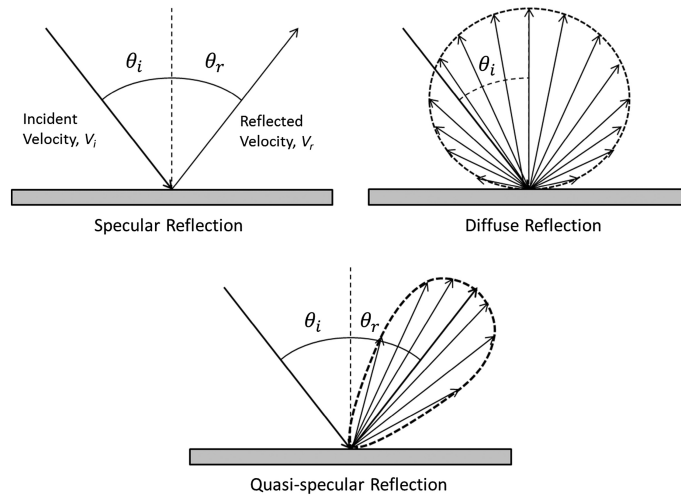


Figure 4.1: Different cases of particle reflection. Credit (4).

The majority of interactions below 500 km are diffuse in character with a scattering kernel that approximates the cosine (12). This happens when the velocities of reflected molecules are centered around the surface normal vector in a cosine distribution, as shown in the left panel of Figure 4.1. In the second model, the molecules are reflected in a narrow lobe centered around the specular direction.

A successful technique when computing  $C_D$  is to adapt the diffuse model developed by Sentman (6) and the quasi-specular model developed by Schamberg (5). Sentman's model adapts the diffuse energy distribution with a variable accommodation coefficient, which is a diffuse reflection with incomplete accommodation, while Schamberg's model reflects the majority of molecules around a direction such that the angle of reflection  $\theta_r$  is smaller or equal than  $\theta_i$ .

Once a set of incident and reflected velocity distribution is set, we can compute the drag coefficient for diffuse reflection on a flat plate and a sphere by using the following equations:

When computing the drag coefficients using this model, we will use Schamberg's approximation (5) for the temperature quotient defined as:

$$\sqrt{\frac{T_{k,out}}{T_a}} = \sqrt{1 - \alpha} \quad (4.4)$$

So then

$$C_{D,plate} = \frac{2}{s\sqrt{\pi}} \exp[-s^2 \sin^2(\theta_i)] + \frac{\sin(\theta_i)}{s^2} (1 + 2s^2) \operatorname{erf}[s \sin(\theta_i)] + \frac{\sqrt{\pi}}{s} \sin^2(\theta_i) \sqrt{1 - \alpha} \quad (4.5)$$

$$C_{D,sphere,diff} = \frac{2s^2 + 1}{\sqrt{\pi}s^3} \exp(-s^2) + \frac{4s^4 + 4s^2 - 1}{2s^4} \operatorname{erf}(s) + \frac{2\sqrt{\pi}}{3s} \sqrt{1 - \alpha} \quad (4.6)$$

Now, using Eqs.4.5 and 4.6 we can compute the drag coefficient from a flat plate and a sphere. In the flat plate case we are going to consider the variation of the angle of attack,  $\theta_i$ , and different values for the accommodation coefficient,  $\alpha$ , [0 0.25 0.5 0.75 1].

It can be seen from Figure 4.2 that  $C_D$  goes from values close to zero for small angles of attack, where  $\alpha$  does not have any influence, to values around 2 for high angles of attack, and then  $\alpha$  plays an important role causing  $\Delta C_D$ . From Figure 4.3, we can say that the variation of the accommodation coefficient does not cause a big change in  $C_D$  for a sphere.

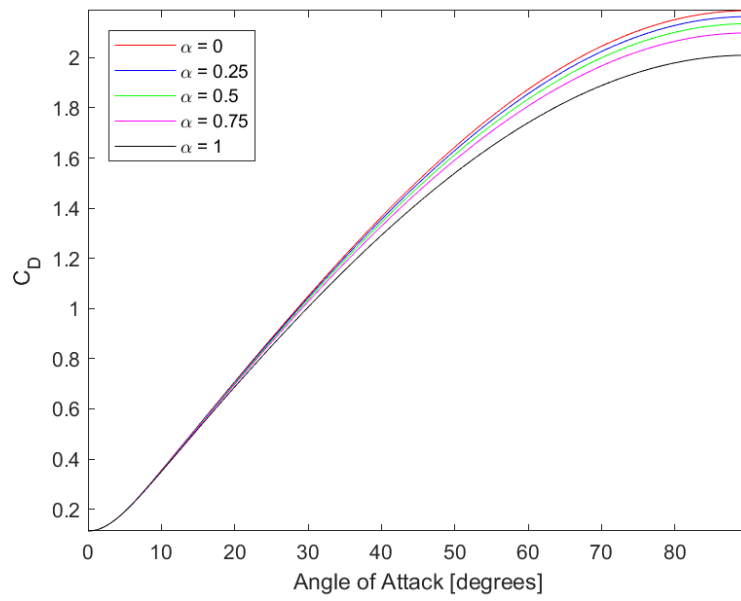


Figure 4.2: Analyzed Sentman model for flat plate and diffuse reflection.  $\theta = [0, 90]^\circ$  and  $\alpha = [0 \ 0.25 \ 0.5 \ 0.75 \ 1]$ .

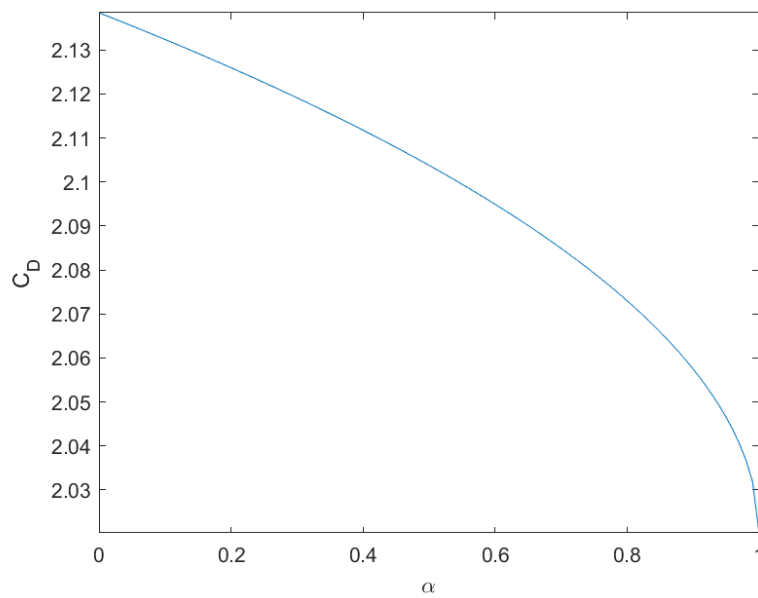


Figure 4.3: Analyzed Sentman model for sphere and diffuse reflection.  $\alpha = [0, 1]$ .



## CHAPTER 5. CONCLUSIONS

A proper assessment of satellite  $C_D$  values is critical for the understanding of satellite dynamics and, in particular, for a more accurate determination of atmospheric density values. We have calculated and analyzed the  $C_D$  values in free molecular flow and hyper-thermal velocity conditions, both for flat plate and spherical symmetries, using the models by Schaaf and Chambre, (3), SC58, and by Schamberg (5), S58.

In spite of being relatively old, these models (or slightly modified versions, like (6)) are still used nowadays, and remain important touchstones when checking the results of more sophisticated recent numerical models. The determination of the parameters on which they depend (see Chapters 2 and 3) can be achieved by performing experiments, both on Earth and in space (13). Improved determinations have been made in the last decades (14), and their influence on  $C_D$ , as well as the dependence of the latter coefficient on altitude and solar activity has been studied.

We have analyzed the limitations of both models, whose performance varies depending on the specific problem conditions. Frequent (and most successful) applications actually combine both models. S58 is usually applied in cases of quasi-specular reflection (see Figure 4.1), for moderate-to-high angles of attack. Its main advantage lays in the fact that it can consider re-emission of particles in a certain angular range. However, as we have seen, this model yields unrealistically low values at low grazing angles probably because it uses a rather simplistic approximation for the incident stream of particles. SC58 is applied in cases of diffuse reflection. It can describe more accurately momentum transfer at moderate to low grazing angles. In practice, we express the total  $C_D$  as:

$$C_D = f C_{D,spec} + (1 - f) C_{D,diff} \quad (5.1)$$

where  $f$  is a parameter to be determined in each case.

The main conclusions derived from our calculations are:

- i) Comparing both Schaaf and Chambre and Schamberg models, we can deduce that for the specular case, we obtain similar values of  $C_D$  when varying the angle of attack.

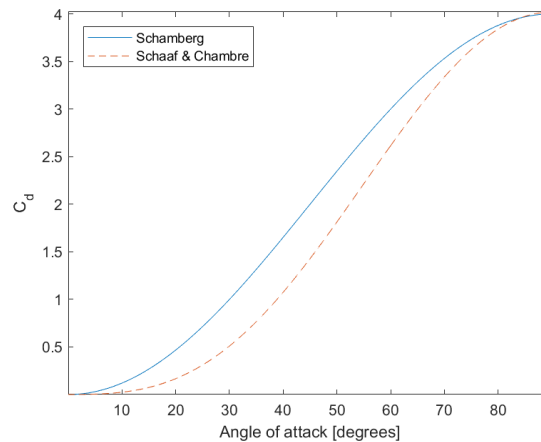


Figure 5.1: Comparison between Schaaf and Chambre (3) and Schamberg (5) for flat plate and specular reflection.

Figure 5.1 shows that both models return very similar values for a flat-plate when specular reflection occurs. For spherical geometries the result is also analogous and both models return a constant value. For SC58 we obtain  $C_D = 2.02$ , and for S58 we obtain  $C_D = 2$ . These constant values are due to the fact that perfect spherical symmetry the effect of the angle of attack is irrelevant. Neither there is a variation with  $T_w$ , because there no accommodation for these cases.

- ii) If we analyze the diffuse cases, we realize that both SC58 and S58 models return different values for some particular conditions (see Figure 5.2).

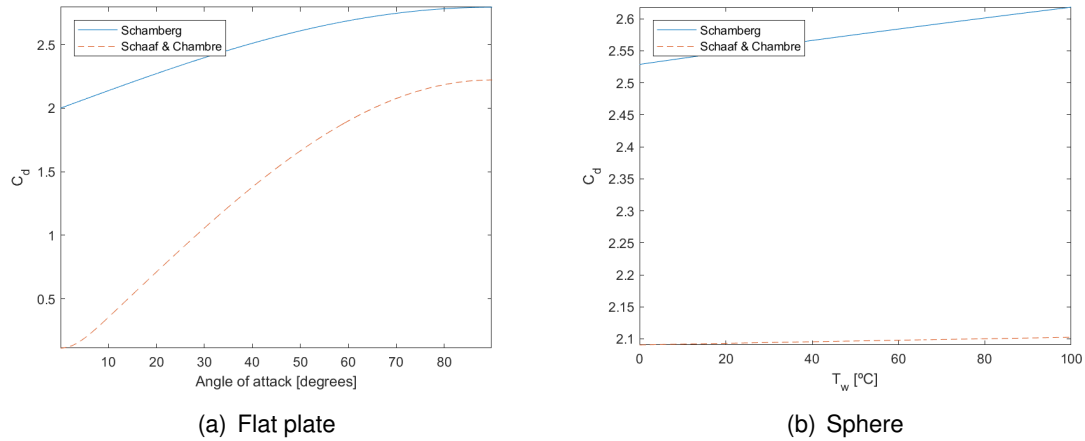


Figure 5.2: Comparison between Schaaf and Chambre (3) and Schamberg (5) for (a) Flat plate and (b) Sphere and diffuse reflection.

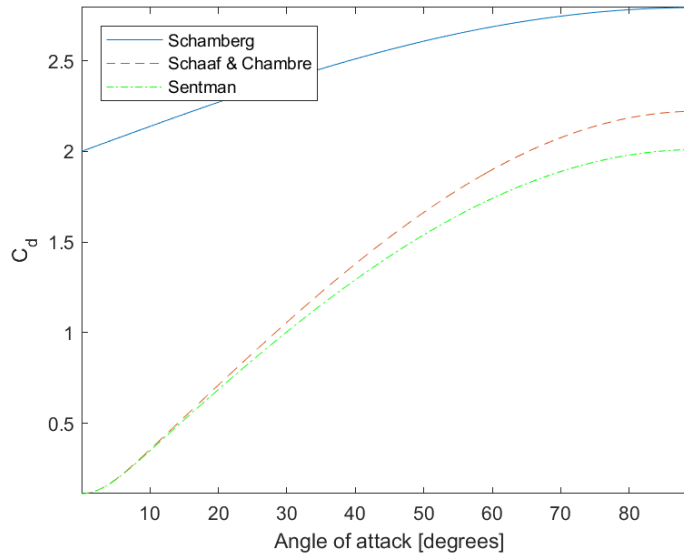


Figure 5.3: Comparison between Schaaf and Chambre (3), Schamberg (5) and Sentman (6) for flat plate and diffuse reflection.

Left panel of Figure 5.2 shows that there is a huge difference between both models, specially for low angles of attack, where Schamberg's values are about 20 times



greater than the ones obtained with Schaaf and Chambre model. Schamberg's values do not reproduce empirical values of  $C_D$  obtained from real satellites, for example CubeSats.

Actually, it is well known that S58 model is not accurate enough for diffuse reflection (and low grazing angles). Actually, an alternative popular  $C_D$  model, Sentman's model (6) yields values for diffuse reflection cases very similar to the ones by SC58, and widely different from the results with S58 (see Figure 5.3). A combination of both models (S58 for quasi-specular reflection and SC58 or Sentman's for diffuse reflection and low grazing angles) can handle almost any physical situation (14).

- iii) We have checked the crucial role of the angle of attack when determining the value of  $C_D$  and flat-plate geometries, as it can cause variations of around 30 times between  $0^\circ$  and  $90^\circ$ . This is particularly important when specular reflection occurs. Given the problems for controlling satellite' attitude during the re-entry stage, any experiment designed for the assessment of  $C_D$  would be greatly simplified by using spherical symmetry.
- iv) Regarding the accommodation coefficient,  $\alpha$ , it is important to recall that it is a determining factor for  $C_D$  calculations both for Schamberg and Sentman models. Schaaf and Chambre equations for  $C_D$  calculations do not include the parameter  $\alpha$ , but it is assumed to be  $\alpha = 0$  for specular reflection and  $\alpha = 1$  for diffuse reflection. Orbital velocity, as well as surface temperature, are the determining factor for different  $\alpha$  values.
- v) When analyzing the results from the different models, we can see that the surface temperature,  $T_w$ , must be taken into account for drag modelling. Variations in  $T_w$  do not cause wide  $C_D$  variations by themselves, as far as  $\alpha$  is constant. However it is important to remind that in real cases,  $T_w$  changes imply variations in  $\alpha$ , which themselves causes high variations in  $C_D$ .
- vi) Finally, it should be recalled that flat-plate geometry is an oversimplification of real satellite geometries, such as CubeSats. In particular, considering a single angle of attack is not realistic for most satellite orientations. We have seen that spherical satellites are more useful for the assessment of  $C_D$  (and ultimately, of atmospheric density). However, they pose the technical problem of the adaptation to standard launchers, which are rather designed for the launch of CubeSats. Probably, the solution to allow both interfacing with launchers and good orbital performance would be to design a dispenser that, once in orbit, could open and release one (or several) spherical satellite(s) as the final geometry that is going to orbit around the Earth and, eventually, obtain experimental data.



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Escola d'Enginyeria de Telecomunicació  
i Aeroespacial de Castelldefels

# APPENDICES

**TFG TITLE:** Determination of drag coefficients in a free molecular flow

**DEGREE:** Grau en Enginyeria d'Aeronavegació

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**DATE:** July 17, 2018



# APPENDIX A. TYPICAL THERMAL ACCOMMODATION COEFFICIENT VALUES FOR AIR

Surface	$\alpha$
Flat lacquer on bronze	0.88 – 0.89
Polished bronze	0.91 – 0.94
Machined bronze	0.89 – 0.93
Etched bronze	0.93 – 0.95
Polished cast iron	0.87 – 0.93
Machined cast iron	0.87 – 0.88
Etched cast iron	0.89 – 0.96
Polished aluminum	0.89 – 0.95
Machined aluminum	0.95 – 0.97
Etched aluminum	0.89 – 0.97

Table A.1: Thermal accommodation coefficient  $\alpha$  for air depending on the surface.





## APPENDIX B. EXTRA FIGURES

This appendix contains some extra figures related to the models studied that are intended to show some exceptional cases mentioned during the project.

### B.1. Schaaf and Chambre model

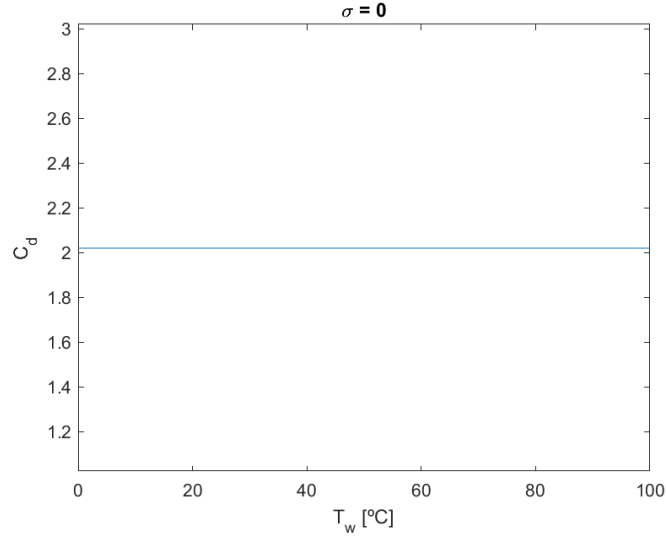


Figure B.1: Analyzed Schaaf and Chambre models for sphere and specular reflection.  $\sigma = 0$  and  $T_w = [0, 100]^\circ\text{C}$ .

### B.2. Schamberg model

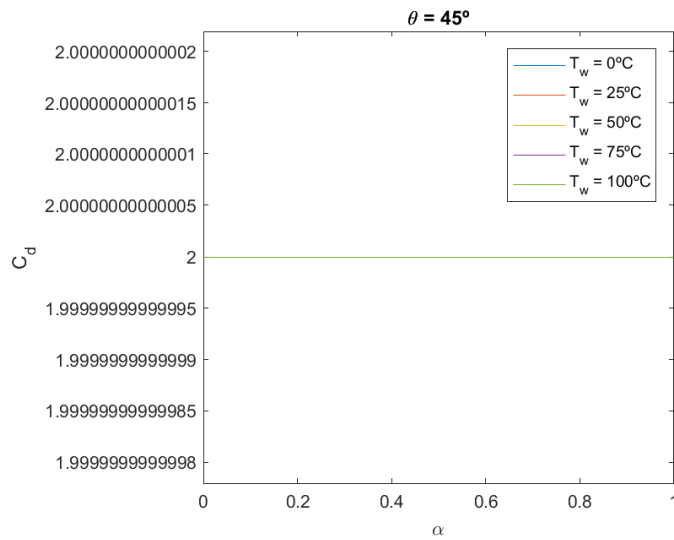


Figure B.2: Analyzed Schamberg models for flat plate and specular reflection.  $T_w = [0, 25, 50, 75, 100]^\circ\text{C}$ ,  $\alpha = [0, 1]$ ,  $\theta = 45^\circ$ ,  $\phi = 1$  and  $\frac{1}{v} = 1$ .

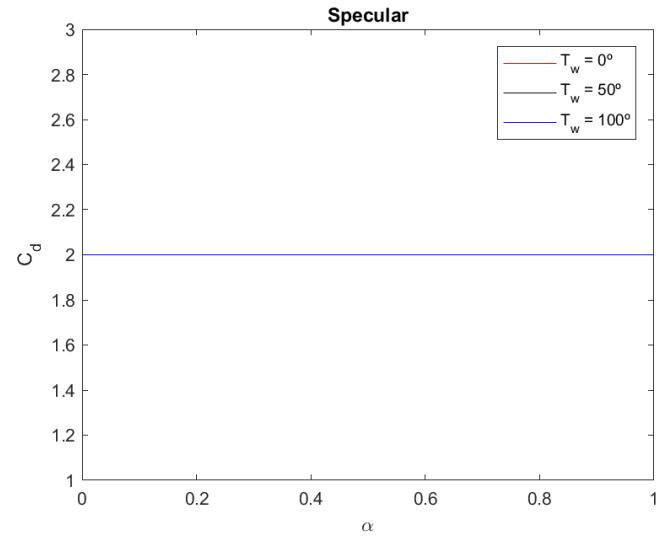


Figure B.3: Analyzed Schamberg models for sphere and specular reflection.  $T_w = [0 \ 50 \ 100]^\circ\text{C}$  and  $\alpha = [0, 1]$

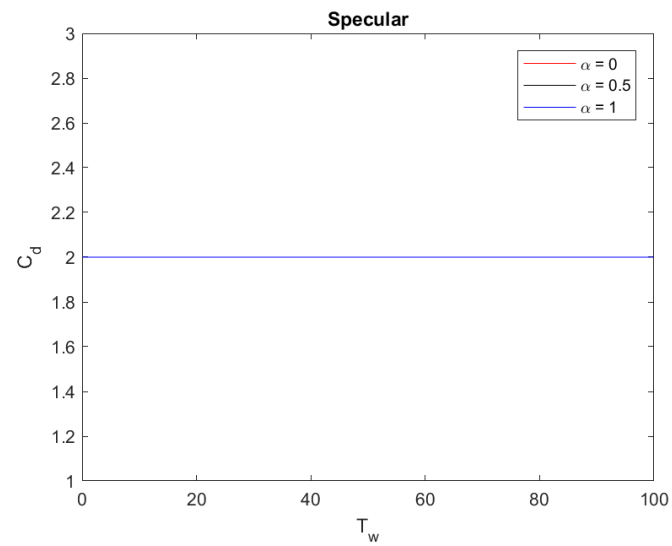


Figure B.4: Analyzed Schamberg models for sphere and specular reflection.  $T_w = [0, 100]^\circ\text{C}$  and  $\alpha = [0 \ 0.5 \ 1]$

# APPENDIX C. MATLAB CODE

## C.1. Schaaf and Chambre model

### C.1.1. Flat plate

#### C.1.1.1. Main menu

```
1 clear all
2 close all
3
4 list = { 'Specular', 'Diffuse', 'Variable' };
5 [indx,tf] = listdlg('PromptString','Select an option:', '
    SelectionMode','single', 'ListSize',[400,200], 'ListString', list
    );
6
7 if(indx == 1)
8     list = { 'Constant Temperature', 'Constant angle of attack' };
9     [indx,tf] = listdlg('PromptString','Select an option:', '
        SelectionMode','single', 'ListSize',[400,200], 'ListString',
        list);
10    if(indx == 1)
11        T_w = [0 25 50 75 100] + 273;
12        sigma=0;
13        theta = 0.00001:0.00001:(pi/2);
14        model=1;
15        [C_d_SchaafChambre_flat] = SchaafChambre_plots_flat(T_w,
            sigma, theta, model);
16        figure()
17        plot(theta*180/pi, C_d_SchaafChambre_flat(:,1))
18        xlabel('\theta [degrees]');
19        ylabel('C_d');
20        title('\sigma = 0');
21        axis tight
22        saveas(gcf, 'Model_1_1.png')
23    elseif(indx == 2)
24        T_w = (0:1:100)+273;
25        theta = 0:((pi/2)/(length(T_w)-1)):(pi/2);
26        sigma=0;
27        model=2;
28        [C_d_SchaafChambre_flat] = SchaafChambre_plots_flat(T_w,
            sigma, theta, model);
29        figure()
30        plot(T_w, C_d_SchaafChambre_flat);
31        xlabel('T_w');
32        ylabel('C_d');
```

```

33     title ( '\sigma = 0' );
34     axis tight
35     saveas(gcf, 'Model_1_2.png')
36 end
37 end
38 if (indx == 2)
39     list = { 'Constant Temperature', 'Constant angle of attack' };
40     [indx, tf] = listdlg ( 'PromptString', 'Select an option:', '
        SelectionMode', 'single', 'ListSize', [400,200], 'ListString',
        list );
41     if (indx == 1)
42         T_w = [0 25 50 75 100] + 273;
43         sigma=0.5;
44         theta = 0:0.01:(pi/2);
45         model=3;
46         [C_d_SchaafChambre_flat] = SchaafChambre_plots_flat(T_w,
            sigma, theta, model);
47         figure ()
48         plot(theta*180/pi, C_d_SchaafChambre_flat(:,1), 'r');
49         hold on
50         plot(theta*180/pi, C_d_SchaafChambre_flat(:,2), 'b');
51         plot(theta*180/pi, C_d_SchaafChambre_flat(:,3), 'g');
52         plot(theta*180/pi, C_d_SchaafChambre_flat(:,4), 'm');
53         plot(theta*180/pi, C_d_SchaafChambre_flat(:,5), 'k');
54         hold off
55         legend('T_w = 0 C', 'T_w = 25 C', 'T_w = 50 C', 'T_w = 75
            C', 'T_w = 100 C', 'Location', 'northwest');
56         xlabel('\theta [degrees]');
57         ylabel('C_d');
58         title ( '\sigma = 1' );
59         axis tight
60         saveas(gcf, 'Model_2_1.png')
61     elseif (indx == 2)
62         T_w = (0:1:100)+273;
63         theta = [0 15 30 45 60 75 90]*pi/180;
64         sigma=1;
65         model=4;
66         [C_d_SchaafChambre_flat] = SchaafChambre_plots_flat(T_w,
            sigma, theta, model);
67         figure ()
68         plot(T_w-273, C_d_SchaafChambre_flat(:,1), 'r');
69         hold on
70         plot(T_w-273, C_d_SchaafChambre_flat(:,2), 'b');
71         plot(T_w-273, C_d_SchaafChambre_flat(:,3), 'g');
72         plot(T_w-273, C_d_SchaafChambre_flat(:,4), 'm');
73         plot(T_w-273, C_d_SchaafChambre_flat(:,5), 'k');
74         plot(T_w-273, C_d_SchaafChambre_flat(:,6), 'c');
75         plot(T_w-273, C_d_SchaafChambre_flat(:,7), 'y');

```

```

76         hold off
77         legend( '\theta = 0 ', '\theta = 15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ', '\theta = 75 ', '\theta = 90 ' );
78         xlabel( 'T_w [ C ] ' );
79         ylabel( 'C_d ' );
80         title( '\sigma = 1 ' );
81         axis tight
82         saveas(gcf, 'Model_2.2.png')
83     end
84 end
85 if(indx == 3)
86     T_w = 50+273;
87     theta = [0 15 30 45 60 75 90]*pi/180;
88     sigma=0:0.001:1;
89     model=5;
90     [C_d_SchaafChambre_flat] = SchaafChambre_plots_flat(T_w,
91         sigma, theta, model);
92     figure()
93     plot(sigma, C_d_SchaafChambre_flat(:,1), 'r');
94     hold on
95     plot(sigma, C_d_SchaafChambre_flat(:,2), 'b');
96     plot(sigma, C_d_SchaafChambre_flat(:,3), 'g');
97     plot(sigma, C_d_SchaafChambre_flat(:,4), 'm');
98     plot(sigma, C_d_SchaafChambre_flat(:,5), 'k');
99     plot(sigma, C_d_SchaafChambre_flat(:,6), 'c');
100    plot(sigma, C_d_SchaafChambre_flat(:,7), 'y');
101    hold off
102    legend( '\theta = 0 ', '\theta = 15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ', '\theta = 75 ', '\theta = 90 ' );
103    xlabel( '\sigma ' );
104    ylabel( 'C_d ' );
105    title( 'T_w = 50 C ' );
106    axis tight
107    saveas(gcf, 'Model_3.png')
108 end

```

#### C.1.1.2. Function

```

1 function [C_d_SchaafChambre_flat] = SchaafChambre_plots_flat(T_w,
2     sigma, theta, model)
3
4     M = xlsread('Datos1.xlsx'); % Read Excel
5     Elements = [M(:,2), M(:,3), M(:,4), M(:,8), M(:,9), M(:,10),
6         M(:,11), M(:,12)]; % Atmospheric composition matrix
7     Height = M(:,1); % Height
8     Temp_neutral = M(:,6); % Neutral Temperature
9     Rt = 6.37E6; % Earth radius
10    Mt = 5.972E24; % Earth mass

```

```

9      G = 6.674E-11; % Gravitational constant
10     mu = []; % Empty mu vector
11     k = 1.3806488E-23; % Boltzmann constant
12     RIG = 8.314472; % Ideal gases constant
13
14     % Molar masses in g/mol
15     mass_O = 15.999; mass_N2 = 28.01340; mass_O2 = 31.99880;
16     mass_He = 4.0026020;
17     mass_Ar = 39.9480; mass_H = 1.007940; mass_N = 14.00670;
18     mass_AnOmO = 15.999;
19
20     % Masses vector in Kg/mol
21     mass_total = [mass_O, mass_N2, mass_O2, mass_He, mass_Ar,
22     mass_H, mass_N, mass_AnOmO]/1000;
23
24     % Mu computation
25     for i=1:1:101
26         mu(i)=(Elements(i,:) * mass_total' / sum(Elements(i,:)));
27     end
28
29     R=RIG*ones(1, length(mu)) ./ (mu);
30
31     V_a = sqrt(2*R' .* Temp_neutral);
32     V_orb = sqrt((G*Mt*ones(101,1)) ./ (Height.*1000+Rt));
33
34     V_a_r = sqrt(2*R(21)*T_w);
35
36     s = (V_orb(21)/V_a(21));
37     s_w = (V_orb(21)*ones(1, length(V_a_r)) ./ V_a_r);
38
39     if (model == 1)
40         for i=1:1:length(T_w)
41             C_d_SchaafChambre_flat_spec(:, i) = ((4*sin(theta)) ./ (
42                 sqrt(pi)*s^2)) .* ((s*sin(theta)) .* exp(-(s*sin(theta)).^2)) + sqrt(pi) * (0.5 + (s*sin(theta)).^2) .* erf(s*sin(theta)));
43         end
44         C_d_SchaafChambre_flat = C_d_SchaafChambre_flat_spec;
45     elseif (model == 2)
46         C_d_SchaafChambre_flat_spec = ((4*sin(theta)) ./ (sqrt(pi)*s^2)) .* ((s*sin(theta)) .* exp(-(s*sin(theta)).^2)) + sqrt(pi) * (0.5 + (s*sin(theta)).^2) .* erf(s*sin(theta)));
47         C_d_SchaafChambre_flat = C_d_SchaafChambre_flat_spec;
48     elseif (model == 3)
49         for i=1:1:length(T_w)
50             C_d_SchaafChambre_flat_diff(:, i) = (2/(sqrt(pi)*s))
51                 * ((exp(-(s*sin(theta)).^2)) + sqrt(pi)*s*sin(theta)
52                 ).*(1+(1/(2*s^2))) .* erf(s*sin(theta)) + ((pi*s)/s_w(

```

```

        i))*(sin(theta).^2));
47     end
48     C_d_SchaafChambre_flat = C_d_SchaafChambre_flat_diff;
49 elseif(model == 4)
50     for i=1:1:length(theta)
51         C_d_SchaafChambre_flat_diff(:,i) = (2/(sqrt(pi)*s))
            *((exp(-(s*sin(theta(i)))^2))+sqrt(pi)*s*sin(
            theta(i))*(1+(1/(2*s^2)))*erf(s*sin(theta(i)))+(
            pi*s)*ones(1,length(s_w))./s_w).*(sin(theta(i))^2
            );
52     end
53     C_d_SchaafChambre_flat = C_d_SchaafChambre_flat_diff;
54 elseif(model == 5)
55     for i=1:1:length(theta)
56         C_d_SchaafChambre_flat_diff(i) = (2/(sqrt(pi)*s))*((
            exp(-(s*sin(theta(i)))^2))+sqrt(pi)*s*sin(theta(
            i))*(1+(1/(2*s^2)))*erf(s*sin(theta(i)))+(pi*s)/
            s_w).*(sin(theta(i))^2));
57         C_d_SchaafChambre_flat_spec(i) = ((4*sin(theta(i)))/(
            sqrt(pi)*s^2))*((s*sin(theta(i)))*exp(-(s*sin(
            theta(i)))^2))+sqrt(pi)*(0.5+(s*sin(theta(i)))^2)*
            erf(s*sin(theta(i))));
58         C_d_SchaafChambre_flat(:,i) =
            C_d_SchaafChambre_flat_diff(i)*sigma +
            C_d_SchaafChambre_flat_spec(i)*(1-sigma);
59     end
60 end
61 end

```

## C.1.2. Sphere

### C.1.2.1. Main menu

```

1 clear all
2 close all
3
4 list = {'Specular','Diffuse','Variable'};
5 [indx,tf] = listdlg('PromptString','Select an option:',
    'SelectionMode','single','ListSize',[400,200],'ListString',list
    );
6
7 if(indx == 1)
8     T_w = (0:1:100)+273;
9     sigma=0;
10    model=1;
11    [C_d_SchaafChambre_sphere] = SchaafChambre_plots_sphere(T_w,
        sigma, model);
12    figure()

```

```

13     plot(T_w-273,C_d_SchaafChambre_sphere);
14     xlabel('T_w [ C ]');
15     ylabel('C_d');
16     title('\sigma = 0');
17     axis tight
18     saveas(gcf,'Model_1.png')
19 end
20 if(indx == 2)
21     T_w = (0:1:100)+273;
22     sigma=1;
23     model=2;
24     [C_d_SchaafChambre_sphere] = SchaafChambre_plots_sphere(T_w,
        sigma, model);
25     figure()
26     plot(T_w-273,C_d_SchaafChambre_sphere);
27     xlabel('T_w [ C ]');
28     ylabel('C_d');
29     title('\sigma = 1');
30     axis tight
31     saveas(gcf,'Model_2.png')
32 end
33 if(indx == 3)
34     T_w = [0 25 50 75 100] + 273;
35     sigma=0:0.01:1;
36     model=3;
37     [C_d_SchaafChambre_sphere] = SchaafChambre_plots_sphere(T_w,
        sigma, model);
38     figure()
39     plot(sigma,C_d_SchaafChambre_sphere(:,1),'r');
40     hold on
41     plot(sigma,C_d_SchaafChambre_sphere(:,2),'b');
42     plot(sigma,C_d_SchaafChambre_sphere(:,3),'g');
43     plot(sigma,C_d_SchaafChambre_sphere(:,4),'m');
44     plot(sigma,C_d_SchaafChambre_sphere(:,5),'k');
45     hold off
46     legend('T_w = 0 C', 'T_w = 25 C', 'T_w = 50 C', 'T_w = 75 C',
        'T_w = 100 C','Location','northwest');
47     xlabel('\sigma');
48     ylabel('C_d');
49     axis tight
50     saveas(gcf,'Model_3.png')
51 end

```

### C.1.2.2. Function

```

1 function [C_d_SchaafChambre_sphere] = SchaafChambre_plots_sphere(
    T_w, sigma, model)
2
3     M = xlsread('Datos1.xlsx'); % Read Excel

```



```

4     Elements = [M(:,2), M(:,3), M(:,4), M(:,8), M(:,9), M(:,10),
      M(:,11), M(:,12)]; % Atmospheric composition matrix
5     Height = M(:,1); % Height
6     Temp_neutral = M(:,6); % Neutral Temperature
7     Rt = 6.37E6; % Earth radius
8     Mt = 5.972E24; % Earth mass
9     G = 6.674E-11; % Gravitational constant
10    mu = []; % Empty mu vector
11    k = 1.3806488E-23; % Boltzmann constant
12    RIG = 8.314472; % Ideal gases constant
13
14    % Molar masses in g/mol
15    mass_O = 15.999; mass_N2 = 28.01340; mass_O2 = 31.99880;
      mass_He = 4.0026020;
16    mass_Ar = 39.9480; mass_H = 1.007940; mass_N = 14.00670;
      mass_Anomo = 15.999;
17
18    % Masses vector in Kg/mol
19    mass_total = [mass_O, mass_N2, mass_O2, mass_He, mass_Ar,
      mass_H, mass_N, mass_Anomo]/1000;
20
21    % Mu computation
22    for i=1:1:101
23        mu(i)=(Elements(i,:) * mass_total' / sum(Elements(i,:)));
24    end
25
26    R=RIG*ones(1, length(mu)) ./ (mu);
27
28    V_a = sqrt(2*R' .* Temp_neutral);
29    V_orb = sqrt((G*Mt*ones(101,1)) ./ (Height.*1000+Rt));
30
31    V_a_r = sqrt(2*R(21)*T_w);
32
33    s = (V_orb(21)/V_a(21));
34    s_w = (V_orb(21)*ones(1, length(V_a_r)) ./ V_a_r);
35
36    if (model == 1 || model == 2)
37        C_d_SchaafChambre_sphere = (exp(-(s^2)/2) / (sqrt(pi)*s^3))
      *(1+2*(s^2)) + ((4*(s^4)+4*(s^2)-1)/(2*(s^4)))*erf(s)
      + (2*sigma*sqrt(pi)*ones(1, length(s_w))) ./ (3*s_w);
38    elseif (model == 3)
39        for i=1:1:length(s_w)
40            C_d_SchaafChambre_sphere(:, i) = (exp(-(s^2)/2) / (sqrt(
      pi)*s^3)) * (1+2*(s^2)) + ((4*(s^4)+4*(s^2)-1)/(2*(s
      ^4)))*erf(s) + (2*sigma.*sqrt(pi)) ./ (3*s_w(i));
41        end
42    end
43 end

```

## C.2. Schamberg model

### C.2.1. Flat plate

#### C.2.1.1. Main menu

```
1 clear all
2 close all
3
4 list = {'Specular', 'Diffuse', 'Variable'};
5 [indx, tf] = listdlg('PromptString', 'Select an option:', '
    SelectionMode', 'single', 'ListSize', [400, 200], 'ListString', list
    );
6
7 %% SPECULAR
8
9 if (indx == 1)
10     list = {'Constant Temperature', 'Constant accomodation factor'
        , 'Constant angle of attack'};
11     [indx, tf] = listdlg('PromptString', 'Select an option:', '
        SelectionMode', 'single', 'ListSize', [400, 200], 'ListString',
        list);
12
13     if (indx == 1)
14         T_w1 = 0+273;
15         T_w2 = 50+273;
16         T_w3 = 100+273;
17         alpha = [0 0.25 0.5 0.75 1];
18         angle_of_attack = 0:0.01:(pi/2);
19         phi = 1;
20         nu_inv = 1;
21         model = 1;
22         [C_d_1, V_rel1, phi] = Schamberg_plots_flat(T_w1, alpha,
            angle_of_attack, phi, nu_inv, model);
23         [C_d_2, V_rel2, phi] = Schamberg_plots_flat(T_w2, alpha,
            angle_of_attack, phi, nu_inv, model);
24         [C_d_3, V_rel3, phi] = Schamberg_plots_flat(T_w3, alpha,
            angle_of_attack, phi, nu_inv, model);
25
26         figure()
27         p1 = plot(angle_of_attack*180/pi, C_d_1(:,1), 'r');
28         hold on
29         p2 = plot(angle_of_attack*180/pi, C_d_1(:,2), 'b');
30         p3 = plot(angle_of_attack*180/pi, C_d_1(:,3), 'g');
31         p4 = plot(angle_of_attack*180/pi, C_d_1(:,4), 'm');
32         p5 = plot(angle_of_attack*180/pi, C_d_1(:,5), 'k');
33         p6 = plot(angle_of_attack*180/pi, C_d_2(:,1), '—r');
34         p7 = plot(angle_of_attack*180/pi, C_d_2(:,2), '—b');
```

```

35     p8 = plot(angle_of_attack*180/pi, C_d_2(:,3), '—g');
36     p9 = plot(angle_of_attack*180/pi, C_d_2(:,4), '—m');
37     p10 = plot(angle_of_attack*180/pi, C_d_2(:,5), '—k');
38     p11 = plot(angle_of_attack*180/pi, C_d_3(:,1), ':r');
39     p12 = plot(angle_of_attack*180/pi, C_d_3(:,2), ':b');
40     p13 = plot(angle_of_attack*180/pi, C_d_3(:,3), ':g');
41     p14 = plot(angle_of_attack*180/pi, C_d_3(:,4), ':m');
42     p15= plot(angle_of_attack*180/pi, C_d_3(:,5), ':k');
43     hold off
44     legend([p1 p2 p3 p4 p5],{'\alpha = 0', '\alpha = 0.25', '\alpha = 0.5', '\alpha = 0.75', '\alpha = 1'}, 'Location','northwest')
45     xlabel('Angle of attack [degrees]');
46     ylabel('C_d');
47     title('T_w = 0 C, 50 C, 100 C');
48     axis tight
49     saveas(gcf, 'Model_1.png')
50 end
51 if (indx == 2)
52     alpha1 = 0;
53     alpha2 = 0.5;
54     alpha3 = 1;
55     T_w = [0 25 50 75 100] + 273;
56     angle_of_attack = 0:0.01:(pi/2);
57     phi = 1;
58     nu_inv = 1;
59     model = 2;
60     [C_d_1, V_rel1, phi] = Schamberg_plots_flat(T_w, alpha1, angle_of_attack, phi, nu_inv, model);
61     [C_d_2, V_rel2, phi] = Schamberg_plots_flat(T_w, alpha2, angle_of_attack, phi, nu_inv, model);
62     [C_d_3, V_rel3, phi] = Schamberg_plots_flat(T_w, alpha3, angle_of_attack, phi, nu_inv, model);
63
64     figure()
65     p1 = plot(angle_of_attack*180/pi, C_d_1(:,1), 'r');
66     hold on
67     p2 = plot(angle_of_attack*180/pi, C_d_1(:,2), 'b');
68     p3 = plot(angle_of_attack*180/pi, C_d_1(:,3), 'g');
69     p4 = plot(angle_of_attack*180/pi, C_d_1(:,4), 'm');
70     p5 = plot(angle_of_attack*180/pi, C_d_1(:,5), 'k');
71     p6 = plot(angle_of_attack*180/pi, C_d_2(:,1), '—r');
72     p7 = plot(angle_of_attack*180/pi, C_d_2(:,2), '—b');
73     p8 = plot(angle_of_attack*180/pi, C_d_2(:,3), '—g');
74     p9 = plot(angle_of_attack*180/pi, C_d_2(:,4), '—m');
75     p10 = plot(angle_of_attack*180/pi, C_d_2(:,5), '—k');
76     p11 = plot(angle_of_attack*180/pi, C_d_3(:,1), ':r');
77     p12 = plot(angle_of_attack*180/pi, C_d_3(:,2), ':b');

```

```

78     p13 = plot(angle_of_attack*180/pi, C_d_3(:,3), 'g');
79     p14 = plot(angle_of_attack*180/pi, C_d_3(:,4), 'm');
80     p15 = plot(angle_of_attack*180/pi, C_d_3(:,5), 'k');
81     hold off
82     legend([p1 p2 p3 p4 p5], {'T_w = 0 C', 'T_w = 25 C', 'T_w
      = 50 C', 'T_w = 75 C', 'T_w = 100 C'}, 'Location', '
      northwest')
83     xlabel('Angle of attack [degrees]');
84     ylabel('C_d');
85     title('\alpha = 0, 0.5, 1');
86     axis tight
87     saveas(gcf, 'Model_2.png')
88 end
89 if (indx == 3)
90     alpha = 0:0.001:1;
91     T_w = [0 25 50 75 100] + 273;
92     angle_of_attack = [0 15 30 45 60 75 90]*pi/180;
93     phi = 1;
94     nu_inv = 1;
95     model = 3;
96     [C_d_1, V_rel1, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(1), phi, nu_inv, model);
97     [C_d_2, V_rel2, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(2), phi, nu_inv, model);
98     [C_d_3, V_rel3, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(3), phi, nu_inv, model);
99     [C_d_4, V_rel4, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(4), phi, nu_inv, model);
100    [C_d_5, V_rel5, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(5), phi, nu_inv, model);
101    [C_d_6, V_rel6, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(6), phi, nu_inv, model);
102    [C_d_7, V_rel7, phi] = Schamberg_plots_flat(T_w, alpha,
      angle_of_attack(7), phi, nu_inv, model);
103
104    figure()
105    p1 = plot(alpha, C_d_1(:,1), 'r');
106    hold on
107    p2 = plot(alpha, C_d_1(:,2), 'b');
108    p3 = plot(alpha, C_d_1(:,3), 'g');
109    p4 = plot(alpha, C_d_1(:,4), 'm');
110    p5 = plot(alpha, C_d_1(:,5), 'k');
111    p6 = plot(alpha, C_d_2(:,1), '—r');
112    p7 = plot(alpha, C_d_2(:,2), '—b');
113    p8 = plot(alpha, C_d_2(:,3), '—g');
114    p9 = plot(alpha, C_d_2(:,4), '—m');
115    p10 = plot(alpha, C_d_2(:,5), '—k');
116    p11 = plot(alpha, C_d_3(:,1), 'r');

```

```

117 p12 = plot(alpha , C_d_3(:,2) , ':b');
118 p13 = plot(alpha , C_d_3(:,3) , ':g');
119 p14 = plot(alpha , C_d_3(:,4) , ':m');
120 p15= plot(alpha , C_d_3(:,5) , ':k');
121 hold off
122 legend([p1 p2 p3 p4 p5],{'T_w = 0 C' , 'T_w = 25 C' , 'T_w
    = 50 C' , 'T_w = 75 C' , 'T_w = 100 C'},'Location','
    northwest')
123 xlabel( '\alpha' );
124 ylabel( 'C_d' );
125 title( '\theta = 0 , 15 , 30 ');
126 axis tight
127 saveas(gcf , 'Model_3_1.png')
128
129 figure()
130 plot(alpha , C_d_4);
131 legend('T_w = 0 C' , 'T_w = 25 C' , 'T_w = 50 C' , 'T_w = 75
    C' , 'T_w = 100 C')
132 xlabel( '\alpha' );
133 ylabel( 'C_d' );
134 title( '\theta = 45 ');
135 axis tight
136 saveas(gcf , 'Model_3_2.png')
137
138 figure()
139 p1 = plot(alpha , C_d_5(:,1) , 'r');
140 hold on
141 p2 = plot(alpha , C_d_5(:,2) , 'b');
142 p3 = plot(alpha , C_d_5(:,3) , 'g');
143 p4 = plot(alpha , C_d_5(:,4) , 'm');
144 p5 = plot(alpha , C_d_5(:,5) , 'k');
145 p6 = plot(alpha , C_d_6(:,1) , '—r');
146 p7 = plot(alpha , C_d_6(:,2) , '—b');
147 p8 = plot(alpha , C_d_6(:,3) , '—g');
148 p9 = plot(alpha , C_d_6(:,4) , '—m');
149 p10 = plot(alpha , C_d_6(:,5) , '—k');
150 p11 = plot(alpha , C_d_7(:,1) , ':r');
151 p12 = plot(alpha , C_d_7(:,2) , ':b');
152 p13 = plot(alpha , C_d_7(:,3) , ':g');
153 p14 = plot(alpha , C_d_7(:,4) , ':m');
154 p15 = plot(alpha , C_d_7(:,5) , ':k');
155 hold off
156 legend([p1 p2 p3 p4 p5],{'T_w = 0 C' , 'T_w = 25 C' , 'T_w
    = 50 C' , 'T_w = 75 C' , 'T_w = 100 C'},'Location','
    northwest')
157 xlabel( '\alpha' );
158 ylabel( 'C_d' );
159 title( '\theta = 60 , 75 , 90 ');

```

```

160         axis tight
161         saveas(gcf, 'Model_3.3.png')
162     end
163 end
164
165 %% DIFFUSE
166
167 if (indx == 2)
168     list = {'Constant Temperature', 'Constant accomodation factor'
169           , 'Constant angle of attack'};
169     [indx, tf] = listdlg('PromptString', 'Select an option:', '
170                         SelectionMode', 'single', 'ListSize', [400, 200], 'ListString',
171                         list);
172
173 if (indx == 1)
174     T_w1 = 0+273;
175     T_w2 = 50+273;
176     T_w3 = 100+273;
177     alpha = [0 0.25 0.5 0.75 1];
178     angle_of_attack = 0.00001:0.00001:(pi/2);
179     phi = 2/3;
180     nu_inv = 0;
181     model = 4;
182     [C_d_1, V_rel1, phi] = Schamberg_plots_flat(T_w1, alpha,
183         angle_of_attack, phi, nu_inv, model);
184     [C_d_2, V_rel2, phi] = Schamberg_plots_flat(T_w2, alpha,
185         angle_of_attack, phi, nu_inv, model);
186     [C_d_3, V_rel3, phi] = Schamberg_plots_flat(T_w3, alpha,
187         angle_of_attack, phi, nu_inv, model);
188
189 figure()
190 p1 = plot(angle_of_attack*180/pi, C_d_1(:,1), 'r');
191 hold on
192 p2 = plot(angle_of_attack*180/pi, C_d_1(:,2), 'b');
193 p3 = plot(angle_of_attack*180/pi, C_d_1(:,3), 'g');
194 p4 = plot(angle_of_attack*180/pi, C_d_1(:,4), 'm');
195 p5 = plot(angle_of_attack*180/pi, C_d_1(:,5), 'k');
196 p6 = plot(angle_of_attack*180/pi, C_d_2(:,1), '—r');
197 p7 = plot(angle_of_attack*180/pi, C_d_2(:,2), '—b');
198 p8 = plot(angle_of_attack*180/pi, C_d_2(:,3), '—g');
199 p9 = plot(angle_of_attack*180/pi, C_d_2(:,4), '—m');
200 p10 = plot(angle_of_attack*180/pi, C_d_2(:,5), '—k');
201 p11 = plot(angle_of_attack*180/pi, C_d_3(:,1), ':r');
202 p12 = plot(angle_of_attack*180/pi, C_d_3(:,2), ':b');
203 p13 = plot(angle_of_attack*180/pi, C_d_3(:,3), ':g');
204 p14 = plot(angle_of_attack*180/pi, C_d_3(:,4), ':m');
205 p15 = plot(angle_of_attack*180/pi, C_d_3(:,5), ':k');
206 hold off

```

```

202     legend([p1 p2 p3 p4 p5],{'\alpha = 0', '\alpha = 0.25', '
        \alpha = 0.5', '\alpha = 0.75', '\alpha = 1'}, '
        Location', 'northwest')
203     xlabel('Angle of attack [degrees]');
204     ylabel('C_d');
205     title('T_w = 0 C , 50 C , 100 C ');
206     axis tight
207     saveas(gcf, 'Model_4.png')
208 end
209 if(indx == 2)
210     alpha1 = 0;
211     alpha2 = 0.5;
212     alpha3 = 1;
213     T_w = [0 25 50 75 100] + 273;
214     angle_of_attack = 0.00001:0.00001:(pi/2);
215     phi = 2/3;
216     nu_inv = 0;
217     model = 5;
218     [C_d_1, V_rel1, phi] = Schamberg_plots_flat(T_w, alpha1,
        angle_of_attack, phi, nu_inv, model);
219     [C_d_2, V_rel2, phi] = Schamberg_plots_flat(T_w, alpha2,
        angle_of_attack, phi, nu_inv, model);
220     [C_d_3, V_rel3, phi] = Schamberg_plots_flat(T_w, alpha3,
        angle_of_attack, phi, nu_inv, model);
221
222     figure()
223     p1 = plot(angle_of_attack*180/pi, C_d_1(:,1), 'r');
224     hold on
225     p2 = plot(angle_of_attack*180/pi, C_d_1(:,2), 'b');
226     p3 = plot(angle_of_attack*180/pi, C_d_1(:,3), 'g');
227     p4 = plot(angle_of_attack*180/pi, C_d_1(:,4), 'm');
228     p5 = plot(angle_of_attack*180/pi, C_d_1(:,5), 'k');
229     p6 = plot(angle_of_attack*180/pi, C_d_2(:,1), '—r');
230     p7 = plot(angle_of_attack*180/pi, C_d_2(:,2), '—b');
231     p8 = plot(angle_of_attack*180/pi, C_d_2(:,3), '—g');
232     p9 = plot(angle_of_attack*180/pi, C_d_2(:,4), '—m');
233     p10 = plot(angle_of_attack*180/pi, C_d_2(:,5), '—k');
234     p11 = plot(angle_of_attack*180/pi, C_d_3(:,1), ':r');
235     p12 = plot(angle_of_attack*180/pi, C_d_3(:,2), ':b');
236     p13 = plot(angle_of_attack*180/pi, C_d_3(:,3), ':g');
237     p14 = plot(angle_of_attack*180/pi, C_d_3(:,4), ':m');
238     p15 = plot(angle_of_attack*180/pi, C_d_3(:,5), ':k');
239     hold off
240     legend([p1 p2 p3 p4 p5],{'T_w = 0 C', 'T_w = 25 C', 'T_w
        = 50 C', 'T_w = 75 C', 'T_w = 100 C'}, 'Location', '
        northwest')
241     xlabel('Angle of attack [degrees]');
242     ylabel('C_d');

```

```

243     title ( '\alpha = 0, 0.5, 1 ');
244     axis tight
245     saveas(gcf, 'Model_5.png')
246 end
247 if (indx == 3)
248     alpha = 0:0.001:1;
249     T_w = [0 25 50 75 100] + 273;
250     angle_of_attack = [0 15 30 45 60 75 90]*pi/180;
251     phi = 2/3;
252     nu_inv = 0;
253     model = 6;
254     [C_d_1, V_rel1, phi] = Schamberg_plots_flat(T_w, alpha,
255         angle_of_attack(1), phi, nu_inv, model);
256     [C_d_2, V_rel2, phi] = Schamberg_plots_flat(T_w, alpha,
257         angle_of_attack(2), phi, nu_inv, model);
258     [C_d_3, V_rel3, phi] = Schamberg_plots_flat(T_w, alpha,
259         angle_of_attack(3), phi, nu_inv, model);
260     [C_d_4, V_rel4, phi] = Schamberg_plots_flat(T_w, alpha,
261         angle_of_attack(4), phi, nu_inv, model);
262     [C_d_5, V_rel5, phi] = Schamberg_plots_flat(T_w, alpha,
263         angle_of_attack(5), phi, nu_inv, model);
264     [C_d_6, V_rel6, phi] = Schamberg_plots_flat(T_w, alpha,
265         angle_of_attack(6), phi, nu_inv, model);
266     [C_d_7, V_rel7, phi] = Schamberg_plots_flat(T_w, alpha,
267         angle_of_attack(7), phi, nu_inv, model);
268
269     figure ()
270     plot(alpha, C_d_1);
271     legend('T_w = 0 C', 'T_w = 25 C', 'T_w = 50 C', 'T_w = 75
272         C', 'T_w = 100 C', 'Location', 'northwest')
273     xlabel( '\alpha ');
274     ylabel( 'C_d' );
275     title ( '\theta = 0 ');
276     axis tight
277     saveas(gcf, 'Model_6_1.png')
278
279     figure ()
280     p1 = plot(alpha, C_d_2(:,1), 'r');
281     hold on
282     p2 = plot(alpha, C_d_2(:,2), 'b');
283     p3 = plot(alpha, C_d_2(:,3), 'g');
284     p4 = plot(alpha, C_d_2(:,4), 'm');
285     p5 = plot(alpha, C_d_2(:,5), 'k');
286     p6 = plot(alpha, C_d_3(:,1), '—r');
287     p7 = plot(alpha, C_d_3(:,2), '—b');
288     p8 = plot(alpha, C_d_3(:,3), '—g');
289     p9 = plot(alpha, C_d_3(:,4), '—m');
290     p10 = plot(alpha, C_d_3(:,5), '—k');

```



```

283 p11 = plot(alpha , C_d_4(:,1) , ':r' );
284 p12 = plot(alpha , C_d_4(:,2) , ':b' );
285 p13 = plot(alpha , C_d_4(:,3) , ':g' );
286 p14 = plot(alpha , C_d_4(:,4) , ':m' );
287 p15 = plot(alpha , C_d_4(:,5) , ':k' );
288 hold off
289 legend([p1 p2 p3 p4 p5],{ 'T_w = 0 C ', 'T_w = 25 C ', 'T_w
    = 50 C ', 'T_w = 75 C ', 'T_w = 100 C '}, 'Location ', '
    northeast ')
290 xlabel( '\alpha' );
291 ylabel( 'C_d' );
292 title( '\theta = 15 , 30 , 45 ');
293 axis tight
294 saveas(gcf, 'Model_6.2.png')
295
296 figure()
297 p1 = plot(alpha , C_d_5(:,1) , 'r' );
298 hold on
299 p2 = plot(alpha , C_d_5(:,2) , 'b' );
300 p3 = plot(alpha , C_d_5(:,3) , 'g' );
301 p4 = plot(alpha , C_d_5(:,4) , 'm' );
302 p5 = plot(alpha , C_d_5(:,5) , 'k' );
303 p6 = plot(alpha , C_d_6(:,1) , '—r' );
304 p7 = plot(alpha , C_d_6(:,2) , '—b' );
305 p8 = plot(alpha , C_d_6(:,3) , '—g' );
306 p9 = plot(alpha , C_d_6(:,4) , '—m' );
307 p10 = plot(alpha , C_d_6(:,5) , '—k' );
308 p11 = plot(alpha , C_d_7(:,1) , ':r' );
309 p12 = plot(alpha , C_d_7(:,2) , ':b' );
310 p13 = plot(alpha , C_d_7(:,3) , ':g' );
311 p14 = plot(alpha , C_d_7(:,4) , ':m' );
312 p15 = plot(alpha , C_d_7(:,5) , ':k' );
313 hold off
314 legend([p1 p2 p3 p4 p5],{ 'T_w = 0 C ', 'T_w = 25 C ', 'T_w
    = 50 C ', 'T_w = 75 C ', 'T_w = 100 C '}, 'Location ', '
    northeast ')
315 xlabel( '\alpha' );
316 ylabel( 'C_d' );
317 title( '\theta = 60 , 75 , 90 ');
318 axis tight
319 saveas(gcf, 'Model_6.3.png')
320 end
321 end
322
323 %% VARIABLE
324
325 if (indx == 3)
326     list = { 'Alpha = 0 & T_w = 0 C / 50 C / 100 C ', 'Alpha = 0.5 &

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```

    T_w = 0 C / 50 C / 100 C ', 'Alpha = 1 & T_w = 0 C / 50 C /
    100 C '};
327 [indx, tf] = listdlg('PromptString', 'Select an option:', '
    SelectionMode', 'single', 'ListSize', [400, 200], 'ListString',
    list);
328
329 if(indx == 1)
330     alpha = 0;
331     T_w1 = 0+273;
332     T_w2 = 50+273;
333     T_w3 = 100+273;
334     angle_of_attack = [0 15 30 45 60 75 90]*pi/180;
335     phi_0 = 0:0.001:(pi/2);
336     nu_inv = 1:-(1/(length(phi_0)-1)):0;
337     model = 7;
338     [C_d_1, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(1), phi_0, nu_inv, model);
339     [C_d_2, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(2), phi_0, nu_inv, model);
340     [C_d_3, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(3), phi_0, nu_inv, model);
341     [C_d_4, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(4), phi_0, nu_inv, model);
342     [C_d_5, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(5), phi_0, nu_inv, model);
343     [C_d_6, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(6), phi_0, nu_inv, model);
344     [C_d_7, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(7), phi_0, nu_inv, model);
345     [C_d_8, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(1), phi_0, nu_inv, model);
346     [C_d_9, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(2), phi_0, nu_inv, model);
347     [C_d_10, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(3), phi_0, nu_inv, model);
348     [C_d_11, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(4), phi_0, nu_inv, model);
349     [C_d_12, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(5), phi_0, nu_inv, model);
350     [C_d_13, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(6), phi_0, nu_inv, model);
351     [C_d_14, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(7), phi_0, nu_inv, model);
352     [C_d_15, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
        angle_of_attack(1), phi_0, nu_inv, model);
353     [C_d_16, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
        angle_of_attack(2), phi_0, nu_inv, model);
354     [C_d_17, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,

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```

    angle_of_attack(3), phi_0, nu_inv, model);
355 [C_d_18, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(4), phi_0, nu_inv, model);
356 [C_d_19, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(5), phi_0, nu_inv, model);
357 [C_d_20, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(6), phi_0, nu_inv, model);
358 [C_d_21, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(7), phi_0, nu_inv, model);

359
360 figure()
361 p1 = plot(phi, C_d_1, 'r');
362 hold on
363 p2 = plot(phi, C_d_2, 'b');
364 p3 = plot(phi, C_d_3, 'g');
365 p4 = plot(phi, C_d_4, 'm');
366 p5 = plot(phi, C_d_5, 'k');
367 p6 = plot(phi, C_d_6, 'c');
368 p7 = plot(phi, C_d_7, 'y');
369
370 p8 = plot(phi, C_d_8, '—r');
371 p9 = plot(phi, C_d_9, '—b');
372 p10 = plot(phi, C_d_10, '—g');
373 p11 = plot(phi, C_d_11, '—m');
374 p12 = plot(phi, C_d_12, '—k');
375 p13 = plot(phi, C_d_13, '—c');
376 p14 = plot(phi, C_d_14, '—y');
377
378 p15 = plot(phi, C_d_15, ':r');
379 p16 = plot(phi, C_d_16, ':b');
380 p17 = plot(phi, C_d_17, ':g');
381 p18 = plot(phi, C_d_18, ':m');
382 p19 = plot(phi, C_d_19, ':k');
383 p20 = plot(phi, C_d_20, ':c');
384 p21 = plot(phi, C_d_21, ':y');
385 hold off
386 legend([p1 p2 p3 p4 p5 p6 p7], {'\theta = 0 ', '\theta =
    15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ',
    '\theta = 75 ', '\theta = 90 '});
387 xlabel('\phi(\phi_0)');
388 ylabel('C_d');
389 title('T_w = 0 C / 50 C / 100 C & \alpha = 0');
390 axis tight
391 saveas(gcf, 'Model_7_1.png')
392
393 figure()
394 p1 = plot(nu_inv, C_d_1, 'r');
395 hold on

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```

396 p2 = plot(nu_inv, C_d_2, 'b');
397 p3 = plot(nu_inv, C_d_3, 'g');
398 p4 = plot(nu_inv, C_d_4, 'm');
399 p5 = plot(nu_inv, C_d_5, 'k');
400 p6 = plot(nu_inv, C_d_6, 'c');
401 p7 = plot(nu_inv, C_d_7, 'y');
402
403 p8 = plot(nu_inv, C_d_8, '—r');
404 p9 = plot(nu_inv, C_d_9, '—b');
405 p10 = plot(nu_inv, C_d_10, '—g');
406 p11 = plot(nu_inv, C_d_11, '—m');
407 p12 = plot(nu_inv, C_d_12, '—k');
408 p13 = plot(nu_inv, C_d_13, '—c');
409 p14 = plot(nu_inv, C_d_14, '—y');
410
411 p15 = plot(nu_inv, C_d_15, ':r');
412 p16 = plot(nu_inv, C_d_16, ':b');
413 p17 = plot(nu_inv, C_d_17, ':g');
414 p18 = plot(nu_inv, C_d_18, ':m');
415 p19 = plot(nu_inv, C_d_19, ':k');
416 p20 = plot(nu_inv, C_d_20, ':c');
417 p21 = plot(nu_inv, C_d_21, ':y');
418 hold off
419 legend([p1 p2 p3 p4 p5 p6 p7], {'\theta = 0 ', '\theta =
    15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ',
    '\theta = 75 ', '\theta = 90 '});
420 xlabel('1/\nu');
421 ylabel('C_d');
422 title('T_w = 0 C / 50 C / 100 C & \alpha = 0');
423 axis tight
424 saveas(gcf, 'Model_7.2.png')
425 end
426 if(indx == 2)
427     alpha = 0.5;
428     T_w1 = 0+273;
429     T_w2 = 50+273;
430     T_w3 = 100+273;
431     angle_of_attack = [0 15 30 45 60 75 90]*pi/180;
432     phi_0 = 0:0.001:(pi/2);
433     nu_inv = 1:-(1/(length(phi_0)-1)):0;
434     model = 7;
435     [C_d_1, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(1), phi_0, nu_inv, model);
436     [C_d_2, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(2), phi_0, nu_inv, model);
437     [C_d_3, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(3), phi_0, nu_inv, model);
438     [C_d_4, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,

```

```

    angle_of_attack(4), phi_0, nu_inv, model);
439 [C_d_5, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
    angle_of_attack(5), phi_0, nu_inv, model);
440 [C_d_6, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
    angle_of_attack(6), phi_0, nu_inv, model);
441 [C_d_7, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
    angle_of_attack(7), phi_0, nu_inv, model);
442 [C_d_8, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(1), phi_0, nu_inv, model);
443 [C_d_9, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(2), phi_0, nu_inv, model);
444 [C_d_10, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(3), phi_0, nu_inv, model);
445 [C_d_11, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(4), phi_0, nu_inv, model);
446 [C_d_12, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(5), phi_0, nu_inv, model);
447 [C_d_13, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(6), phi_0, nu_inv, model);
448 [C_d_14, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
    angle_of_attack(7), phi_0, nu_inv, model);
449 [C_d_15, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(1), phi_0, nu_inv, model);
450 [C_d_16, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(2), phi_0, nu_inv, model);
451 [C_d_17, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(3), phi_0, nu_inv, model);
452 [C_d_18, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(4), phi_0, nu_inv, model);
453 [C_d_19, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(5), phi_0, nu_inv, model);
454 [C_d_20, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(6), phi_0, nu_inv, model);
455 [C_d_21, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
    angle_of_attack(7), phi_0, nu_inv, model);
456
457 figure ()
458 p1 = plot(phi, C_d_1, 'r');
459 hold on
460 p2 = plot(phi, C_d_2, 'b');
461 p3 = plot(phi, C_d_3, 'g');
462 p4 = plot(phi, C_d_4, 'm');
463 p5 = plot(phi, C_d_5, 'k');
464 p6 = plot(phi, C_d_6, 'c');
465 p7 = plot(phi, C_d_7, 'y');
466
467 p8 = plot(phi, C_d_8, '—r');
468 p9 = plot(phi, C_d_9, '—b');

```

```

469 p10 = plot(phi,C_d_10,'—g');
470 p11 = plot(phi,C_d_11,'—m');
471 p12 = plot(phi,C_d_12,'—k');
472 p13 = plot(phi,C_d_13,'—c');
473 p14 = plot(phi,C_d_14,'—y');
474
475 p15 = plot(phi,C_d_15,':r');
476 p16 = plot(phi,C_d_16,':b');
477 p17 = plot(phi,C_d_17,':g');
478 p18 = plot(phi,C_d_18,':m');
479 p19 = plot(phi,C_d_19,':k');
480 p20 = plot(phi,C_d_20,':c');
481 p21 = plot(phi,C_d_21,':y');
482 hold off
483 legend([p1 p2 p3 p4 p5 p6 p7],{'\theta = 0 ', '\theta =
    15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ',
    '\theta = 75 ', '\theta = 90 '});
484 xlabel('\phi(\phi_0)');
485 ylabel('C_d');
486 title('T_w = 0 C / 50 C / 100 C & \alpha = 0.5');
487 axis tight
488 saveas(gcf,'Model_8_1.png')
489
490 figure()
491 %T_w = 0 C
492 p1 = plot(nu_inv,C_d_1,'r');
493 hold on
494 p2 = plot(nu_inv,C_d_2,'b');
495 p3 = plot(nu_inv,C_d_3,'g');
496 p4 = plot(nu_inv,C_d_4,'m');
497 p5 = plot(nu_inv,C_d_5,'k');
498 p6 = plot(nu_inv,C_d_6,'c');
499 p7 = plot(nu_inv,C_d_7,'y');
500 %T_w = 50 C
501 p8 = plot(nu_inv,C_d_8,'—r');
502 p9 = plot(nu_inv,C_d_9,'—b');
503 p10 = plot(nu_inv,C_d_10,'—g');
504 p11 = plot(nu_inv,C_d_11,'—m');
505 p12 = plot(nu_inv,C_d_12,'—k');
506 p13 = plot(nu_inv,C_d_13,'—c');
507 p14 = plot(nu_inv,C_d_14,'—y');
508 %T_w = 100 C
509 p15 = plot(nu_inv,C_d_15,':r');
510 p16 = plot(nu_inv,C_d_16,':b');
511 p17 = plot(nu_inv,C_d_17,':g');
512 p18 = plot(nu_inv,C_d_18,':m');
513 p19 = plot(nu_inv,C_d_19,':k');
514 p20 = plot(nu_inv,C_d_20,':c');

```

```

515     p21 = plot(nu_inv, C_d_21, ':y');
516     hold off
517     legend([p1 p2 p3 p4 p5 p6 p7], {'\theta = 0 ', '\theta =
        15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ',
        '\theta = 75 ', '\theta = 90 '});
518     xlabel('1/\nu');
519     ylabel('C_d');
520     title('T_w = 0 C / 50 C / 100 C & \alpha = 0.5');
521     axis tight
522     saveas(gcf, 'Model_8.2.png')
523 end
524 if(indx == 3)
525     alpha = 1;
526     T_w1 = 0+273;
527     T_w2 = 50+273;
528     T_w3 = 100+273;
529     angle_of_attack = [0 15 30 45 60 75 90]*pi/180;
530     phi_0 = 0:0.001:(pi/2);
531     nu_inv = 1:-(1/(length(phi_0)-1)):0;
532     model = 7;
533     [C_d_1, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(1), phi_0, nu_inv, model);
534     [C_d_2, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(2), phi_0, nu_inv, model);
535     [C_d_3, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(3), phi_0, nu_inv, model);
536     [C_d_4, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(4), phi_0, nu_inv, model);
537     [C_d_5, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(5), phi_0, nu_inv, model);
538     [C_d_6, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(6), phi_0, nu_inv, model);
539     [C_d_7, V_rel, phi] = Schamberg_plots_flat(T_w1, alpha,
        angle_of_attack(7), phi_0, nu_inv, model);
540     [C_d_8, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(1), phi_0, nu_inv, model);
541     [C_d_9, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(2), phi_0, nu_inv, model);
542     [C_d_10, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(3), phi_0, nu_inv, model);
543     [C_d_11, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(4), phi_0, nu_inv, model);
544     [C_d_12, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(5), phi_0, nu_inv, model);
545     [C_d_13, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(6), phi_0, nu_inv, model);
546     [C_d_14, V_rel, phi] = Schamberg_plots_flat(T_w2, alpha,
        angle_of_attack(7), phi_0, nu_inv, model);

```

```

547 [C_d_15, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(1), phi_0, nu_inv, model);
548 [C_d_16, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(2), phi_0, nu_inv, model);
549 [C_d_17, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(3), phi_0, nu_inv, model);
550 [C_d_18, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(4), phi_0, nu_inv, model);
551 [C_d_19, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(5), phi_0, nu_inv, model);
552 [C_d_20, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(6), phi_0, nu_inv, model);
553 [C_d_21, V_rel, phi] = Schamberg_plots_flat(T_w3, alpha,
      angle_of_attack(7), phi_0, nu_inv, model);

554
555 figure()
556 p1 = plot(phi, C_d_1, 'r');
557 hold on
558 p2 = plot(phi, C_d_2, 'b');
559 p3 = plot(phi, C_d_3, 'g');
560 p4 = plot(phi, C_d_4, 'm');
561 p5 = plot(phi, C_d_5, 'k');
562 p6 = plot(phi, C_d_6, 'c');
563 p7 = plot(phi, C_d_7, 'y');
564
565 p8 = plot(phi, C_d_8, '—r');
566 p9 = plot(phi, C_d_9, '—b');
567 p10 = plot(phi, C_d_10, '—g');
568 p11 = plot(phi, C_d_11, '—m');
569 p12 = plot(phi, C_d_12, '—k');
570 p13 = plot(phi, C_d_13, '—c');
571 p14 = plot(phi, C_d_14, '—y');
572
573 p15 = plot(phi, C_d_15, ':r');
574 p16 = plot(phi, C_d_16, ':b');
575 p17 = plot(phi, C_d_17, ':g');
576 p18 = plot(phi, C_d_18, ':m');
577 p19 = plot(phi, C_d_19, ':k');
578 p20 = plot(phi, C_d_20, ':c');
579 p21 = plot(phi, C_d_21, ':y');
580 hold off
581 legend([p1 p2 p3 p4 p5 p6 p7],{'\theta = 0 ', '\theta =
      15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ',
      '\theta = 75 ', '\theta = 90 '});
582 xlabel('\phi(\phi_0)');
583 ylabel('C_d');
584 title('T_w = 0 C / 50 C / 100 C & \alpha = 1');
585 axis tight

```



```

586     saveas(gcf, 'Model_9_1.png')
587
588     figure()
589     %T_w = 0 C
590     p1 = plot(nu_inv, C_d_1, 'r');
591     hold on
592     p2 = plot(nu_inv, C_d_2, 'b');
593     p3 = plot(nu_inv, C_d_3, 'g');
594     p4 = plot(nu_inv, C_d_4, 'm');
595     p5 = plot(nu_inv, C_d_5, 'k');
596     p6 = plot(nu_inv, C_d_6, 'c');
597     p7 = plot(nu_inv, C_d_7, 'y');
598     %T_w = 50 C
599     p8 = plot(nu_inv, C_d_8, '—r');
600     p9 = plot(nu_inv, C_d_9, '—b');
601     p10 = plot(nu_inv, C_d_10, '—g');
602     p11 = plot(nu_inv, C_d_11, '—m');
603     p12 = plot(nu_inv, C_d_12, '—k');
604     p13 = plot(nu_inv, C_d_13, '—c');
605     p14 = plot(nu_inv, C_d_14, '—y');
606     %T_w = 100 C
607     p15 = plot(nu_inv, C_d_15, ':r');
608     p16 = plot(nu_inv, C_d_16, ':b');
609     p17 = plot(nu_inv, C_d_17, ':g');
610     p18 = plot(nu_inv, C_d_18, ':m');
611     p19 = plot(nu_inv, C_d_19, ':k');
612     p20 = plot(nu_inv, C_d_20, ':c');
613     p21 = plot(nu_inv, C_d_21, ':y');
614     hold off
615     legend([p1 p2 p3 p4 p5 p6 p7],{'\theta = 0 ', '\theta = 15 ', '\theta = 30 ', '\theta = 45 ', '\theta = 60 ', '\theta = 75 ', '\theta = 90 '});
616     xlabel('1/\nu');
617     ylabel('C_d');
618     title('T_w = 0 C / 50 C / 100 C & \alpha = 1');
619     axis tight
620     saveas(gcf, 'Model_9_2.png')
621 end
622 end

```

### C.2.1.2. Function

```

1 function [C_d_Schamberg_flat, V_rel, phi] = Schamberg_plots_flat(
    T_w, alpha, angle_of_attack, phi, nu_inv, model)
2
3 M = xlsread("Datos1.xlsx"); % Read Excel
4 Elements = [M(:,2), M(:,3), M(:,4), M(:,8), M(:,9), M(:,10),
    M(:,11), M(:,12)]; % Atmospheric composition matrix
5 Height = M(:,1); % Height

```

```

6     Temp_neutral = M(:,6); % Neutral Temperature
7     Rt = 6.37E6; % Earth radius
8     Mt = 5.972E24; % Earth mass
9     G = 6.674E-11; % Gravitational constant
10    mu = []; % Empty mu vector
11    k = 1.3806488E-23; % Boltzmann constant
12    RIG = 8.314472; % Ideal gases constant
13
14    % Molar masses in g/mol
15    mass_O = 15.999; mass_N2 = 28.01340; mass_O2 = 31.99880;
16    mass_He = 4.0026020;
17    mass_Ar = 39.9480; mass_H = 1.007940; mass_N = 14.00670;
18    mass_Anomo = 15.999;
19
20    % Masses vector in Kg/mol
21    mass_total = [mass_O, mass_N2, mass_O2, mass_He, mass_Ar,
22    mass_H, mass_N, mass_Anomo]/1000;
23
24    % Mu computation
25    for i=1:1:101
26        mu(i)=(Elements(i,:) * mass_total' / sum(Elements(i,:)));
27    end
28
29    R=RIG*ones(1, length(mu)) ./ (mu);
30
31    V_a = sqrt(2*R' .* Temp_neutral);
32    V_orb = sqrt((G*Mt*ones(101,1)) ./ (Height.*1000+Rt));
33
34    if (model == 1 || model == 4)
35        f_shape_flat = sin(angle_of_attack) .* sqrt(1-cos(
36            angle_of_attack)^(2/nu_inv)) - cos(angle_of_attack)
37            .^(1+1/nu_inv);
38        V_rel = sqrt(1+alpha.*(T_w/Temp_neutral(21) -1));
39        for i=1:1:length(V_rel)
40            C_d_Schamberg_flat(:, i) = 2*(1+f_shape_flat.*phi*
41                V_rel(i));
42        end
43    elseif (model == 2 || model == 5)
44        f_shape_flat = sin(angle_of_attack) .* sqrt(1-cos(
45            angle_of_attack)^(2/nu_inv)) - cos(angle_of_attack)
46            .^(1+1/nu_inv);
47        V_rel = sqrt(1+alpha.*(T_w./Temp_neutral(21) -1));
48        for i=1:1:length(V_rel)
49            C_d_Schamberg_flat(:, i) = 2*(1+f_shape_flat.*phi*
50                V_rel(i));
51        end
52    elseif (model == 3 || model == 6)
53        f_shape_flat = sin(angle_of_attack) .* sqrt(1-cos(

```

```

        angle_of_attack)^(2/nu_inv)) - cos(angle_of_attack)
        ^(1+1/nu_inv);
45     for i=1:1:length(T_w)
46         V_rel(:,i) = sqrt(1+alpha.*(T_w(i)/Temp_neutral(21) -1));
47     end
48     C_d_Schamberg_flat = 2*(1+V_rel.*phi*f_shape_flat);
49     else
50         phi_var = ((1-(2*phi./pi).^2)./(1-4*(2*phi./pi).^2))
            .*((0.5*sin(2*phi)-(2*phi./pi))./(sin(phi)-(2*phi./pi)
            ));
51         phi_var(1) = 1;
52         f_shape_flat = sin(angle_of_attack)*sqrt(1-cos(
            angle_of_attack).^(2*ones(1,length(nu_inv))./nu_inv))
            - cos(angle_of_attack).^(1+ones(1,length(nu_inv))./
            nu_inv);
53         V_rel = sqrt(1+alpha*(T_w/Temp_neutral(21) -1));
54         C_d_Schamberg_flat = 2*(1+V_rel*phi_var.*f_shape_flat);
55         phi=phi_var;
56     end
57 end

```

## C.2.2. Sphere

### C.2.2.1. Main menu

```

1  clear all
2  close all
3
4  list = {'Specular', 'Diffuse', 'Variable', 'Customize'};
5  [indx,tf] = listdlg('PromptString','Select an option:', '
    SelectionMode','single','ListSize',[400,200], 'ListString',list
    );
6
7  %% SPECULAR
8
9  if(indx == 1)
10     list = {'Constant Temperature', 'Constant accomodation factor'
        };
11     [indx,tf] = listdlg('PromptString','Select an option:', '
        SelectionMode','single','ListSize',[400,200], 'ListString',
        list);
12     if(indx == 1)
13         T_w1 = 0+273;
14         T_w2 = 50+273;
15         T_w3 = 100+273;
16         alpha = 0:0.01:1;
17         model = 1;
18         [C_d_1, V_rel1, phi] = Schamberg_plots_sphere(T_w1, 0,

```

```

        alpha, 1, model);
19 [C_d_2, V_rel2, phi] = Schamberg_plots_sphere(T_w2, 0,
        alpha, 1, model);
20 [C_d_3, V_rel3, phi] = Schamberg_plots_sphere(T_w3, 0,
        alpha, 1, model);
21 figure()
22 plot(alpha, C_d_1, 'r');
23 hold on
24 plot(alpha, C_d_2, 'k');
25 plot(alpha, C_d_3, 'b');
26 hold off
27 legend('T_w = 0 ', 'T_w = 50 ', 'T_w = 100 ')
28 xlabel('\alpha');
29 ylabel('C_d');
30 title('Specular')
31 saveas(gcf, 'Model_1.png')
32 end
33 if(indx == 2)
34     alpha_1 = 0;
35     alpha_2 = 0.5;
36     alpha_3 = 1;
37     T_w = 0:1:100;
38     model = 2;
39     [C_d_1, V_rel1, phi] = Schamberg_plots_sphere(T_w, 0,
        alpha_1, 1, model);
40     [C_d_2, V_rel2, phi] = Schamberg_plots_sphere(T_w, 0,
        alpha_2, 1, model);
41     [C_d_3, V_rel3, phi] = Schamberg_plots_sphere(T_w, 0,
        alpha_3, 1, model);
42     figure()
43     plot(T_w, C_d_1, 'r');
44     hold on
45     plot(T_w, C_d_2, 'k');
46     plot(T_w, C_d_3, 'b');
47     hold off
48     legend('\alpha = 0 ', '\alpha = 0.5 ', '\alpha = 1 ');
49     xlabel('T_w');
50     ylabel('C_d');
51     title('Specular');
52     saveas(gcf, 'Model_2.png')
53 end
54 end
55
56 %% DIFFUSE
57
58 if(indx == 2)
59     list = {'Constant Temperature', 'Constant accomodation factor'
        };

```

```

60 [indx,tf] = listdlg('PromptString','Select an option:', '
    SelectionMode','single','ListSize',[400,200], 'ListString',
    list);
61 if(indx == 1)
62     T_w1 = 0+273;
63     T_w2 = 50+273;
64     T_w3 = 100+273;
65     alpha = 0:0.01:1;
66     model = 1;
67     [C_d_1, V_rel1, phi] = Schamberg_plots_sphere(T_w1, pi/2,
        alpha, 0, model);
68     [C_d_2, V_rel2, phi] = Schamberg_plots_sphere(T_w2, pi/2,
        alpha, 0, model);
69     [C_d_3, V_rel3, phi] = Schamberg_plots_sphere(T_w3, pi/2,
        alpha, 0, model);
70     figure()
71     plot(alpha,C_d_1,'r');
72     hold on
73     plot(alpha,C_d_2,'k');
74     plot(alpha,C_d_3,'b');
75     hold off
76     legend('T_w = 0 ', 'T_w = 50 ', 'T_w = 100 ')
77     xlabel('\alpha');
78     ylabel('C_d');
79     title('Diffuse');
80     axis tight
81     saveas(gcf,'Model_3.png')
82 end
83 if(indx == 2)
84     alpha_1 = 0;
85     alpha_2 = 0.5;
86     alpha_3 = 1;
87     T_w = (0:1:100) + 273;
88     model = 2;
89     [C_d_1, V_rel1, phi] = Schamberg_plots_sphere(T_w, pi/2,
        alpha_1, 0, model);
90     [C_d_2, V_rel2, phi] = Schamberg_plots_sphere(T_w, pi/2,
        alpha_2, 0, model);
91     [C_d_3, V_rel3, phi] = Schamberg_plots_sphere(T_w, pi/2,
        alpha_3, 0, model);
92     figure()
93     plot(T_w-273,C_d_1,'r');
94     hold on
95     plot(T_w-273,C_d_2,'k');
96     plot(T_w-273,C_d_3,'b');
97     hold off
98     legend('\alpha = 0', '\alpha = 0.5', '\alpha = 1', '
        Location','northwest');

```

```

99         xlabel('T_w [ C ]');
100        ylabel('C_d');
101        title('Diffuse');
102        axis tight
103        saveas(gcf, 'Model_4.png')
104    end
105 end
106 %% VARIABLE
107
108 if(indx == 3)
109     list = {'alpha = 0 / 0.5 / 1', 'T_w = 0 / 50 C / 100 C'};
110     [indx, tf] = listdlg('PromptString', 'Select an option:', '
        SelectionMode', 'single', 'ListSize', [400, 200], 'ListString',
        list);
111     if(indx == 1)
112         T_w1 = 0+273;
113         T_w2 = 50+273;
114         T_w3 = 100+273;
115         alpha = 0;
116         alpha2 = 0.5;
117         alpha3 = 1;
118         phi_0 = 0:0.001:(pi/2);
119         nu_inv = 1:-(1/(length(phi_0)-1)):0;
120         model = 3;
121         [C_d_1, V_rel1, phi1] = Schamberg_plots_sphere(T_w1,
        phi_0, alpha, nu_inv, model);
122         [C_d_2, V_rel2, phi2] = Schamberg_plots_sphere(T_w2,
        phi_0, alpha, nu_inv, model);
123         [C_d_3, V_rel3, phi3] = Schamberg_plots_sphere(T_w3,
        phi_0, alpha, nu_inv, model);
124         [C_d_4, V_rel4, phi4] = Schamberg_plots_sphere(T_w1,
        phi_0, alpha2, nu_inv, model);
125         [C_d_5, V_rel5, phi5] = Schamberg_plots_sphere(T_w2,
        phi_0, alpha2, nu_inv, model);
126         [C_d_6, V_rel6, phi6] = Schamberg_plots_sphere(T_w3,
        phi_0, alpha2, nu_inv, model);
127         [C_d_7, V_rel7, phi7] = Schamberg_plots_sphere(T_w1,
        phi_0, alpha3, nu_inv, model);
128         [C_d_8, V_rel8, phi8] = Schamberg_plots_sphere(T_w2,
        phi_0, alpha3, nu_inv, model);
129         [C_d_9, V_rel9, phi9] = Schamberg_plots_sphere(T_w3,
        phi_0, alpha3, nu_inv, model);
130        figure()
131        p1 = plot(phi1, C_d_1, 'r');
132        hold on
133        p2 = plot(phi2, C_d_2, 'k');
134        p3 = plot(phi3, C_d_3, 'b');
135        p4 = plot(phi4, C_d_4, '—r');

```

```

136     p5 = plot(phi5 , C_d_5 , '—k');
137     p6 = plot(phi6 , C_d_6 , '—b');
138     p7 = plot(phi7 , C_d_7 , ':r');
139     p8 = plot(phi8 , C_d_8 , ':k');
140     p9 = plot(phi9 , C_d_9 , ':b');
141     hold off
142     legend([p1 p2 p3],{ 'T_w = 0 ', 'T_w = 50 ', 'T_w = 100 '
143         });
144     xlabel( '\phi(\phi_0) ');
145     ylabel( 'C_d' );
146     title( '\alpha = 0 / 0.5 / 1 ');
147     axis tight
148     saveas(gcf , 'Model_5_1.png')
149
150     figure ()
151     p1 = plot(nu_inv , C_d_1 , 'r');
152     hold on
153     p2 = plot(nu_inv , C_d_2 , 'k');
154     p3 = plot(nu_inv , C_d_3 , 'b');
155     p4 = plot(nu_inv , C_d_4 , '—r');
156     p5 = plot(nu_inv , C_d_5 , '—k');
157     p6 = plot(nu_inv , C_d_6 , '—b');
158     p7 = plot(nu_inv , C_d_7 , ':r');
159     p8 = plot(nu_inv , C_d_8 , ':k');
160     p9 = plot(nu_inv , C_d_9 , ':b');
161     hold off
162     legend([p1 p2 p3],{ 'T_w = 0 ', 'T_w = 50 ', 'T_w = 100 '
163         });
164     xlabel( '1/\nu' );
165     ylabel( 'C_d' );
166     title( '\alpha = 0 / 0.5 / 1 ');
167     axis tight
168     saveas(gcf , 'Model_5_2.png')
169
170 end
171
172 if(indx == 2)
173     alpha1 = 0;
174     alpha2 = 0.5;
175     alpha3 = 1;
176     T_w = 0+273;
177     T_w2 = 50+273;
178     T_w3 = 100+273;
179     phi_0 = 0:0.001:(pi/2);
180     nu_inv = 1:-(1/(length(phi_0)-1)):0;
181     model = 3;
182     [C_d_1 , V_rel1 , phi1] = Schamberg_plots_sphere(T_w , phi_0
183         , alpha1 , nu_inv , model);
184     [C_d_2 , V_rel2 , phi2] = Schamberg_plots_sphere(T_w , phi_0
185         , alpha2 , nu_inv , model);

```

```

180 [C_d_3, V_rel3, phi3] = Schamberg_plots_sphere(T_w, phi_0
      , alpha3, nu_inv, model);
181 [C_d_4, V_rel4, phi4] = Schamberg_plots_sphere(T_w2,
      phi_0, alpha1, nu_inv, model);
182 [C_d_5, V_rel5, phi5] = Schamberg_plots_sphere(T_w2,
      phi_0, alpha2, nu_inv, model);
183 [C_d_6, V_rel6, phi6] = Schamberg_plots_sphere(T_w2,
      phi_0, alpha3, nu_inv, model);
184 [C_d_7, V_rel7, phi7] = Schamberg_plots_sphere(T_w3,
      phi_0, alpha1, nu_inv, model);
185 [C_d_8, V_rel8, phi8] = Schamberg_plots_sphere(T_w3,
      phi_0, alpha2, nu_inv, model);
186 [C_d_9, V_rel9, phi9] = Schamberg_plots_sphere(T_w3,
      phi_0, alpha3, nu_inv, model);
187 figure()
188 p1 = plot(phi1, C_d_1, 'r');
189 hold on
190 p2 = plot(phi2, C_d_2, 'k');
191 p3 = plot(phi3, C_d_3, 'b');
192 p4 = plot(phi4, C_d_4, '—r');
193 p5 = plot(phi5, C_d_5, '—k');
194 p6 = plot(phi6, C_d_6, '—b');
195 p7 = plot(phi7, C_d_7, ':r');
196 p8 = plot(phi8, C_d_8, ':k');
197 p9 = plot(phi9, C_d_9, ':b');
198 hold off
199 legend([p1 p2 p3],{'\alpha = 0', '\alpha = 0.5', '\alpha
      = 1'});
200 xlabel('\phi(\phi_0)');
201 ylabel('C_d');
202 title('T_w = 0 C / 50 C / 100 C');
203 axis tight
204 saveas(gcf, 'Model_6_1.png')
205
206 figure()
207 p1 = plot(nu_inv, C_d_1, 'r');
208 hold on
209 p2 = plot(nu_inv, C_d_2, 'k');
210 p3 = plot(nu_inv, C_d_3, 'b');
211 p4 = plot(nu_inv, C_d_4, '—r');
212 p5 = plot(nu_inv, C_d_5, '—k');
213 p6 = plot(nu_inv, C_d_6, '—b');
214 p7 = plot(nu_inv, C_d_7, ':r');
215 p8 = plot(nu_inv, C_d_8, ':k');
216 p9 = plot(nu_inv, C_d_9, ':b');
217 hold off
218 legend([p1 p2 p3],{'\alpha = 0', '\alpha = 0.5', '\alpha
      = 1'});

```



```

219         xlabel('1/\nu');
220         ylabel('C_d');
221         title('T_w = 0 C / 50 C / 100 C');
222         axis tight
223         saveas(gcf, 'Model_6_2.png')
224     end
225 end

```

#### C.2.2.2. Function

```

1 function [C_d_Schamberg, V_rel, phi] = Schamberg_plots_sphere(T_w,
    phi_0, alpha, nu_inv, model)
2     M = xlsread("Datos1.xlsx"); % Read Excel
3     Elements = [M(:,2), M(:,3), M(:,4), M(:,8), M(:,9), M(:,10),
        M(:,11), M(:,12)]; % Atmospheric composition matrix
4     Height = M(:,1); % Height
5     Temp_neutral = M(:,6); % Neutral Temperature
6     Rt = 6.37E6; % Earth radius
7     Mt = 5.972E24; % Earth mass
8     G = 6.674E-11; % Gravitational constant
9     mu = []; % Empty mu vector
10    k = 1.3806488E-23; % Boltzmann constant
11    RIG = 8.314472; % Ideal gases constant
12
13    % Molar masses in g/mol
14    mass_O = 15.999; mass_N2 = 28.01340; mass_O2 = 31.99880;
        mass_He = 4.0026020;
15    mass_Ar = 39.9480; mass_H = 1.007940; mass_N = 14.00670;
        mass_Anomo = 15.999;
16
17    % Masses vector in Kg/mol
18    mass_total = [mass_O, mass_N2, mass_O2, mass_He, mass_Ar,
        mass_H, mass_N, mass_Anomo]/1000;
19
20    % Mu computation
21    for i=1:1:101
22        mu(i)=(Elements(i,:) * mass_total' / sum(Elements(i,:)));
23    end
24
25    R=RIG*ones(1, length(mu)) ./ (mu);
26
27    V_a = sqrt(2*R' .* Temp_neutral);
28    V_orb = sqrt((G*Mt*ones(101,1)) ./ (Height.*1000+Rt));
29
30    if (model==1)
31        phi = ((1-(2*phi_0/pi)^2)/(1-4*(2*phi_0/pi)^2))*((0.5*sin
            (2*phi_0)-(2*phi_0/pi))/(sin(phi_0)-(2*phi_0/pi)));
32        if (phi_0==pi/2)
33            phi=2/3;

```

```

34     elseif(phi_0 == 0)
35         phi = 1;
36     end
37     fun = @(x,nu) x.*sqrt((1-x.^(2/nu)).*(1-x.^2));
38     l_1 = integral(@(x)fun(x,nu_inv),0,1);
39     f_shape = 2*(l_1 - (1/(1/nu_inv + 3)));
40     V_rel = sqrt(1+alpha.*(T_w/Temp_neutral(21) -1));
41     C_d_Schamberg = 2*(1+phi*V_rel.*f_shape);
42 elseif(model==2)
43     phi = ((1-(2*phi_0/pi)^2)/(1-4*(2*phi_0/pi)^2))*((0.5*sin
44         (2*phi_0)-(2*phi_0/pi))/(sin(phi_0)-(2*phi_0/pi)));
45     if(phi_0==pi/2)
46         phi=2/3;
47     elseif(phi_0 == 0)
48         phi = 1;
49     end
50     fun = @(x,nu) x.*sqrt((1-x.^(2/nu)).*(1-x.^2));
51     l_1 = integral(@(x)fun(x,nu_inv),0,1);
52     f_shape = 2*(l_1 - (1/(1/nu_inv + 3)));
53     V_rel = sqrt(1+alpha.*(T_w./Temp_neutral(21) -1));
54     C_d_Schamberg = 2*(1+phi*V_rel.*f_shape);
55 elseif(model==3)
56     phi = ((1-(2*phi_0./pi).^2)/(1-4*(2*phi_0./pi).^2))
57         .*((0.5*sin(2*phi_0)-(2*phi_0./pi))./(sin(phi_0)-(2*
58         phi_0./pi)));
59     phi(1) = 1;
60     fun = @(x,nu) x.*sqrt((1-x.^(2/nu)).*(1-x.^2));
61     for i=1:1:length(nu_inv)
62         l_1(i) = integral(@(x)fun(x,nu_inv(i)),0,1);
63     end
64     f_shape = 2*(l_1 - (1*ones(1,length(l_1)))./((1*ones(1,
65         length(l_1))./nu_inv)+3));
66     V_rel = sqrt(1+alpha.*(T_w/Temp_neutral(21) -1));
67     C_d_Schamberg = 2*(1+V_rel*phi.*f_shape);
68 elseif(model==4)
69     phi = phi_0;
70     fun = @(x,nu) x.*sqrt((1-x.^(2/nu)).*(1-x.^2));
71     l_1 = integral(@(x)fun(x,nu_inv),0,1);
72     f_shape = 2*(l_1 - (1/(1/nu_inv + 3)));
73     V_rel = sqrt(1+alpha.*(T_w./Temp_neutral(21) -1));
74     C_d_Schamberg = 2*(1+phi*V_rel.*f_shape);
75 end
76 end

```

### C.3. Sentman model

```

1 clear all
2 close all

```

```

3
4 M = xlsread("Datos1.xlsx"); % Read Excel
5 Elements = [M(:,2), M(:,3), M(:,4), M(:,8), M(:,9), M(:,10), M
    (:,11), M(:,12)]; % Atmospheric composition matrix
6 Height = M(:,1); % Height
7 Temp_neutral = M(:,6); % Neutral Temperature
8 Rt = 6.37E6; % Earth radius
9 Mt = 5.972E24; % Earth mass
10 G = 6.674E-11; % Gravitational constant
11 mu = []; % Empty mu vector
12 k = 1.3806488E-23; % Boltzmann constant
13 RIG = 8.314472; % Ideal gases constant
14
15 % Molar masses in g/mol
16 mass_O = 15.999; mass_N2 = 28.01340; mass_O2 = 31.99880; mass_He
    = 4.0026020;
17 mass_Ar = 39.9480; mass_H = 1.007940; mass_N = 14.00670;
    mass_Anomo = 15.999;
18
19 % Masses vector in Kg/mol
20 mass_total = [mass_O, mass_N2, mass_O2, mass_He, mass_Ar, mass_H,
    mass_N, mass_Anomo]/1000;
21
22 % Mu computation
23 for i=1:1:101
24 mu(i)=(Elements(i,:) * mass_total' / sum(Elements(i,:)));
25 end
26
27 R=RIG*ones(1,length(mu)) ./ (mu);
28 V_a = sqrt(2*R' .* Temp_neutral);
29 V_orb = sqrt((G*Mt*ones(101,1)) ./ (Height.*1000+Rt));
30 s = (V_orb(21)/V_a(21));
31
32 angle_of_attack = 0:0.00001:(pi/2);
33 alpha = [0 0.25 0.5 0.75 1];
34 for i=1:length(alpha)
35     C_d_flat(:,i) = (2/(s*sqrt(pi)))*exp(-s^2*(sin(
        angle_of_attack).^2)) + (sin(angle_of_attack)./(s^2))
        .*(1+2*s^2).*erf(s*sin(angle_of_attack)) + (sqrt(pi)/s)*(
        sin(angle_of_attack).^2).*sqrt(1-alpha(i));
36 end
37 alpha=0:0.01:1;
38 C_d_sph = ((2*s^2+1)/(sqrt(pi)*s^3))*exp(-s^2) + ((4*s^4+4*s^2-1)
    ./(2*s^4))*erf(s) + ((2*sqrt(pi))/(3*s))*sqrt(1-alpha);
39
40 figure()
41 plot(angle_of_attack*180/pi, C_d_flat(:,1), 'r')
42 hold on

```

```

43 plot( angle_of_attack*180/pi , C_d_flat(:,2) , 'b')
44 plot( angle_of_attack*180/pi , C_d_flat(:,3) , 'g')
45 plot( angle_of_attack*180/pi , C_d_flat(:,4) , 'm')
46 plot( angle_of_attack*180/pi , C_d_flat(:,5) , 'k')
47 xlabel( 'Angle of Attack [rad]' );
48 ylabel( 'C_D' );
49 legend( '\alpha = 0' , '\alpha = 0.25' , '\alpha = 0.5' , '\alpha = 0.75'
        , '\alpha = 1' , 'Location' , 'northwest' );
50 axis tight
51
52 figure()
53 plot( alpha , C_d_sph)
54 xlabel( '\alpha' );
55 ylabel( 'C_D' );
56 axis tight

```